

**Cien Años de Gravedad**

.

**One Hundred Years  
of Nigritude**

25 Nov 1915

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = 8\pi G T_{\mu\nu}$$

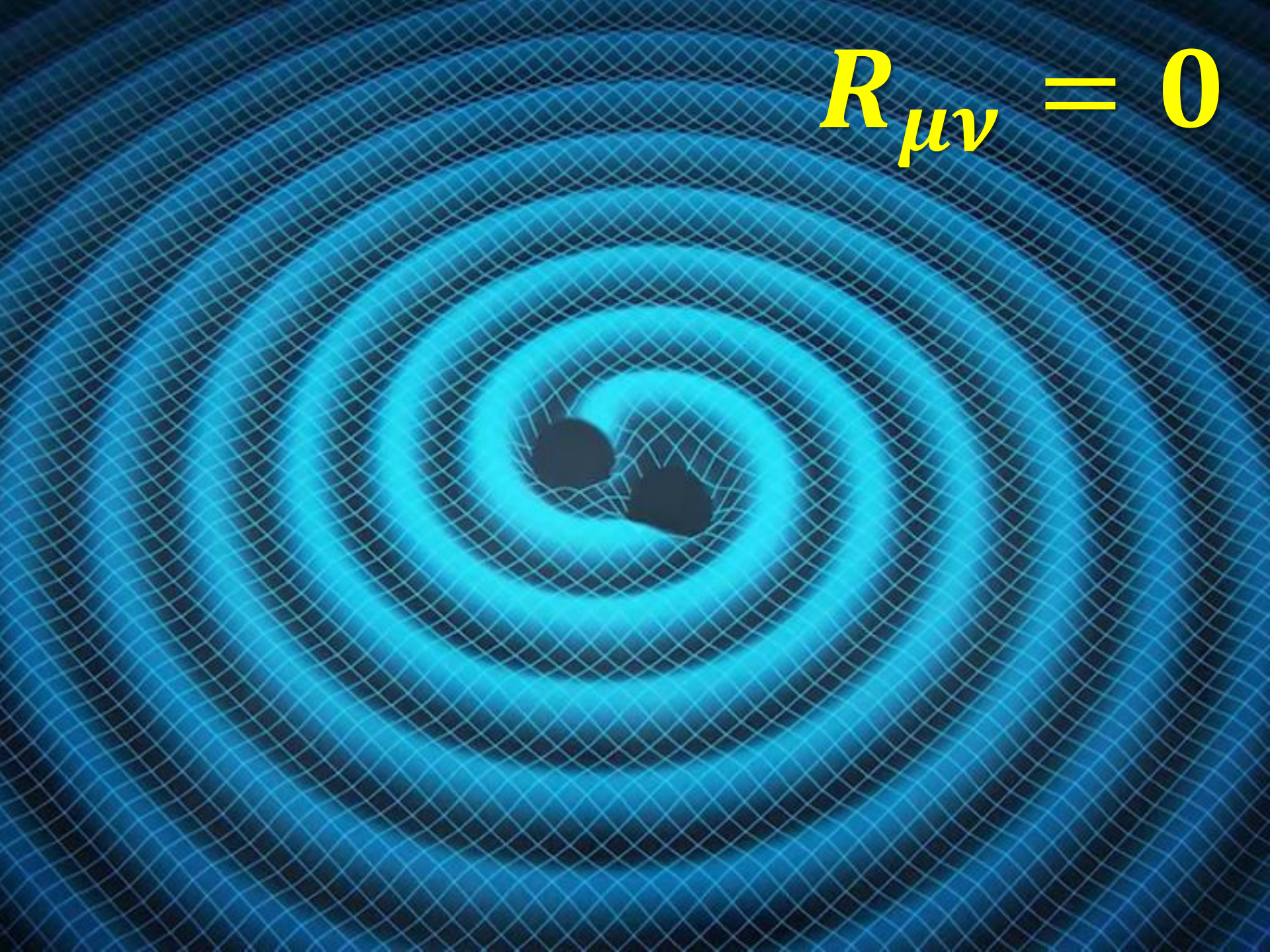
## Die Feldgleichungen der Gravitation.

VON A. EINSTEIN.

$$R_{im} = \sum_i \frac{\partial \Gamma_{im}^i}{\partial x_i} + \sum_{i,i'} \Gamma_{i'i}^i \Gamma_{im}^{i'} = -\kappa \left( T_{im} - \frac{1}{2} g_{im} T \right)$$

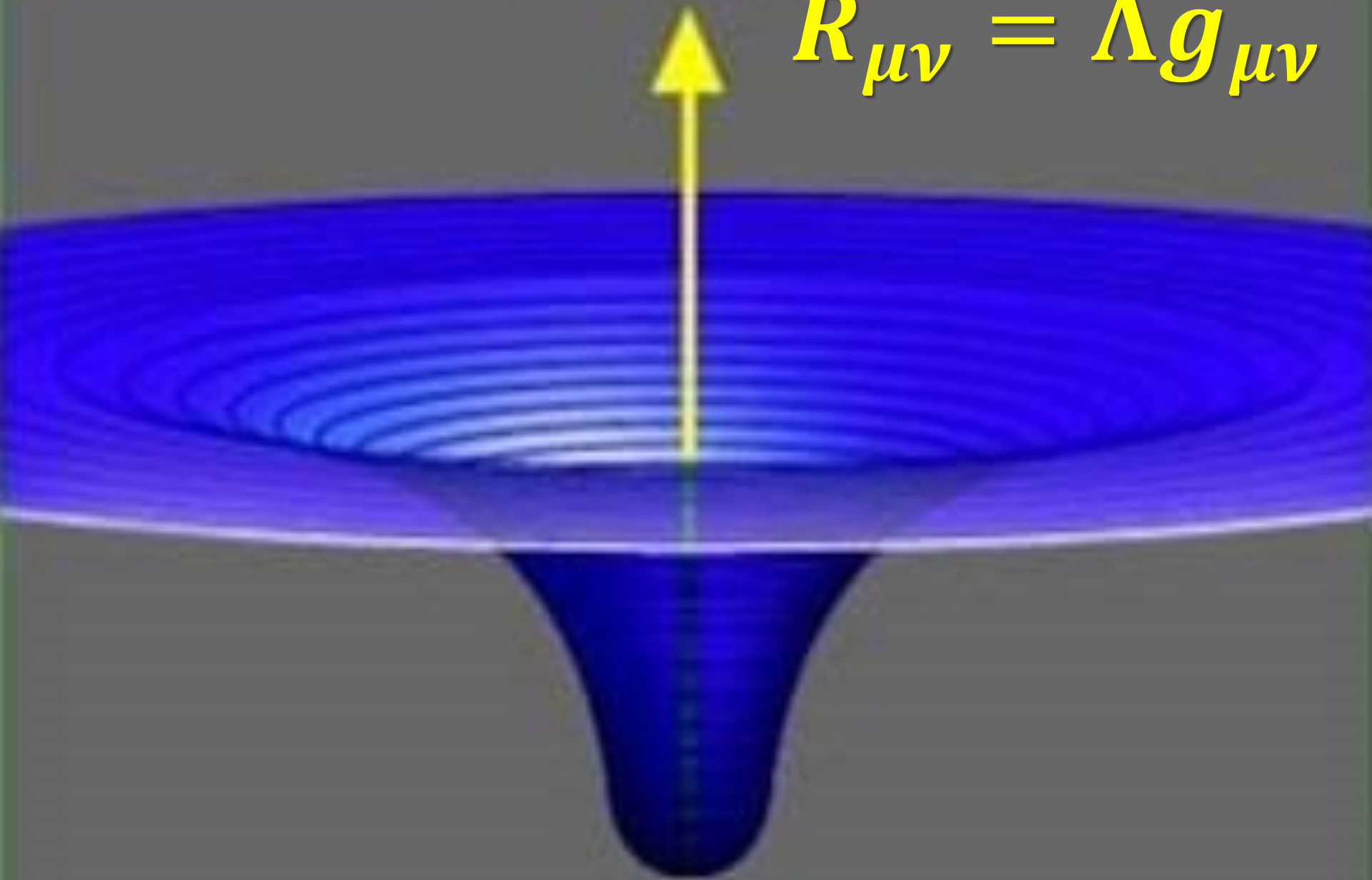


$$R_{\mu\nu} = 0$$



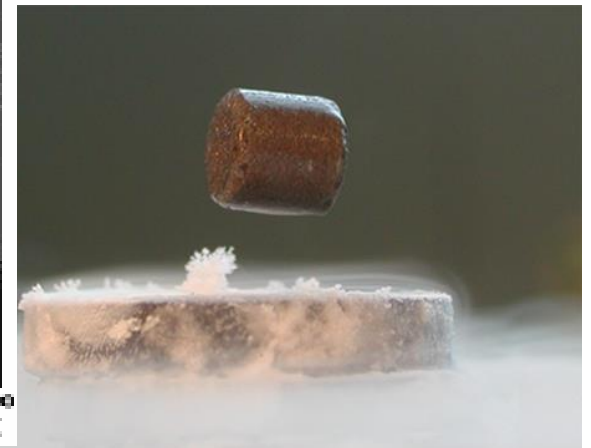
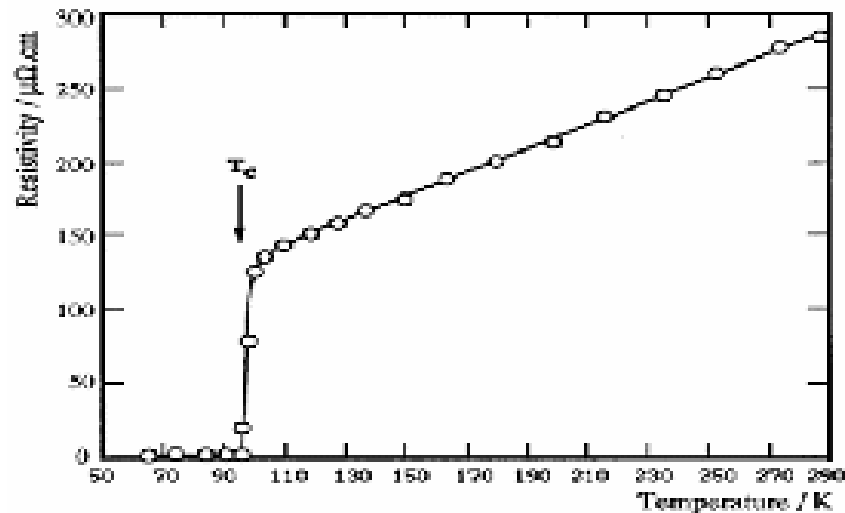


$$R_{\mu\nu} = \Lambda g_{\mu\nu}$$

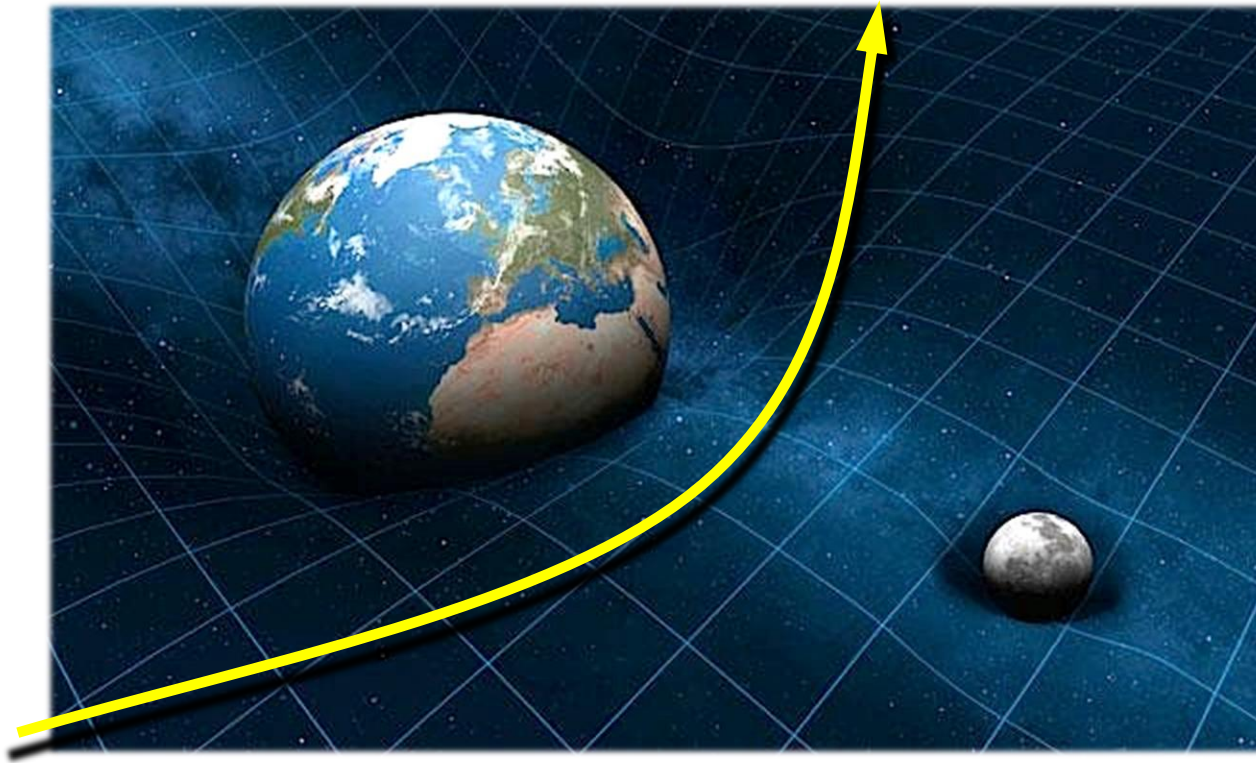


quark-gluon plasma

$$R_{\mu\nu} = -\Lambda g_{\mu\nu}$$



# Gravity as curvature



What's wrong with this picture?

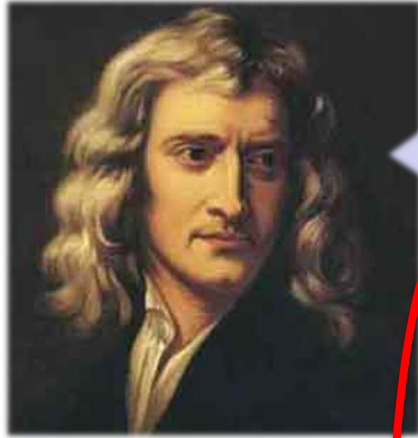
Particles follow **geodesics**  
in a curved space:

“most straight” (extremal)  
paths

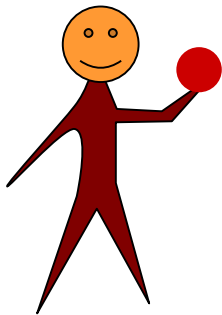
On **Earth**, the curvature is  
**small**



On a **weakly curved** space,  
geodesics are **very** close to  
straight lines



Are you telling *me* that  
this trajectory is  
**almost straight** and  
**weakly curved ???**





Massive particles follow trajectories in space *and time*

Since  $c = 300,000 \text{ km/s}$  is large,  
they move **much more in time**  
than in space

$$1 \text{ s} = 300,000,000 \text{ m}$$



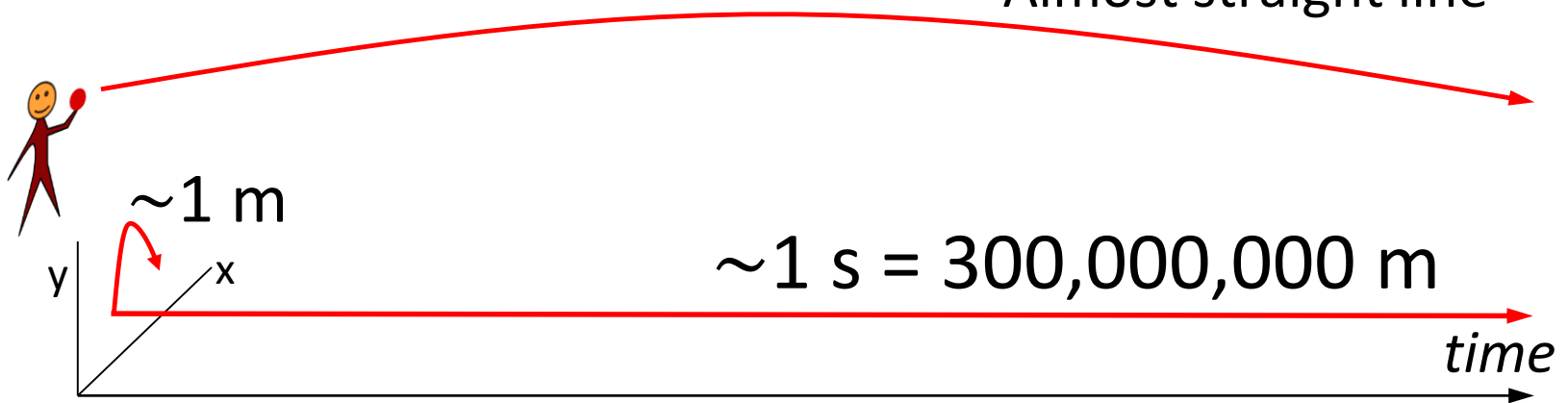


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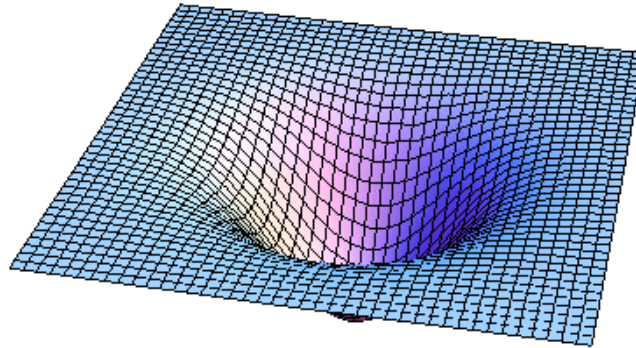
This trajectory is **weakly curved**

Almost straight line



# Gravity as curvature

When we represent a gravitational field as

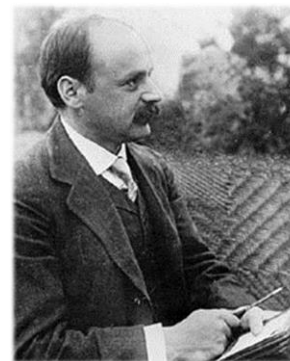


it is to be understood that there is also  
*curvature* in the *time* direction

# K Schwarzschild to A Einstein

letter dated 22 December 1915

from the Russian war front



*“I made at once by good luck a search for a full solution. A not too difficult calculation gave the following result:”*

$$ds^2 = - \left( 1 - \frac{2GM}{r} \right) dt^2 + \frac{dr^2}{1 - \frac{2GM}{r}} + r^2 (d\theta^2 + \sin^2 \theta d\phi^2)$$



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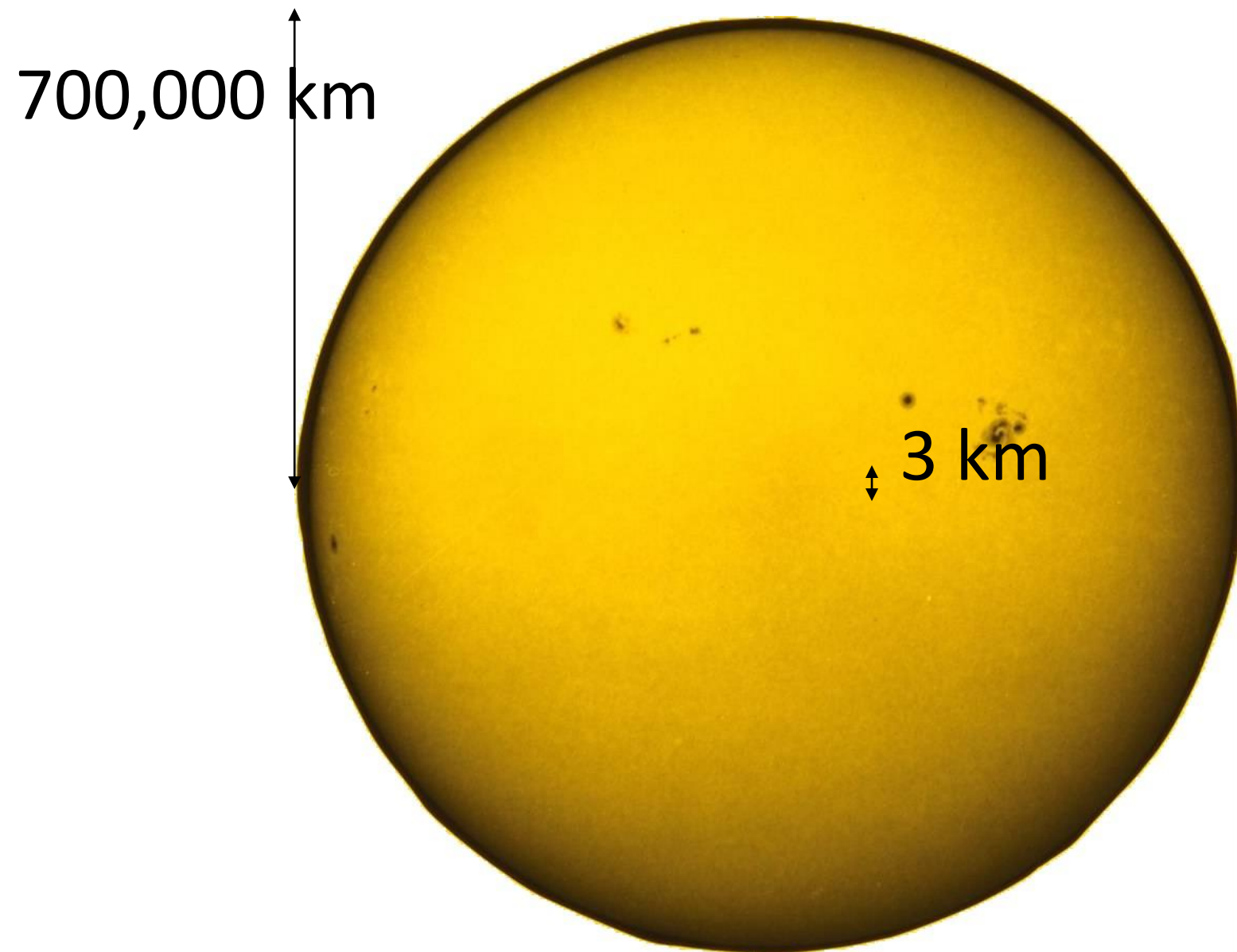
Something odd going on at  $r = r_s \equiv 2GM$

“Schwarzschild singularity”

So?

For a star:  $r_S \simeq 3 \text{ km}$





700,000 km

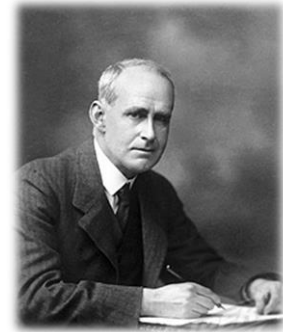
3 km

That's *way too small!*

Ignore it...

# BUT...

## Eddington (1924)



found new coordinates that are

not singular at  $r = r_S$

(he did not mention this)

## Lemaitre (1932-33)



$r = r_S$  just a coordinate singularity

So,  
is this physical or not?



Can gravitational collapse  
shrink a star beyond the  
“Schwarzschild singularity”?

# MEETING OF THE ROYAL ASTRONOMICAL SOCIETY.

Friday, 1935 January 11.

*The President* then closed the discussion on Nova Herculis, and asked **Dr. Chandrasekhar** to give an account of his recent investigation of Stellar Configurations.

If the star's mass is greater than  $M$  the star cannot have a degenerate core, but if the star's mass is less than  $M$  it will tend, at the end of its life history, towards a **com-pletely collapsed state.**



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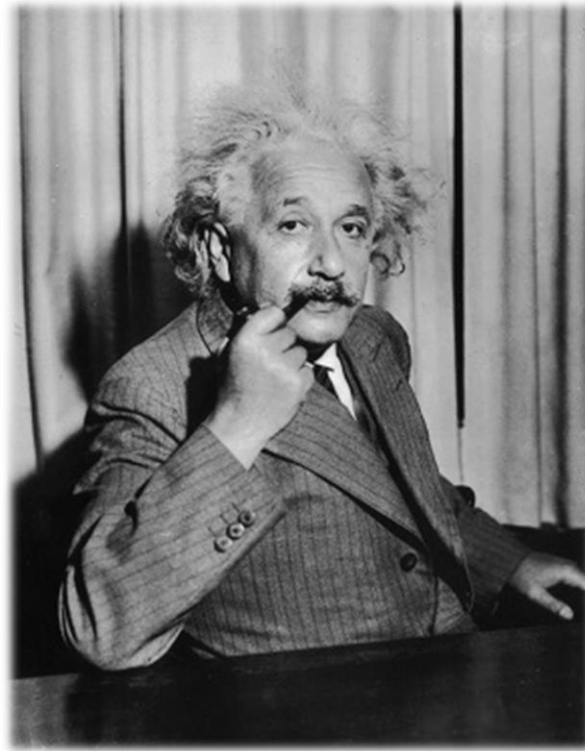
*Sir Arthur Eddington.*



The star has to go on radiating and radiating and contracting and contracting until, I suppose, it gets down to **a few km. radius**, when **gravity becomes strong enough to hold in the radiation**, and the star can at last find peace.

I felt driven to the conclusion that this was almost a **reductio ad absurdum** of the relativistic degeneracy formula. Various accidents may intervene to save the star, but I want more protection than that. **I think there should be a law of Nature to prevent a star from behaving in this absurd way !**

# Ask the master





# Einstein's *worst* blunder

He failed to recognize a *most striking* prediction of his theory:

Schwarzschild's solution describes a **Black Hole**

He even wrote a regrettable, confused paper **denying** the possibility

ON A STATIONARY SYSTEM WITH SPHERICAL SYMMETRY  
CONSISTING OF MANY GRAVITATING MASSES

BY ALBERT EINSTEIN

(Received May 10, 1939)

The essential result of this investigation is a clear understanding as to why the “Schwarzschild singularities” do not exist in physical reality. Although the theory given here treats only clusters whose particles move along circular paths it does not seem to be subject to reasonable doubt that more general cases will have analogous results. The “Schwarzschild singularity” does not appear for the reason that matter cannot be concentrated arbitrarily. And this is due to the fact that otherwise the constituting particles would reach the velocity of light.

# ON A STATIONARY SYSTEM WITH SPHERICAL SYMMETRY CONSISTING OF MANY GRAVITATING MASSES

BY ALBERT EINSTEIN

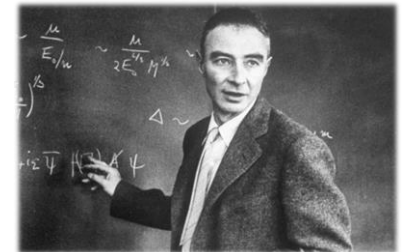
(Received May 10, 1939)

## On Continued Gravitational Contraction

J. R. OPPENHEIMER AND H. SNYDER

*University of California, Berkeley, California*

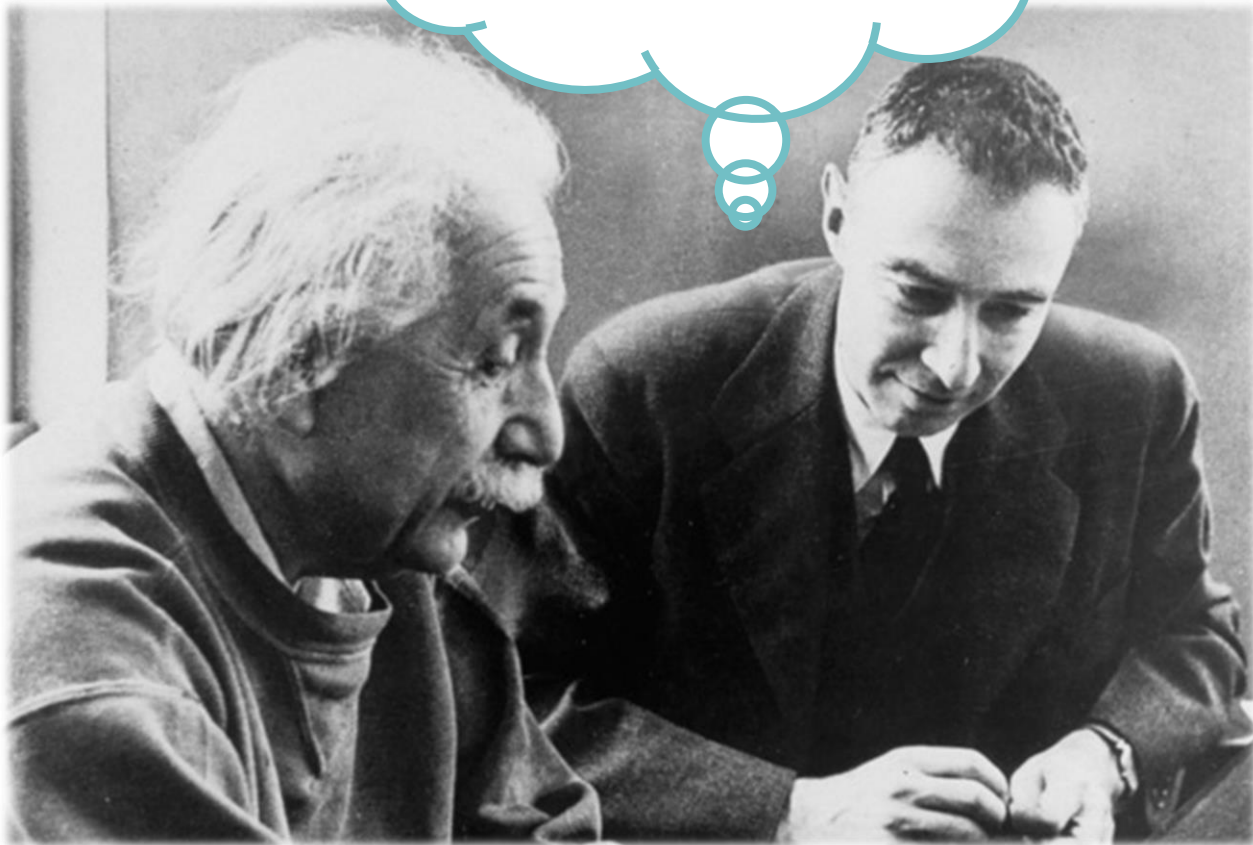
(Received July 10, 1939)



But Einstein never knew about his blunder

Years later,  
at IAS...

Let's not  
talk about  
it...



# 1950's-1960's

“Schwarzschild singularities”  
are physical,  
unavoidable consequences  
of General Relativity



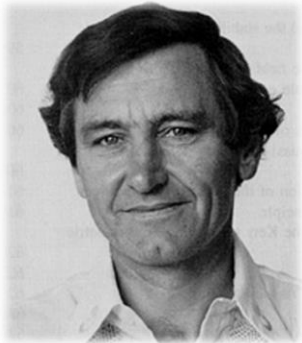
JA Wheeler



D Finkelstein



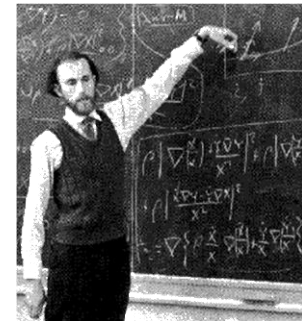
W Israel



R Kerr



R Penrose

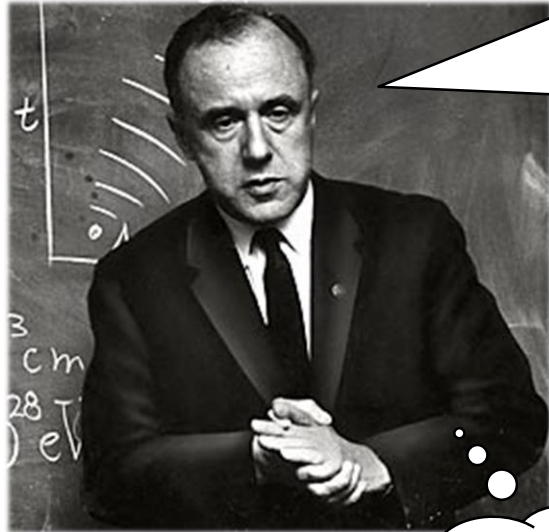


B Carter



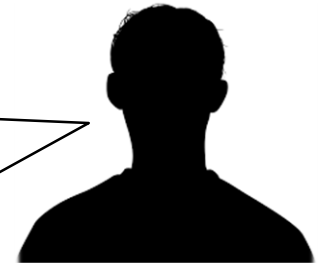
Since the baby has survived childbirth  
let's give him a proper name...

JA Wheeler 1967



One can't keep calling it  
“gravitationally completely  
collapsed object” ...

How about  
**BLACK HOLE?**

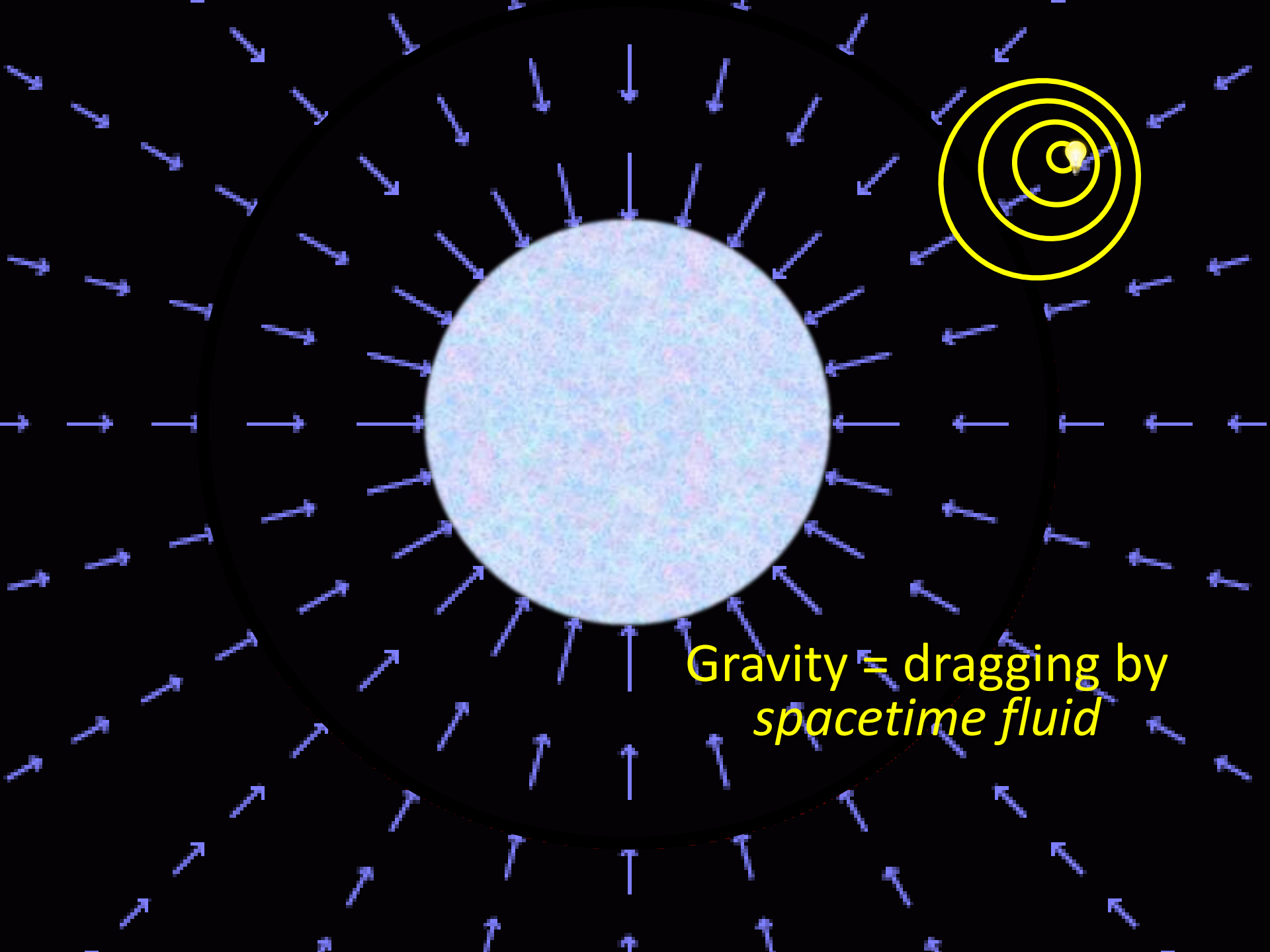


...yes!

That's an  
obscene  
name!



# Space and time *flow*



Gravity = dragging by  
*spacetime fluid*

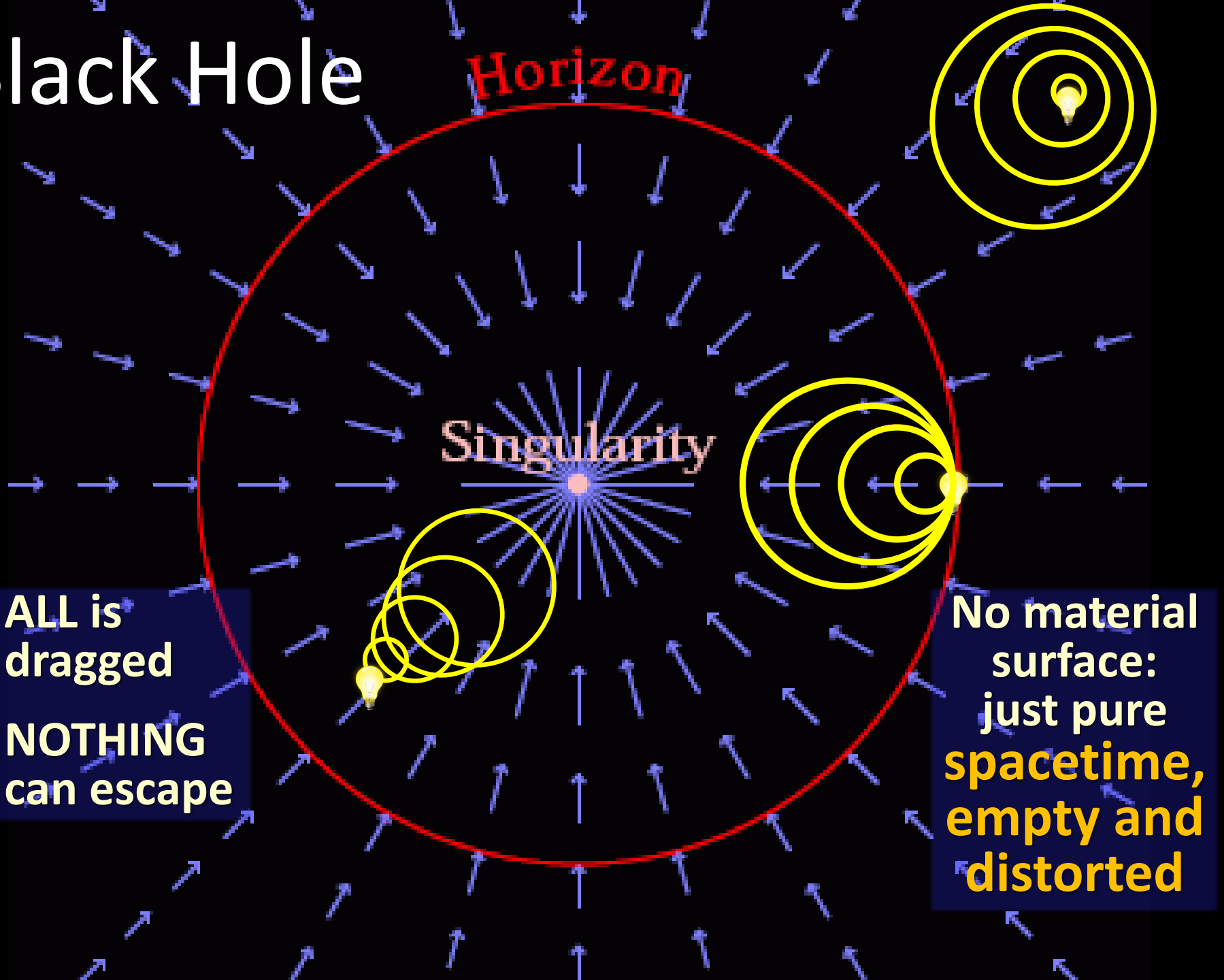
# Black Hole

Horizon

Singularity

ALL is  
dragged  
**NOTHING**  
can escape

No material  
surface:  
just pure  
**spacetime,**  
**empty and**  
**distorted**





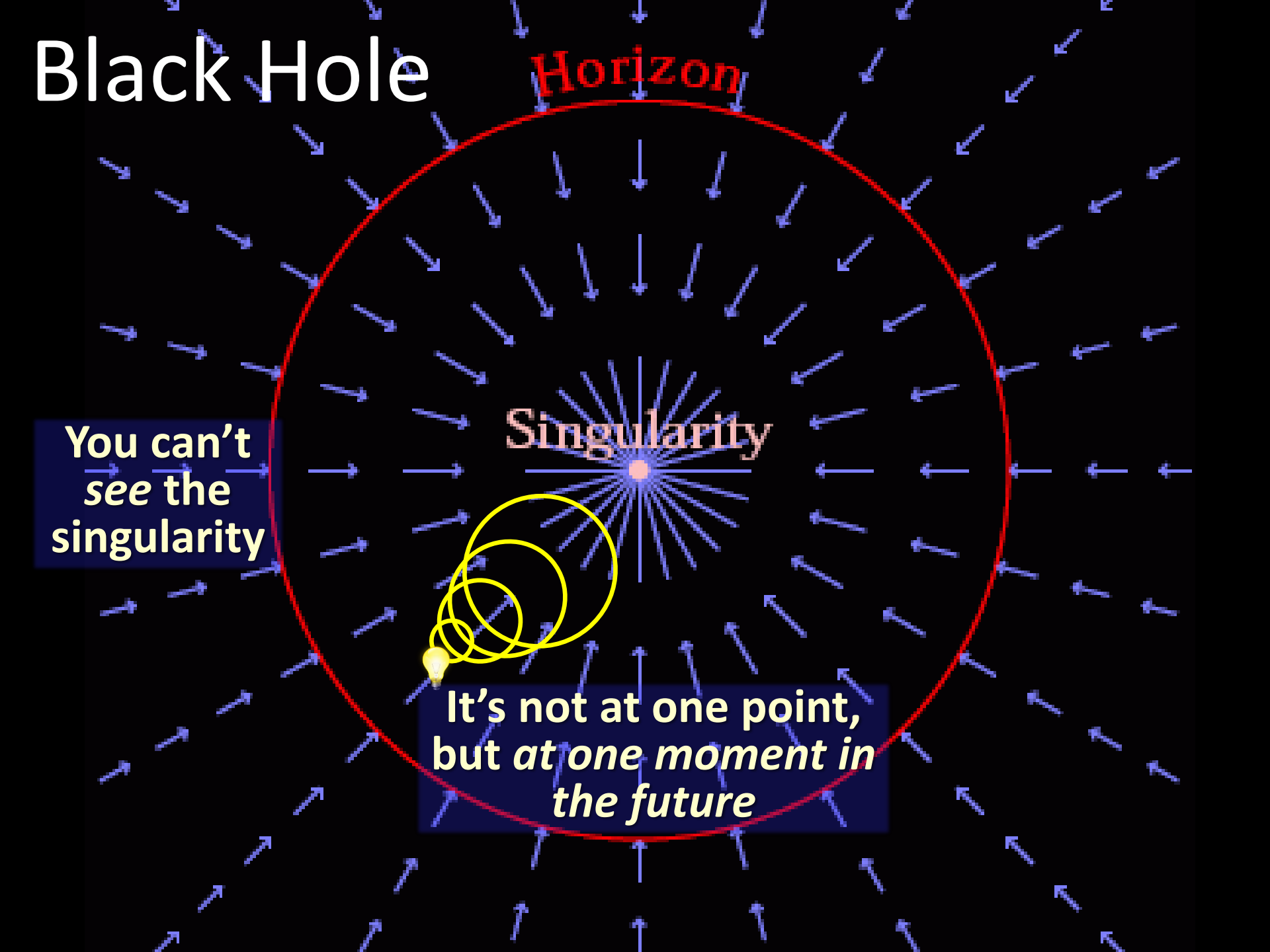
# Black Hole

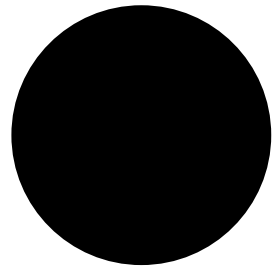
Horizon

Singularity

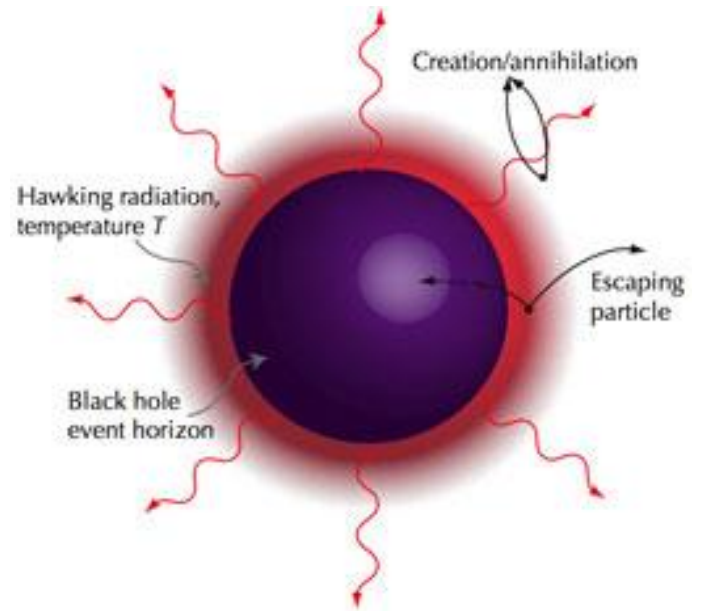
You can't  
see the  
singularity

It's not at one point,  
but *at one moment in  
the future*





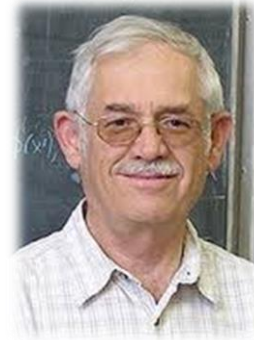
+  $\hbar$



1974

# First equation of Quantum Gravity

$$S_{BH} = \frac{Area}{4G\hbar}$$



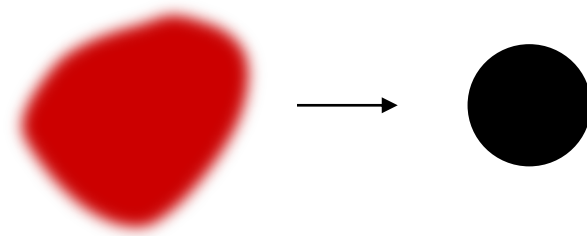
Bekenstein – Hawking

# Black Hole Information Problem

*Hawking 1976*

A problem of *fundamental irreversibility*

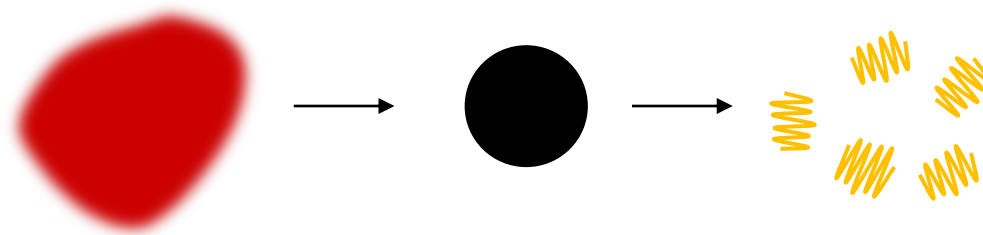
Black holes have no hair



Where does the information about the initial state go?

Maybe it's still inside the BH...

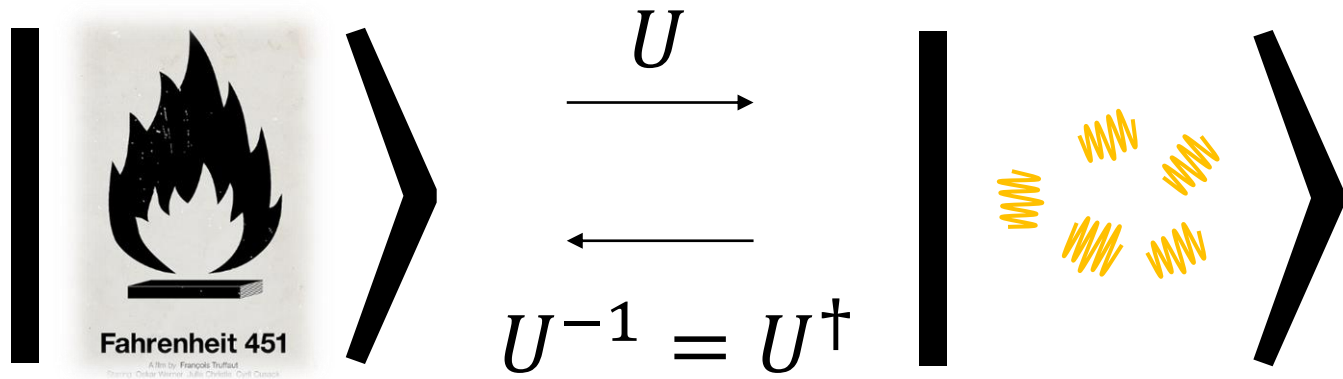
# What about this?



Thermal radiation

# Quantum evolution is reversible

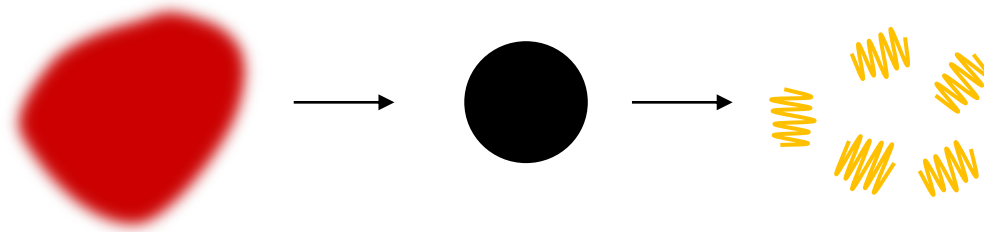
Pure state  $\rightleftharpoons$  Pure state





# Non-Unitary Quantum evolution?

Pure state  $\rightarrow$  Mixed state ?



no hair  $\rightarrow$  no info in radiation

# Possibility A

Information is **fundamentally lost**

Yes, Non-unitary evolution in Quantum Gravity

Most of Quantum Field Theory, incl.  
perturbative Quantum Gravity, remains as usual

In the presence of bhs, *small subtle* violations of  
unitarity

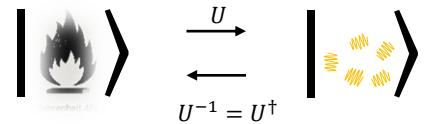
We just have to live with this

# Possibility B

Information is **not lost**

Quantum mechanics is extremely difficult to modify in a consistent manner

Info must be retrievable outside bh and after evaporation, much like

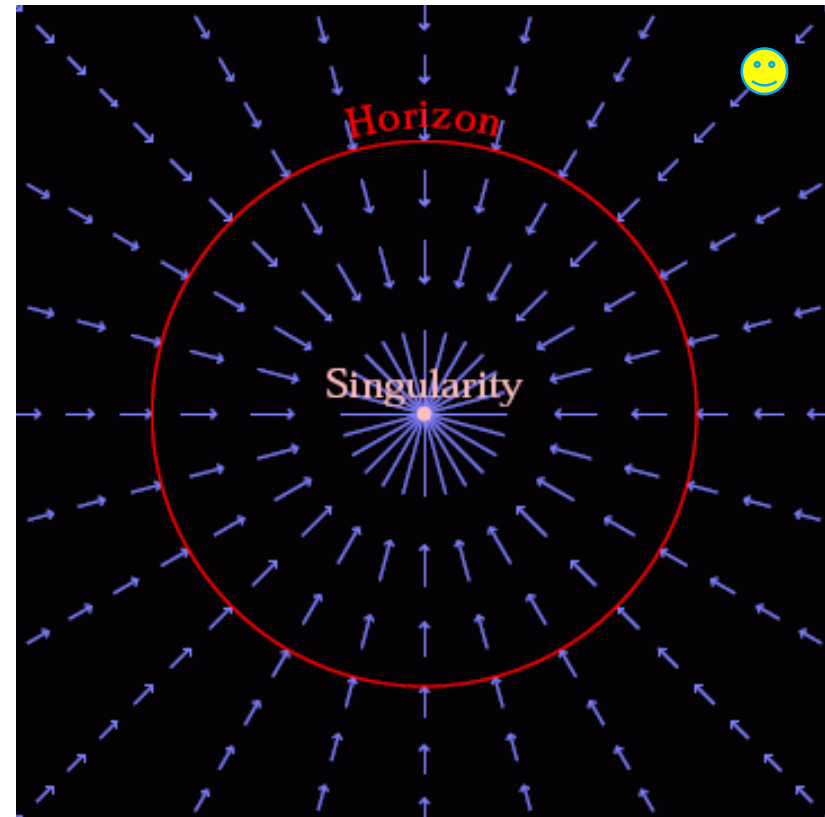


But how? Info that has fallen inside the bh can't be cloned outside

# Firewall argument – simplified version

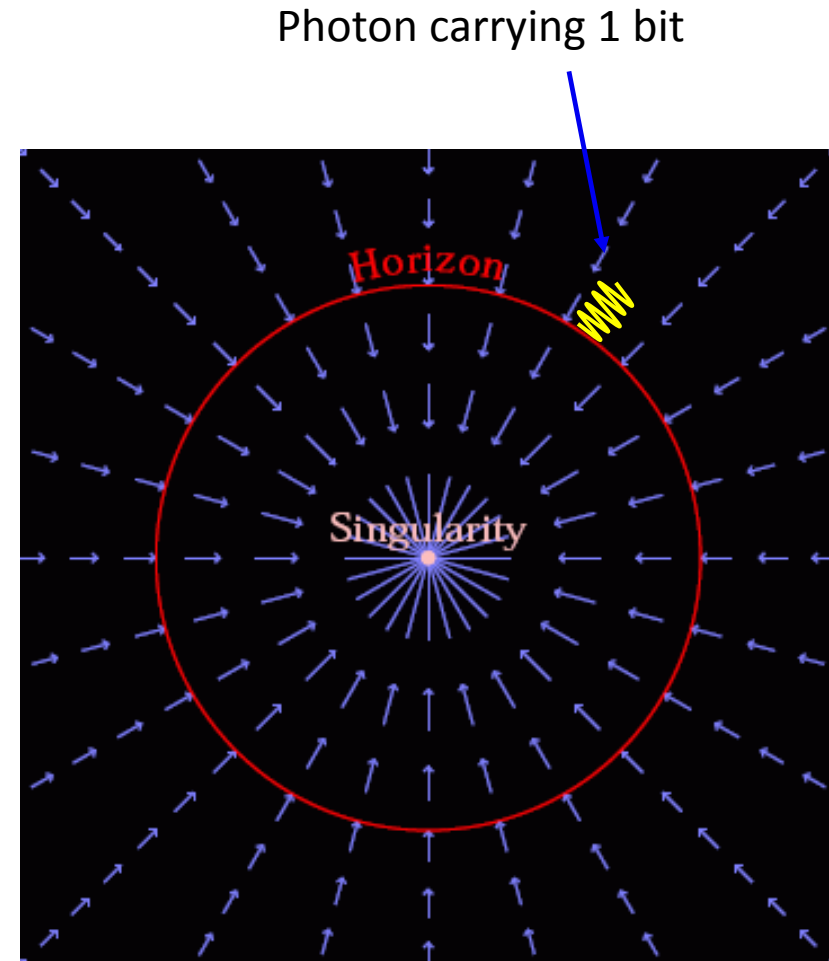
*Almheiri+Marolf+Polchinski+Sully*

1. An observer in free-fall would **not feel anything special** when crossing the horizon



# Firewall argument – simplified version

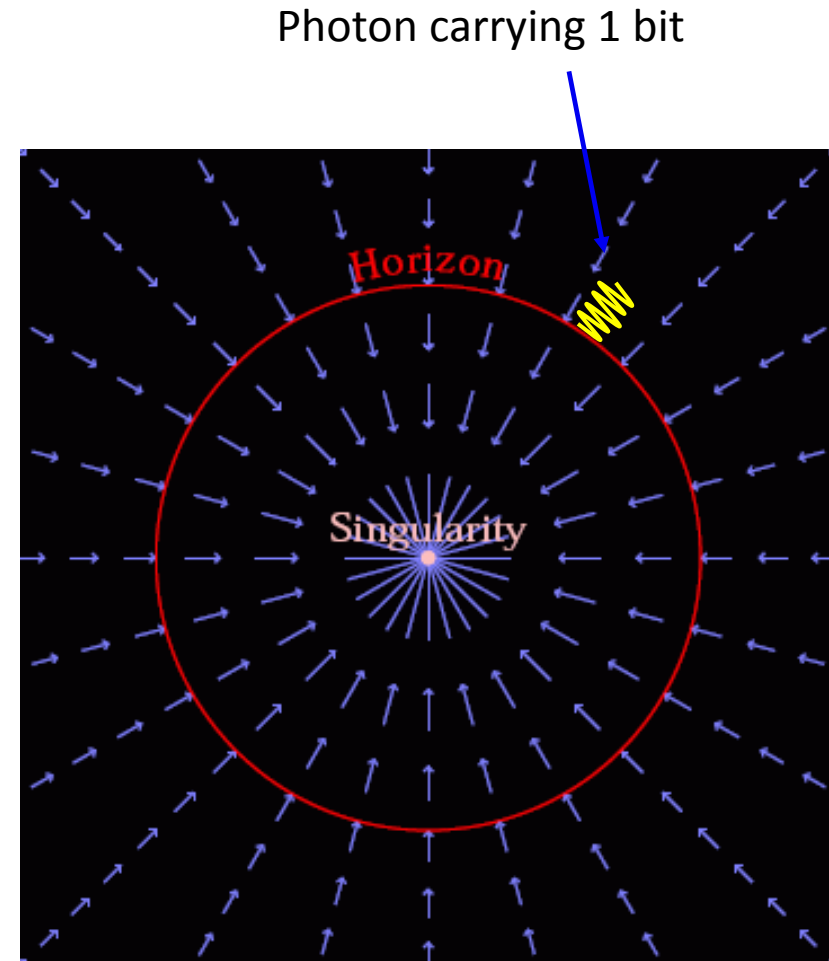
2. **If the information** about what formed the black hole **is not lost** in the interior, then it must somehow be outside the horizon
3. Storing **information** requires **energy**
4. Hence **information gravitates** and tends to fall into the black hole
5. To keep it outside requires applying a **huge force**
6. Huge force is like a ‘**firewall**’ outside the horizon
7. An observer in free fall would **not** perceive the horizon as **a smooth place**
8. **GR breaks down at places of very low curvature**



# Firewall argument – simplified version

2. If the information  
is not lost

8. GR breaks down at places of very low  
curvature



So, black holes  
don't exist, after all?



## A possible way out

*Papadodimas+Raju*

Operators describing the **interior** of the black hole are a **highly-scrambled combination** of operators that describe the **exterior**

The **interior of the bh does not exist** at a fundamental level, but it can be an extremely good **effective description** for certain macroscopic observers

It's quintessentially a theorist's conundrum

Even if LHC had been churning out  $10^7$  bhs/year, extremely hard to decide the problem experimentally

But it remains as a problem of principle and a very confusing one

# Is it leading anywhere, or just a waste of time?

This problem is an old chestnut.

Proposed in 1976, largely ignored in the 1980's.

Brought to fore in the early 1990's

Led to intense activity to understand bhs in string theory

Problem was not solved, but had fruitful spinoffs:

- BH entropy from microscopic stat-mech
  - AdS/CFT and holography

2012-2015: disagreement among parties  
remains significant

But once again, the debate has sparked fruitful  
progress in other directions

This time by emphasizing the role of concepts  
from quantum information

# Quantum Entanglement and Spacetime Geometry

One can convincingly argue that  
the existence of **black holes** with  
**thermodynamic** properties  
leads to **Holography**

Physics at the highest energy densities is  
dominated by black holes,  
with a degeneracy of states  $\propto \text{Area}$  (*not*  $\propto \text{Vol}$ )

Quantum Gravity described by fundamental  
degrees of freedom on the boundary of  
spacetime

This insight is not dependent on String Theory



Lower-dimensional quantum field theory  $\Rightarrow$   
higher-dimensional theory of quantum gravity

QFT at boundary<sub>D</sub>  $\rightarrow$  QG in bulk<sub>D+1</sub>

Extra dimension is emergent: not directly visible  
in field theory

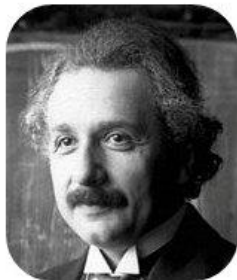
Qualitatively, it corresponds to the energy scale  
of field theory

QFT magnitudes must have a  
gravitational counterpart

- in some cases in terms of classical geometry (not necessarily simple)

Quantum entanglement  
seems to admit simple  
geometric counterparts

# ER = EPR

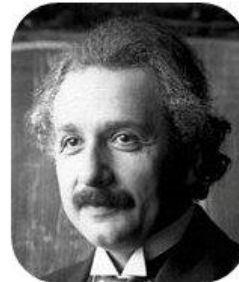


A. Einstein



N. Rosen

=



A. Einstein



B. Podolsky



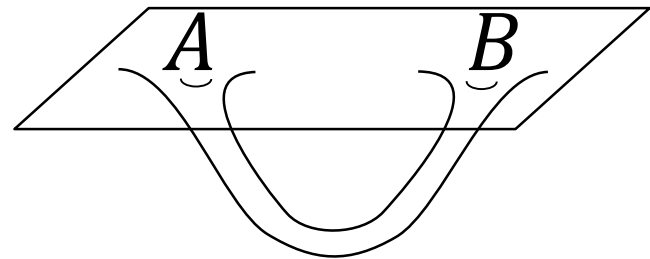
N. Rosen

Maldacena+Susskind 2013

A **non-traversable** (ER) **wormhole** is a geometric encoding of EPR-type quantum correlations

EPR-type quantum correlations are quantum-gravity wormholes

$$\frac{1}{\sqrt{2}}(|0\rangle_A|0\rangle_B + |1\rangle_A|1\rangle_B) \iff$$



Bell pair

Wormhole

## **EPR pairs (Bell states)**

describe spatially-separated entangled states  
cannot be used to send superluminal information  
can be formed by pair-creation (Schwinger)

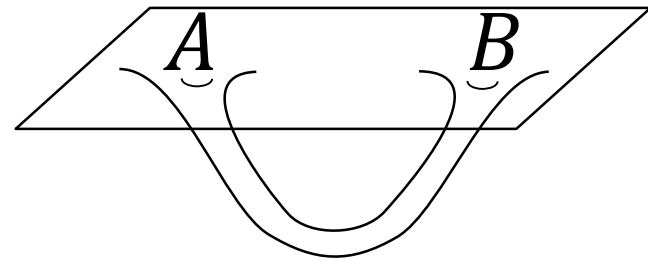
## **ER wormholes**

describe spatially-separated regions  
cannot be crossed by timelike trajectories  
can be formed through pair-creation (eg of black holes)

Few Bell pairs = highly quantum wormhole

Many Bell pairs = (semi)classical wormhole

$$\frac{1}{\sqrt{2}}(|0\rangle_A|0\rangle_B + |1\rangle_A|1\rangle_B) \iff$$



Bell pair

Wormhole

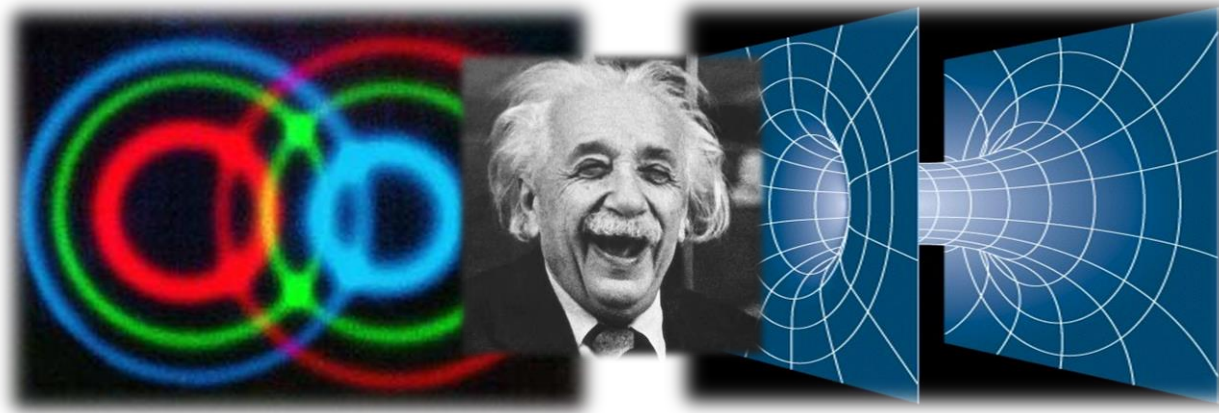


A spacetime is a geometric way of  
encoding quantum correlations  
organized by scale

Einstein understood the fundamental  
importance of both:

spacetime geometry (1915)

quantum entanglement (1935)



**Time to merge them?**

Questions?

