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Recall the Naturalness argument:

$$m_{H}^{2} = \int_{0}^{\infty} dE \frac{dm_{H}^{2}}{dE} (E; p_{\rm FT}) = \int_{0}^{\lesssim \Lambda_{\rm SM}} dE(\ldots) + \int_{\lesssim \Lambda_{\rm SM}}^{\infty} dE(\ldots)$$
$$= \delta_{\rm SM} m_{H}^{2} + \delta_{\rm BSM} m_{H}^{2}$$
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Low Higgs mass-term is Un-Natural because:

1. has dimension 2 > 0 (given that $[H^{\dagger}H] = 2 < 4$)

2. is not protected by a symmetry

If any of the two was violated, no Naturalness problem.

For example, consider the Yukawa couplings:

$$y = \int_0^\infty dE \frac{dy}{dE} (E; p_{\rm FT}) = \int_0^{\lesssim \Lambda_{\rm SM}} dE(\dots) + \int_{\lesssim \Lambda_{\rm SM}}^\infty dE(\dots)$$
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No Problem: $\delta_{\rm SM} y \simeq \frac{g^3 - \log(\Lambda_{\rm SM}/M_{\rm EW})}{16\pi^2}$ some d=0 couplings Moreover, Just dim. an. Moreover, Yukawas are also protected by a symmetry: SM gains Flavor symmetry if y = 0

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 $y_{d,e,\ldots} \ll y_t$ is Natural that's why flavor origin might be at $\Lambda_{SM} \gg TeV$

Other example: a **fermion with mass** (allowed by EW)

$$m_F = \int_0^\infty \frac{dm_F}{dE} (E; p_{\rm FT}) = \int_0^{\lesssim \Lambda} \frac{dE(\ldots) + \int_{\lesssim \Lambda}^\infty dE(\ldots)}{dE(\ldots)}$$
$$= \delta_{\rm IR} m_F + \delta_{\rm UV} m_F$$
$$\qquad \qquad \text{theory with new fermion is not} \\ \text{the SM, } \Lambda \text{ is its own cutoff.} \\ \text{Problem? } \delta_{\rm IR} m_F \stackrel{?}{\simeq} \frac{g^2}{16\pi^2} \Lambda \text{ is allowed by dim. analysis.}$$

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Other example: a fermion with mass (allowed by EW)

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No Problem: $\delta_{\rm IR} m_F \simeq \frac{g^2}{16\pi^2} m_F \log \Lambda$ $m_F \ll \Lambda$ is Natural. Fermions can be Naturally light





















We must break SUSY, easy to do it preserving Naturalness Add stop mass-term $M_{\tilde{t}}^2 |\tilde{t}_{L,R}|^2$, and use dim. an.:

$$\delta_{\mathrm{IR}} m_H^2 \simeq \frac{3}{8\pi^2} \Lambda^2 [y_t^2 - \tilde{y}_t^2] + \frac{3y_t^2}{8\pi^2} M_{\tilde{t}}^2 \log(\Lambda/M_{\mathrm{EW}})$$

SOFT BREAKING

all terms that break SUSY preserving Naturalness

- scalar mass terms $-m_{\phi_i}^2 |\phi_i|^2$, and
- trilinear scalar interactions $-A_{ijk}\phi_i\phi_j\phi_k + h.c.$
- gaugino mass terms $-\frac{1}{2}m_l\bar{\lambda}_l\lambda_l$, where *l* again labels the group factor;
- bilinear terms $-B_{ij}\phi_i\phi_j + h.c.$; and
- linear terms $-C_i\phi_i$.

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• linear terms $-C_i\phi_i$.

only d>0 SUST parameters (not all d>0) for sure you **cannot** have $y_t \neq \tilde{y}_t$.

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The "low-energy SUSY" picture for high energy physics



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- Unobserved sparticles have soft masses
- Soft breaking generated at $\Lambda_{\rm SUSY}$

The "low-energy SUSY" picture for high energy physics



Natural SUSY cannot hide above the TeV scale.

The "low-energy SUSY" picture for high energy physics



Natural SUSY cannot hide above the TeV scale. General tuning estimate **worsened** by the log term.

SUSY and GUT:

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 $au(p
ightarrow \pi^0 e^+) \sim 10^{34} \mathrm{years} \left(\frac{3 \, 10^{15} \mathrm{GeV}}{M_{\mathrm{GUT}}}
ight)^4$ GUT excluded in SM. Viable in SUSY-SM

SUSY, R-Parity and DM: (or, how a problem turns in a virtue)

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SUSY, R-Parity and DM: (or, how a problem turns in a virtue) **Problem:** B# and L# violation allowed in SUSY at d=4. (unlike in SM, no Accidental Symmetries) **Solution: R-Parity** (imposed on both SUSY and soft terms) $\theta \rightarrow -\theta$ $SF_{matter} \rightarrow -SF_{matter}$ $SF_{Higgs,gauge} \rightarrow +SF_{Higgs,gauge}$ $\Phi = \phi(y) + \sqrt{2}\theta\psi(y) - \theta\theta F(y)$ $V = \theta \sigma^{\mu} \overline{\theta} A_{\mu}(x) + i \theta \theta \overline{\theta} \overline{\lambda}(x) - i \overline{\theta} \overline{\theta} \theta \lambda(x) + \frac{1}{2} \theta \theta \overline{\theta} \overline{\theta} D(x)$ **SM** particles are **EVEN** Associated **Sparticles are ODD**
Other virtues of SUSY

SUSY, R-Parity and DM: (or, how a problem turns in a virtue)

Problem: B# and L# violation allowed in SUSY at d=4. (unlike in SM, no Accidental Symmetries)

Solution: R-Parity (imposed on both SUSY and soft terms)



Virtue: lightest SUSY particle (LSP) is stable (can be DM, thanks to WIMP Miracle)

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- 1. A Natural theory
- 2. Viable GUT
- 3. A Dark Matter Candidate
- 4. A string-friendly UV picture

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In these lectures I hope to convince you that supersymmetry (SUSY)¹, will soon provide you with a whole new spectroscopy to investigate. Indeed, it may even be that experiments^{2,3} are already starting to reveal this spectroscopy to us.

> From John Ellis' lecture notes: supersymmetry --spectroscopy of the future? or of the present? 1984

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This led to "some" excitement about SUSY ...

"If SUSY will not be found at LEP, I will cut my ba..."

Riccardo Barbieri private communication (secondhand) circa 1989

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The excitement WAS JUSTIFIED.

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This led to "some" excitement about SUSY ...

The excitement **WAS JUSTIFIED.** But life is not that easy...

In SUSY, fields are promoted to **SuperFields.** One would thus naively expect:

SM Higgs field

 $H \in \mathbf{2}_{1/2}$



SUSY Higgs SF $\Phi \in \mathbf{2}_{1/2}$

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In SM we can freely use conjugate *H*: $H^c = i\sigma_2 H^*$ $\mathcal{L}_{Y}^u = y_u q_L H u_R^c$ $\mathcal{L}_{Y}^d = y_d q_L H^c d_R^c$ In SUSY instead we use Superpotential $W[\Phi, \Phi_R^*]$ $W_Y^u = y_u \Phi_{q_L} \Phi_u \Phi_{u_R^c}$ $W_Y^d = y_d \Phi_{q_L} \Phi_d \Phi_{d_R^c}$

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The SUSY Higgses scalar potential:

$$V[H_{\rm u}, H_{\rm d}] = \mu^2 \left[|H_{\rm u}|^2 + |H_{\rm d}|^2 \right] + \frac{g^2 + g'^2}{8} \left[|H_{\rm u}|^2 - |H_{\rm d}|^2 \right]^2 + \frac{g^2}{2} |H_{\rm u}^{\dagger} H_{\rm d}|^2 + m_{\rm u}^2 |H_{\rm u}|^2 + m_{\rm d}^2 |H_{\rm d}|^2 + B \left[H_{\rm u} H_{\rm d} + H_{\rm u}^* H_{\rm d}^* \right]$$

The SUSY Higgses scalar potential:

 $\begin{array}{c} \hline \mathbf{F}\text{-}\mathbf{Term} \left(|\partial W / \partial \Phi|^2 \right) \\ from \\ W = \mu \Phi_{\mathrm{u}} \Phi_{\mathrm{d}} \end{array} \\ V[H_{\mathrm{u}}, H_{\mathrm{d}}] = \hline \mu^2 \left[|H_{\mathrm{u}}|^2 + |H_{\mathrm{d}}|^2 \right] \\ + \frac{g^2 + g'^2}{8} \left[|H_{\mathrm{u}}|^2 - |H_{\mathrm{d}}|^2 \right]^2 + \frac{g^2}{2} |H_{\mathrm{u}}^{\dagger} H_{\mathrm{d}}|^2 \\ + m_{\mathrm{u}}^2 |H_{\mathrm{u}}|^2 + m_{\mathrm{d}}^2 |H_{\mathrm{d}}|^2 + B \left[H_{\mathrm{u}} H_{\mathrm{d}} + H_{\mathrm{u}}^* H_{\mathrm{d}}^* \right] \end{array}$

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Particular case of generic 2 Higgs doublet model

Four implications of the SUSY Higgs sector structure

Four implications of the SUSY Higgs sector structure. Implication #0: (actually 5 impl.) vacuum is viable (no e.m., color, L and B breaking)

$$\langle |H_{\rm u}|^2 \rangle = \frac{v_{\rm u}^2}{2}$$
 $\langle |H_{\rm d}|^2 \rangle = \frac{v_{\rm d}^2}{2}$
2 sources of EWSB $v_{\rm u}^2 + v_{\rm d}^2 = v^2 = (246 \,{\rm GeV})^2$

$$\langle |H_{\rm u}|^2 \rangle = \frac{v_{\rm u}^2}{2} \qquad \langle |H_{\rm d}|^2 \rangle = \frac{v_{\rm d}^2}{2}$$
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define: $v_{\rm u}/v_{\rm d} = \tan \beta$

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Abbreviations:
 $s_{\beta} = \sin \beta \\ c_{\beta} = \cos \beta \end{cases} t_{\beta} = \frac{s_{\beta}}{c_{\beta}}$
Both Higgses must take VEV, for u and d-type masses:
$$\mathcal{L}_{\rm Y}^{\rm u} = y_{\rm u}q_LH_{\rm u}u_R^c \\ \mathcal{L}_{\rm Y}^{\rm d} = y_{\rm d}q_LH_{\rm d}d_R^c \end{cases} \longrightarrow \begin{cases} m_{\rm u} = y_{\rm u}v_{\rm u}/\sqrt{2} \\ m_{\rm d} = y_{\rm d}v_{\rm d}/\sqrt{2} \end{cases}$$

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2 sources of EWSB $v_{\rm u}^2 + v_{\rm d}^2 = v^2 = (246 \text{GeV})^2$
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Both Higgses must take VEV, for u and d-type masses:
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For $y_{\rm u,d} < 4\pi$ (perturbative): $0.08 \simeq \frac{y_{\rm top}^{\rm SM}}{4\pi} \lesssim t_{\beta} \lesssim \frac{4\pi}{y_{\rm bot}^{\rm SM}} \simeq 500$

Four implications of the SUSY Higgs sector structure. Implication #2: many scalars around

In Unitary Gauge $H_{\rm u} = \begin{bmatrix} 0 \\ \frac{v_{\rm u} + h_{\rm u}}{\sqrt{2}} \end{bmatrix} + \begin{bmatrix} c_{\beta} H_{+} \\ c_{\beta} \frac{iA}{\sqrt{2}} \end{bmatrix} \qquad H_{\rm d} = \begin{bmatrix} \frac{v_{\rm d} + h_{\rm d}}{\sqrt{2}} \\ 0 \end{bmatrix} + \begin{bmatrix} s_{\beta} \frac{iA}{\sqrt{2}} \\ s_{\beta} H_{-} \end{bmatrix}$

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 $H_+ = (H_-)^*$: one **charged** scalar

A : one **neutral** pseudo-scalar (CP-odd)

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 $H_+=(H_-)^*$: one charged scalar

- *A* : one **neutral** pseudo-scalar (CP-odd)
- $h_{\rm u,d}~$: two **neutral** scalars

$$\begin{bmatrix} h_{\rm u} \\ h_{\rm d} \end{bmatrix} = \begin{bmatrix} \cos \alpha & \sin \alpha \\ -\sin \alpha & \cos \alpha \end{bmatrix} \begin{bmatrix} h \\ H \end{bmatrix}$$

Four implications of the SUSY Higgs sector structure. Implication #2: many scalars around

In Unitary Gauge $H_{\rm u} = \begin{vmatrix} 0 \\ \frac{v_{\rm u} + h_{\rm u}}{\sqrt{2}} \end{vmatrix} + \begin{vmatrix} c_{\beta} H_{+} \\ c_{\beta} \frac{iA}{\sqrt{2}} \end{vmatrix} \qquad \qquad H_{\rm d} = \begin{vmatrix} \frac{v_{\rm d} + h_{\rm d}}{\sqrt{2}} \\ 0 \end{vmatrix} + \begin{vmatrix} s_{\beta} \frac{iA}{\sqrt{2}} \\ s_{\beta} H_{-} \end{vmatrix}$ $H_{+}=(H_{-})^{*}$: one **charged** scalar A : one **neutral** pseudo-scalar (CP-odd) $h_{u,d}$: two **neutral** scalars The Higgs we saw $m_h = 125 \text{GeV}$ $\begin{bmatrix} h_{\rm u} \\ h_{\rm d} \end{bmatrix} = \begin{bmatrix} \cos \alpha & \sin \alpha \\ -\sin \alpha & \cos \alpha \end{bmatrix} \begin{vmatrix} h \\ H \end{vmatrix}$ The Other Higgs (maybe heavier)

$$\kappa_{\rm u} = \frac{g_{h\rm uu}}{g_{h\rm uu}^{\rm SM}} = \frac{\sin(\alpha + \pi/2)}{\sin\beta}$$

$$\kappa_{\rm d} = \frac{g_{h\rm dd}}{g_{h\rm dd}^{\rm SM}} = \frac{\cos(\alpha + \pi/2)}{\cos\beta}$$

$$\kappa_{V} = \frac{g_{hVV}}{g_{hVV}^{\rm SM}} = \sin(\beta - \alpha)$$
The form of the potential allows us to express α in terms of β and of the pseudo-scalar A mass:

$$\tan \alpha = \frac{(m_{A}^{2} + m_{Z}^{2})t_{\beta}}{m_{h}^{2}(1 + t_{\beta}^{2}) - m_{Z}^{2} - m_{A}^{2}t_{\beta}^{2}}$$

Four implications of the SUSY Higgs sector structure. Implication #3: modified Higgs couplings



ATLAS arXiv:1509.00672

Direct scalar searches play an important role in this plane.

Four implications of the SUSY Higgs sector structure. Implication #3: modified Higgs couplings

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The form of the potential allows us to express α in terms of β and of the pseudo-scalar A mass:

$$\tan \alpha = \frac{(m_{A}^{2} + m_{Z}^{2})t_{\beta}}{m_{h}^{2}(1 + t_{\beta}^{2}) - m_{Z}^{2} - m_{A}^{2}t_{\beta}^{2}}$$
Decoupling limit: $m_{\rm d}^{2} \to \infty$ (technically natural)

$$m_{A}^{2} = m_{\rm d}^{2} + \ldots \to \infty \Longrightarrow \tan \alpha \simeq -\frac{1}{4}$$

 t_{β}

$$\kappa_{\rm u} = \frac{g_{huu}}{g_{huu}^{\rm SM}} = \frac{\sin(\alpha + \pi/2)}{\sin\beta}$$

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Decoupling limit: $m_{\rm d}^{2} \to \infty$ (technically natural)
 $m_{A}^{2} = m_{\rm d}^{2} + \ldots \to \infty \longrightarrow \tan \alpha \simeq -\frac{1}{t_{\beta}} \longrightarrow \alpha \simeq \beta - \pi/2 \longrightarrow$ SM Higgs

$$\kappa_{\rm u} = \frac{g_{h\rm uu}}{g_{h\rm uu}^{\rm SM}} = \frac{\sin(\alpha + \pi/2)}{\sin\beta}$$

$$\kappa_{\rm d} = \frac{g_{h\rm dd}}{g_{h\rm dd}^{\rm SM}} = \frac{\cos(\alpha + \pi/2)}{\cos\beta}$$

$$\kappa_{V} = \frac{g_{hVV}}{g_{hVV}^{\rm SM}} = \sin(\beta - \alpha)$$
The form of the potential allows us to express α in terms of β and of the pseudo-scalar A mass:

$$\tan \alpha = \frac{(m_{A}^{2} + m_{Z}^{2})t_{\beta}}{m_{h}^{2}(1 + t_{\beta}^{2}) - m_{Z}^{2} - m_{A}^{2}t_{\beta}^{2}}$$
Decoupling limit: $m_{\rm d}^{2} \to \infty$ (technically natural)
 $m_{A}^{2} = m_{\rm d}^{2} + \ldots \to \infty \Longrightarrow \tan \alpha \simeq -\frac{1}{t_{\beta}} \Longrightarrow \alpha \simeq \beta - \pi/2 \Longrightarrow$ SM Higgs
In the limit we also have: $\sin 2\beta = \frac{2B}{m_{A}^{2}} \Rightarrow t_{\beta} \simeq \frac{m_{A}^{2}}{B} \to \infty$



Four implications of the SUSY Higgs sector structure. Implication #4: **wrong Higgs mass !!**

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Problem: λ is too small. **Solution:** increase λ .

 $\lambda \to \lambda + \delta \lambda$ $\delta \lambda = \frac{m_H^2 - m_Z^2}{2v^2} \simeq 0.06$

Two ways to increase λ :

First way: rely on large loop corrections (only way in MSSM)



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... which is exponentially bad for tuning:

$$\Delta \ge \left(\frac{M_{\text{soft}}}{500 \text{ GeV}}\right)^2 \log(\Lambda_{\text{SUSY}}/M_{\text{EW}})$$

$$\downarrow \log(\Lambda_{\text{SUSY}} = 10 \text{ TeV}$$

SUSY and Naturalness

The "low-energy SUSY" picture for high energy physics



Natural SUSY cannot hide above the TeV scale. General tuning estimate **worsened** by the log term.



 $\Delta\gtrsim100$





 $\Delta\gtrsim 100$

Higgs discovery) a **Natural** theory.

moreover ...

LHC discovery not expected (heavy spart.) even if true.

from arXiv:1112.2703



Second way to make m_H right: Add an extra singlet SF. (NMSSM or λ SUSY)

$$W_S = \lambda_S \Phi_S \Phi_u \Phi_d \quad \longrightarrow \quad V_S = \lambda_S^2 |H_u H_d|^2$$

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Interesting to study **Higgs couplings** and **extra scalars** in this framework.

Caveat: needed values of $\lambda_S \sim 1$ give ~10 TeV cutoff.

Higgs **should be lighter** in Natural MSSM:

Due to peculiar structure of Higgs potential, at tree-level: $\lambda = \frac{g^2 + g'^2}{8}$

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MSSM is **Un-Natural** with 125 GeV Higgs!

Higgs **should be lighter** in Natural MSSM:



 $\Delta \gtrsim 100$ The MSSM is **not anymore** (after Higgs discovery) a **Natural** theory. moreover ... LHC discovery not expected (heavy spart.) even if true. look for SUSY beyond MSSM !

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Impressive search program Quantitative illustration



Qualitative illustration



from arXiv:1309.0528

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Naturalness is one of those questions, not the only one.

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Experimentalists should not **blindly trust** theorists. They should **critically listen** to theorists. And get convinced (or not). Nobody has the truth.

