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Quantum Field Theory and Standard Model

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Universidad de Salamanca

Taller de Altas Energías 2017, Benasque



An “ideal” list of topics

- * The very basics of QFT
- * Loops, divergences and renormalization
- * What to ask from a “healthy” QFT
 - Lorentz invariance
 - Locality
 - Unitarity
 - Renormalizability?
- * Symmetries and their breaking
- * Gauge invariance
- * Massive gauge fields
- * Building the standard model

Schedule

* **Lectures:** Monday to Thursday (first week), from 9:00 to 10:00.

* **Tutorials:**

- ◆ Monday 4th, 15:30 to 16:30
- ◆ Tuesday 5th, 16:30 to 17:30
- ◆ Thursday 7th, 18:00 to 19:00

Tutor: Ramon Miravitllas.

A sample of textbooks

- * L. Álvarez-Gaumé & M.A. Vázquez-Mozo, “An Invitation to Quantum Field Theory”, Springer 2012.
- * M.E. Peskin & D.V. Schroeder, “An Introduction to Quantum Field Theory”, Perseus Books 1995.
- * C. Quigg, “Gauge theories of the strong, weak, and electromagnetic interactions” (2nd edition), Princeton University Press 2013.
- * M. Schwartz, “Quantum field theory and the standard model”, Cambridge 2014.

A note about conventions:

* We use the “**mostly minus**” metric (a.k.a. West Coast metric):

$$\eta_{\mu\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}$$

* Unless otherwise said, **natural units** are used throughout:

$$\hbar = c = 1$$

* We use **Heaviside-Lorentz** electromagnetic units:

$$\nabla \cdot \mathbf{E} = \rho$$

$$\nabla \cdot \mathbf{B} = 0$$

$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}$$

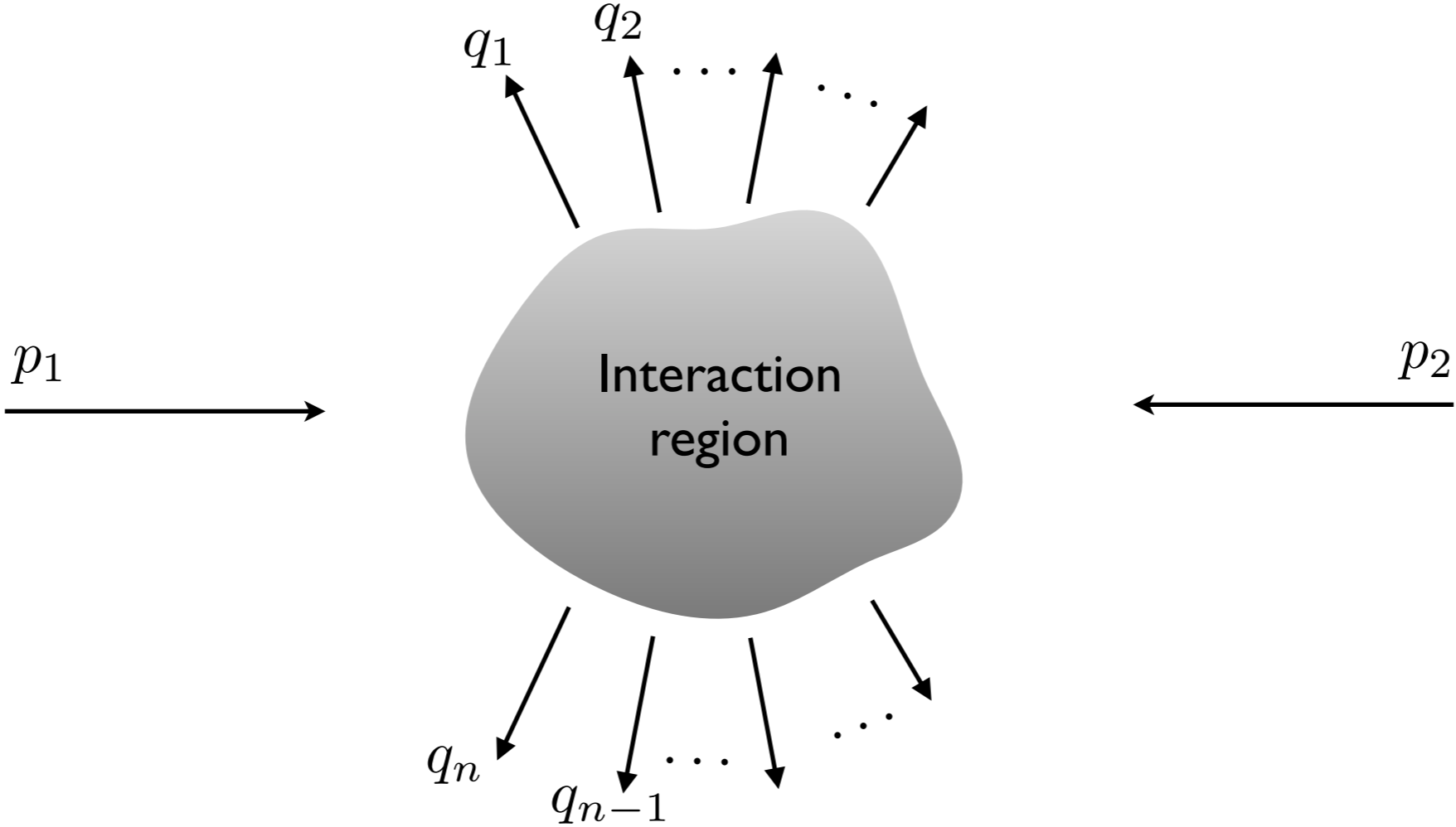
$$\nabla \times \mathbf{B} = \mathbf{j} + \frac{\partial \mathbf{E}}{\partial t}$$

$$\alpha = \frac{e^2}{4\pi} \quad (\text{fine structure constant})$$

Part I

From elementary particles to quantum fields

Elementary particles are studied through **scattering experiments**, typically



Quantum mechanics, even relativistic, is **not enough** to describe these **high energy** experiments...

Let us consider the **relativistic** quantum evolution of a **localized, single-particle wave packet**:

$$\psi(t, \mathbf{x}) = e^{-it \overbrace{\sqrt{-\nabla^2 + m^2}}^{H = \sqrt{p^2 + m^2}}} \delta^{(3)}(\mathbf{x}) = \int \frac{d^3 k}{(2\pi)^3} e^{i\mathbf{k} \cdot \mathbf{x} - it \sqrt{k^2 + m^2}}$$

positive frequency

$$= -\frac{i}{4\pi^2 |\mathbf{x}|} \int_{-\infty}^{\infty} k dk e^{ik|\mathbf{x}| - it \sqrt{k^2 + m^2}} = \frac{1}{2\pi^2 |\mathbf{x}|} \int_0^{\infty} k dk \sin(k|\mathbf{x}|) e^{-it \sqrt{k^2 + m^2}}$$

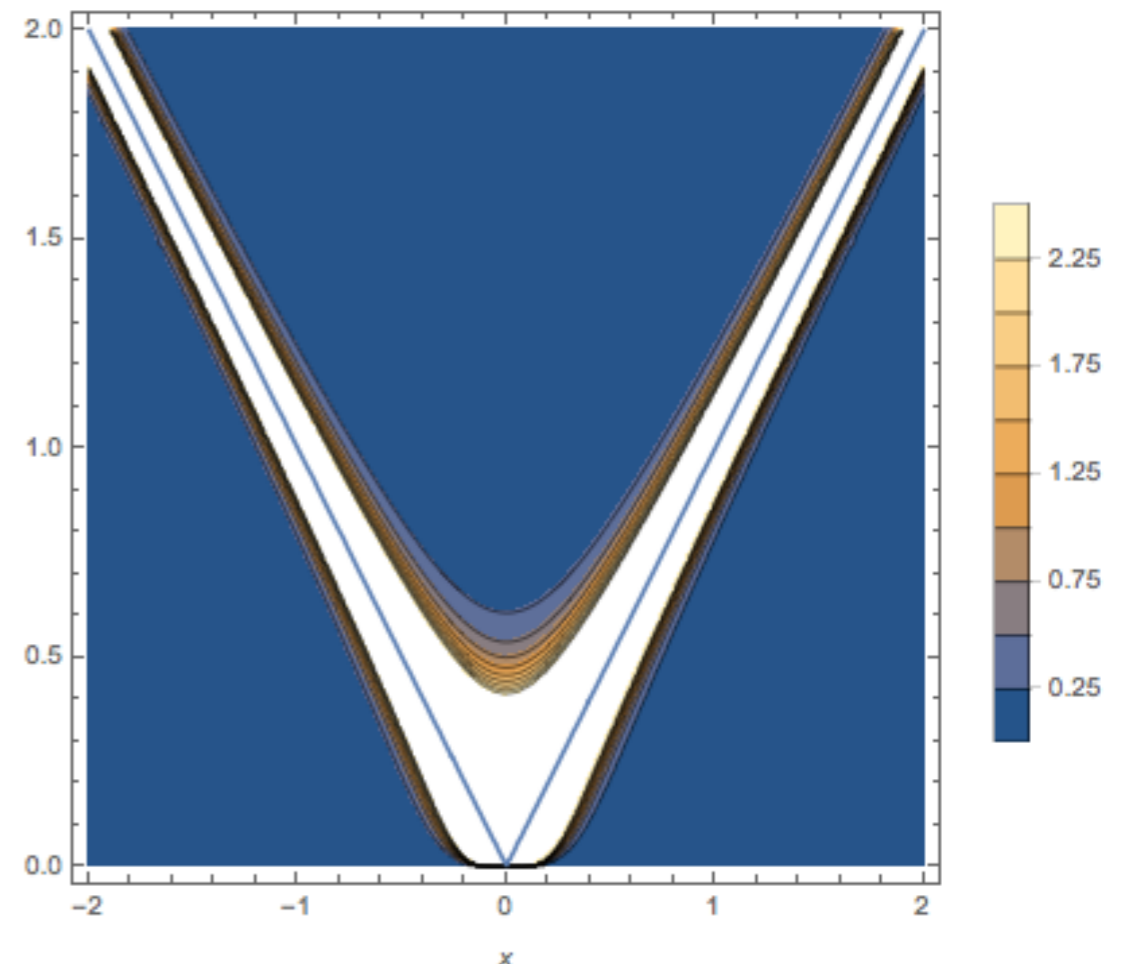
The integral can be regularized by $t \rightarrow t - i\epsilon$, to give

$$\psi(t, \mathbf{x}) = -\frac{i}{2\pi^2} \frac{m^2 t}{t^2 - \mathbf{x}^2} K_2\left(im \sqrt{t^2 - \mathbf{x}^2}\right)$$

The probability $|\psi(t, \mathbf{x})|^2$ **spills** outside the light-cone



Causality is violated!



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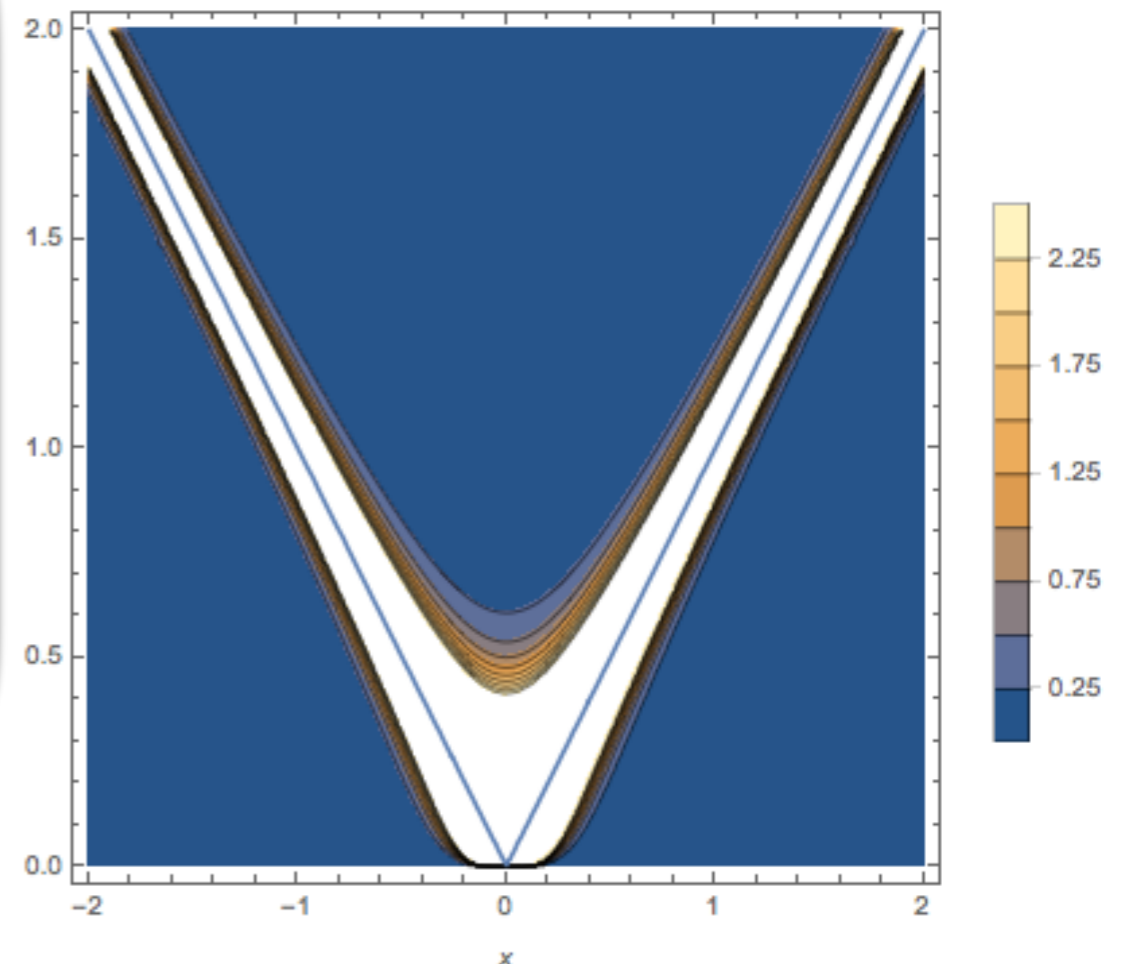
What we are computing is the **propagator** of a relativistic particle:

$$\psi(t, \mathbf{x}) = \langle \mathbf{x} | e^{-itH} | \mathbf{0} \rangle \equiv G(t, \mathbf{x}; 0, \mathbf{0})$$

Relativistic quantum mechanics propagates states **outside** the light cone

$$G(t', \mathbf{x}'; t, \mathbf{x}) \neq 0 \quad \text{when} \quad (t' - t)^2 - (\mathbf{x}' - \mathbf{x})^2 < 0$$

Causality is violated!



But when $t^2 - \mathbf{x}^2 < 0$ there are **frames** in which (t, \mathbf{x}) **happens before** $(0, \mathbf{0})$. Thus,

$$\psi(t, \mathbf{x}) = \begin{cases} \langle \mathbf{x} | e^{-itH} | \mathbf{0} \rangle & \text{when } t^2 - \mathbf{x}^2 > 0 \\ \langle \mathbf{x} | e^{-itH} | \mathbf{0} \rangle + \langle \mathbf{0} | e^{itH} | \mathbf{x} \rangle = 2\text{Re} \langle \mathbf{x} | e^{-itH} | \mathbf{0} \rangle & \text{when } t^2 - \mathbf{x}^2 < 0 \end{cases}$$

But since we have computed

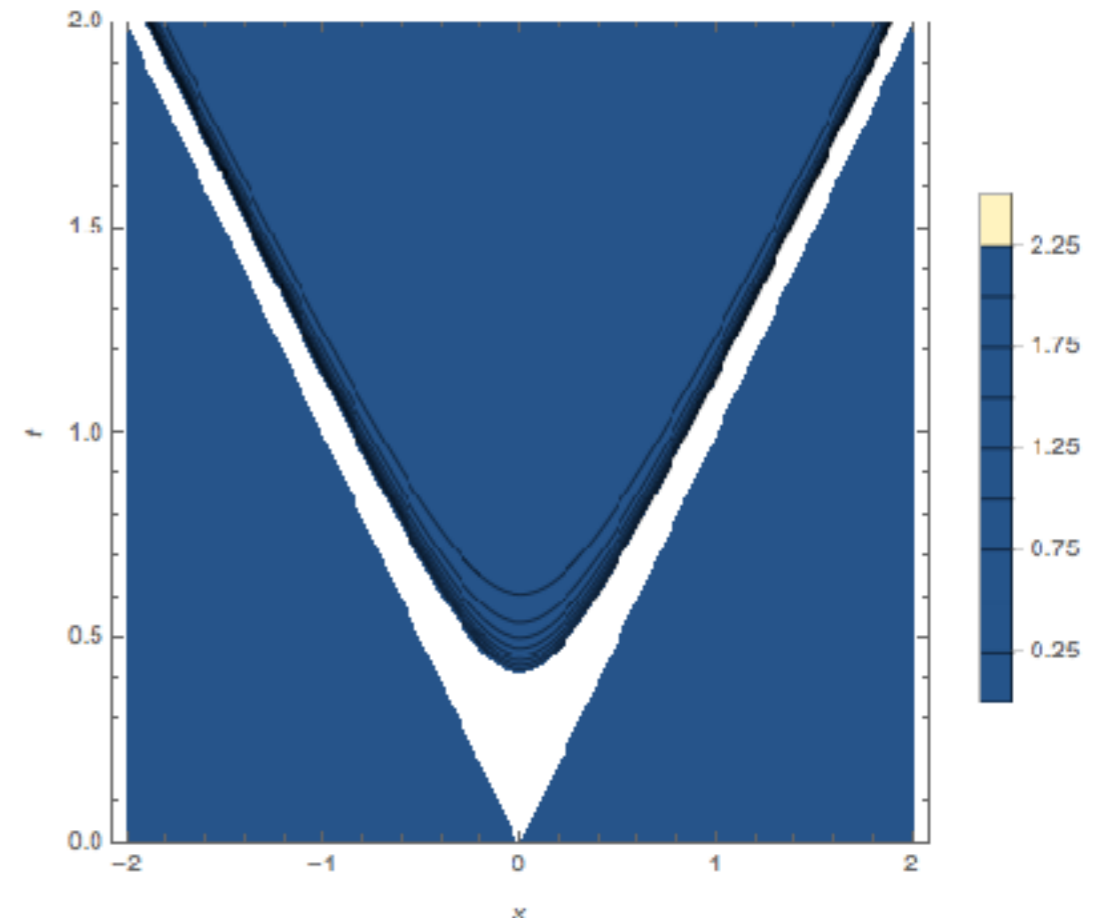
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the result is

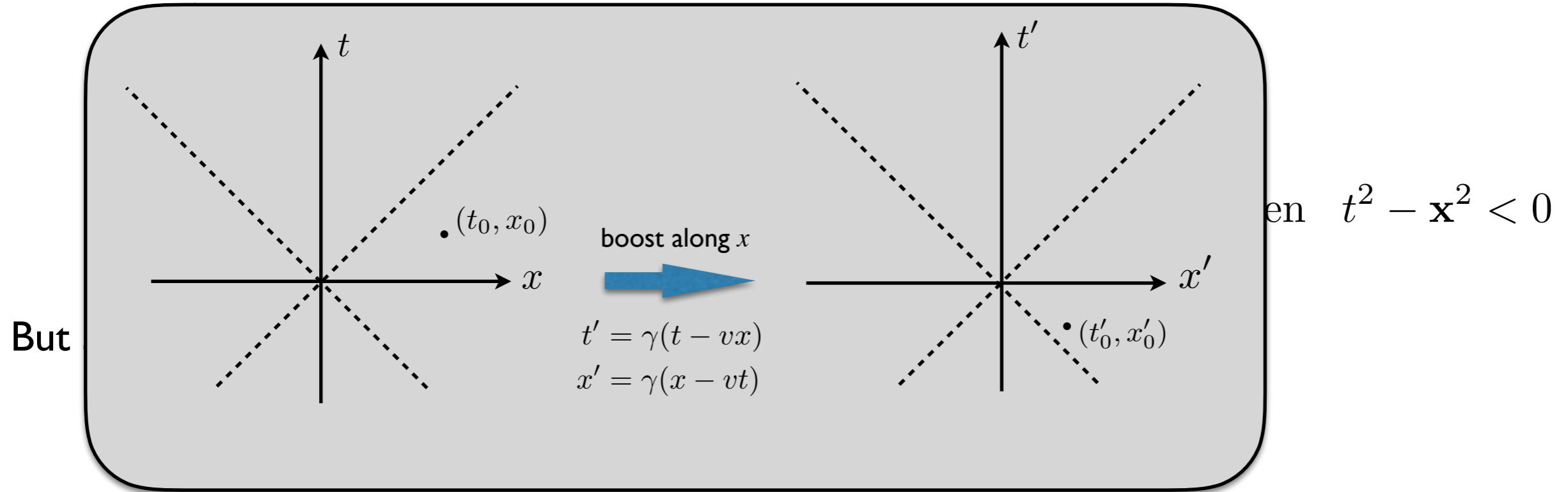
$$\psi(t, \mathbf{x}) = -\frac{i}{2\pi^2} \frac{m^2 t}{t^2 - \mathbf{x}^2} K_2 \left(im \sqrt{t^2 - \mathbf{x}^2} \right) \theta(t^2 - \mathbf{x}^2)$$



Causality is restored!



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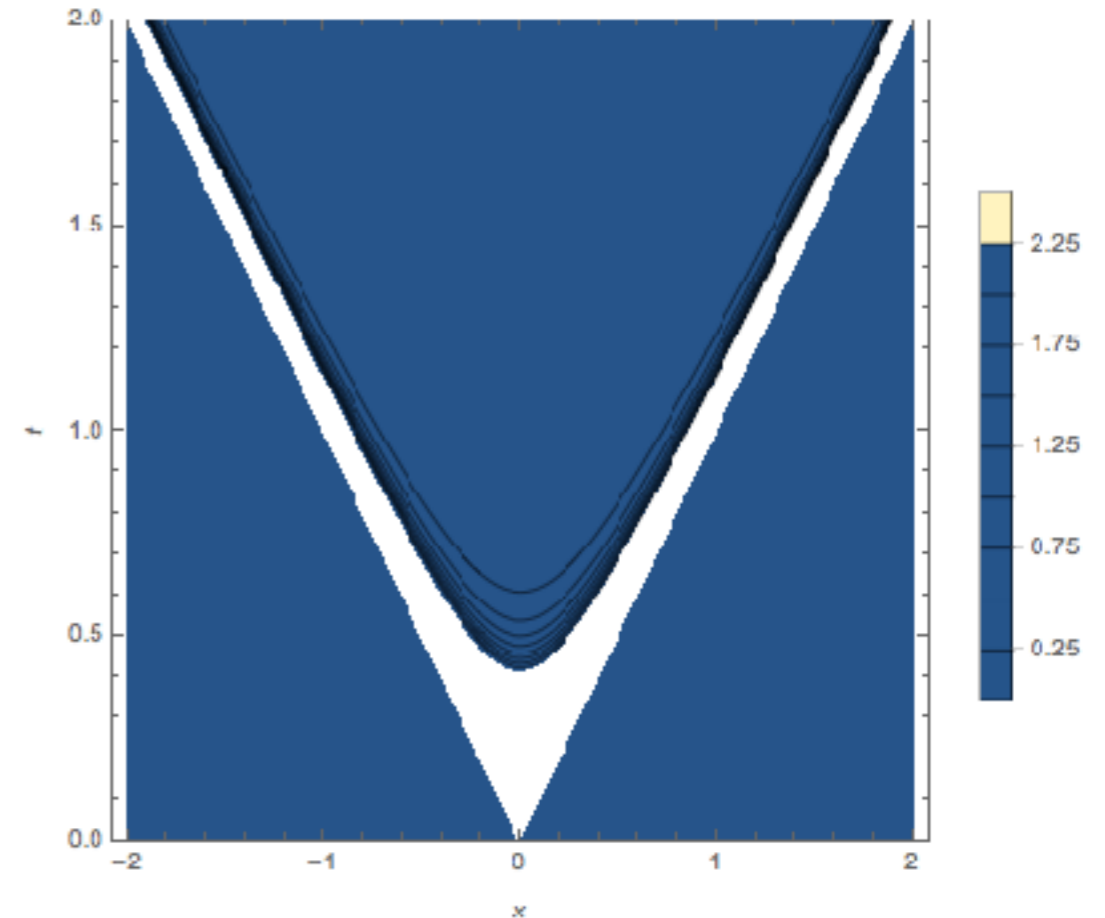


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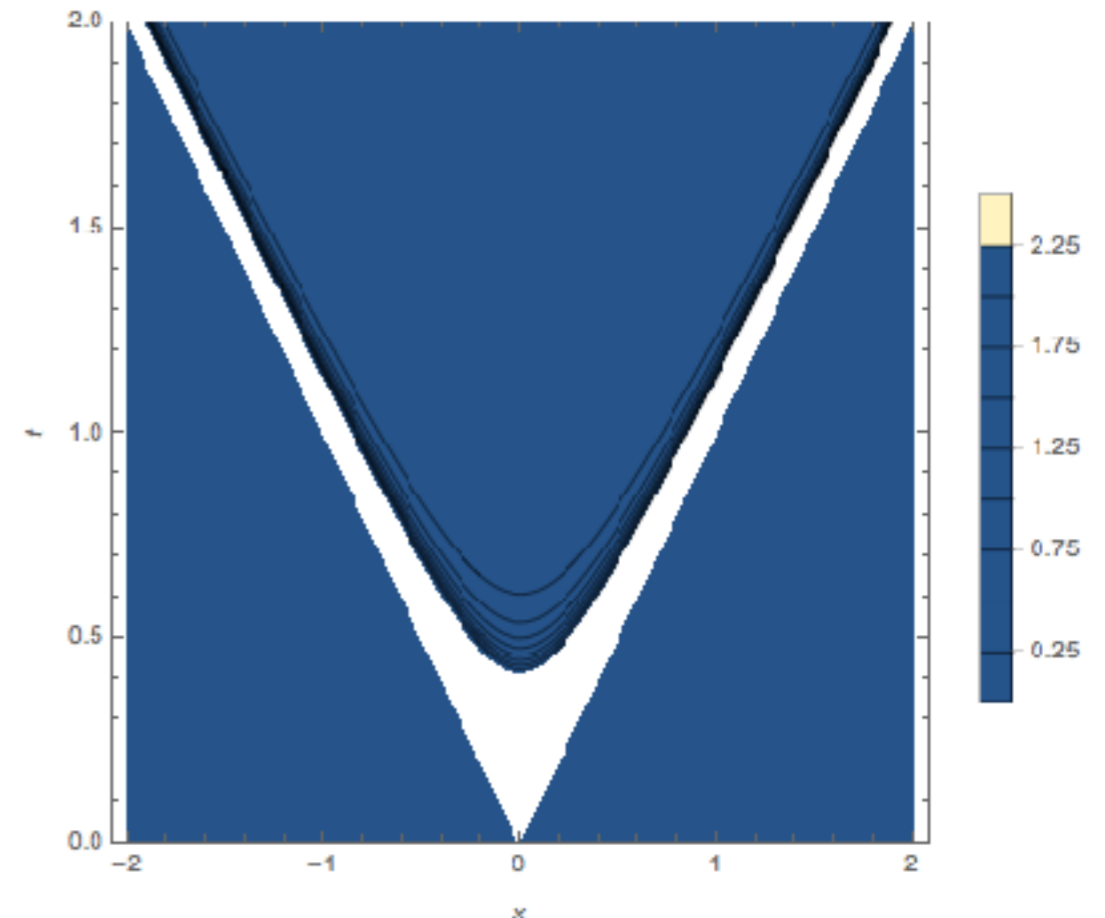
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Causality is restored!



We have to allow **particles travelling backward in time!!**

Their wave functions are

$$\psi(t, \mathbf{x})_{\downarrow} = \langle \mathbf{0} | e^{itH} | \mathbf{x} \rangle = \langle \mathbf{x} | e^{-itH} | \mathbf{0} \rangle^* = \psi(t, \mathbf{x})_{\uparrow}^*$$

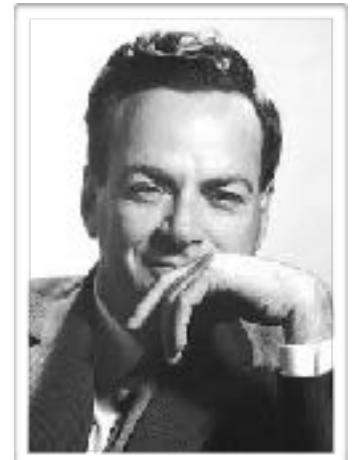
Thus, under any **global U(1) symmetry**

$$\psi(t, \mathbf{x})_{\uparrow} \longrightarrow e^{iq\theta} \psi(t, \mathbf{x})_{\uparrow} \quad \longrightarrow \quad \psi(t, \mathbf{x})_{\downarrow} \longrightarrow e^{-iq\theta} \psi(t, \mathbf{x})_{\downarrow}$$

these particles have **opposite charges**, $q_{\downarrow} = -q_{\uparrow}$ (but the same mass! $H_{\uparrow, \downarrow} = \sqrt{-\nabla^2 + m^2}$)



Ernst Stückelberg
(1905-1984)



Richard Feynman
(1918-1988)

To **restore causality** we are forced to introduce **antiparticles!!**

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Ernst Stückelberg
(1905-1984)



Richard Feynman
(1981-1988)

States moving backward in time can be reinterpreted as **negative frequency states** with **reversed** momentum, propagating **forward** in time:

$$\begin{aligned} \psi(t, \mathbf{x})_{\downarrow} &= \int \frac{d^3k}{(2\pi)^3} e^{-i\mathbf{k}\cdot\mathbf{x} + it\sqrt{k^2 + m^2}} \\ &= \int \frac{d^3k}{(2\pi)^3} e^{i(-\mathbf{k})\cdot\mathbf{x} - it(-\sqrt{k^2 + m^2})} \end{aligned}$$

negative frequency

$$\rightarrow e^{-iq\theta} \psi(t, \mathbf{x})_{\downarrow}$$

mass! $H_{\uparrow, \downarrow} = \sqrt{-\nabla^2 + m^2}$

To **restore causality** we are forced to introduce **antiparticles!!**

Switching on **interactions**, charge conservation allows the creation of **particle-antiparticle pairs**, provided **enough energy** is available.

For example, localizing particle below their **Compton wavelength**

$$\Delta x \sim \frac{1}{m} \quad \xrightarrow{\Delta x \Delta p \sim 1} \quad \Delta p \sim m \quad \xrightarrow{\quad} \quad \Delta E \sim m$$

and due to energy **quantum fluctuations** the creation of particle-antiparticle pairs **cannot be prevented**.



We have to **give up** the **single-particle description!**

Relativistic quantum mechanics is a **dead end** for high energy particle physics...

To handle many particles, **second quantization** seems the best approach, introducing **creation-annihilation operators** for particles with **on-shell momentum** p

$$a(p), a(p)^\dagger \quad \begin{array}{c} \xrightarrow{p^2 = m^2} \\ \xrightarrow{\omega_{\mathbf{p}} = \sqrt{\mathbf{p}^2 + m^2}} \end{array} \quad \begin{array}{l} [a(p), a(p')^\dagger] = (2\pi)^3 (2\omega_{\mathbf{p}}) \delta^{(3)}(\mathbf{p} - \mathbf{p}') \\ [a(p), a(p')] = [a(p)^\dagger, a(p')^\dagger] = 0 \end{array}$$

Lorentz invariant (exercise)

(Multi-)particle states are obtained from the **Poincaré-invariant vacuum** $|0\rangle$

$$|p\rangle = a(p)^\dagger |0\rangle \quad \longrightarrow \quad \langle p|p'\rangle = (2\pi)^3 (2\omega_{\mathbf{p}}) \delta^{(3)}(\mathbf{p} - \mathbf{p}')$$

Lorentz invariant (exercise)

$$|f\rangle = \int \left[\prod_{i=1}^n \frac{d^3 p_i}{(2\pi)^3} \frac{1}{2\omega_{\mathbf{p}_i}} \right] f(\mathbf{p}_1, \dots, \mathbf{p}_n) a(p_1)^\dagger \dots a(p_n)^\dagger |0\rangle$$

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Lorentz invariant (exercise)

$$\mathcal{U}(\Lambda)|0\rangle = e^{-ia \cdot P}|0\rangle = |0\rangle$$

$$\mathcal{U}(\Lambda)a(p)\mathcal{U}(\Lambda)^\dagger = a(\Lambda p)$$

$$\mathcal{U}(\Lambda)|p\rangle = |\Lambda p\rangle$$

where $\mathcal{U}(\Lambda) \in \text{SO}(1,3)$

$$|\mathbf{p}_1, \dots, \mathbf{p}_n\rangle = a(p_1)^\dagger \dots a(p_n)^\dagger |0\rangle$$

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Free fields are **linear combinations** of creation-annihilation operators. E.g., for a free **Hermitian** scalar field

$$\phi(x) = \phi(x)^\dagger \quad \longrightarrow \quad \phi(x) = \int \frac{d^3p}{(2\pi)^3} \frac{1}{2\omega_{\mathbf{p}}} \left[f(x, \mathbf{p}) a(p) + f(x, \mathbf{p})^* a(p)^\dagger \right]$$

Imposing the **equations of motion**,

$$(\square + m^2)\phi(x) = 0 \quad \longrightarrow \quad f(x, \mathbf{p}) = e^{-i\omega_{\mathbf{p}}t + i\mathbf{p}\cdot\mathbf{x}}$$

The free quantum field satisfies:

*** Equal-time** canonical commutation relations

$$[\phi(t, \mathbf{x}), \dot{\phi}(t, \mathbf{x}')] = i\delta^{(3)}(\mathbf{x} - \mathbf{x}'), \quad [\phi(t, \mathbf{x}), \phi(t, \mathbf{x}')] = [\dot{\phi}(t, \mathbf{x}), \dot{\phi}(t, \mathbf{x}')] = 0$$

*** Microcausality**

$$[\phi(x), \phi(x')] = 0 \quad \text{when} \quad (x - x')^2 < 0$$

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positive frequency
 negative frequency

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The **many-particle** Fock states **diagonalize** the free field **Hamiltonian**

$$H = \frac{1}{2} \int d^3x \left[\dot{\phi}^2 + (\nabla\phi)^2 + m^2\phi^2 \right] \xrightarrow{\text{(exercise)}} H = \frac{1}{2} \int \frac{d^3p}{(2\pi)^3} \left[a(p)^\dagger a(p) + (2\pi)^3 \omega_{\mathbf{p}} \delta^{(3)}(\mathbf{0}) \right]$$

$$= \int \frac{d^3p}{(2\pi)^3} \frac{1}{2\omega_{\mathbf{p}}} \left[\omega_{\mathbf{p}} a(p)^\dagger a(p) \right] + E_0$$

Subtracting the (divergent) zero-point energy E_0

$$[a(p), a(p')^\dagger] = (2\pi)^3 (2\omega_{\mathbf{p}}) \delta^{(3)}(\mathbf{p} - \mathbf{p}')$$

$$H|p\rangle = \frac{1}{2} \int \frac{d^3k}{(2\pi)^3} a(k)^\dagger a(k) a(p)^\dagger |0\rangle = \omega_{\mathbf{p}} |p\rangle$$

$$H|p_1, \dots, p_n\rangle \equiv H a(p_1)^\dagger \dots a(p_n)^\dagger |0\rangle = \left(\sum_{i=1}^n \omega_{\mathbf{p}_i} \right) |p_1, \dots, p_n\rangle$$



Particles are the **low-lying excitations** of **quantum fields**

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$$E_0 = \langle 0|H|0\rangle = \sum_{\mathbf{p}} \frac{1}{2} \omega_{\mathbf{p}}$$

$$= \int \frac{d^3p}{(2\pi)^3} \frac{1}{2\omega_{\mathbf{p}}} \left[\omega_{\mathbf{p}} a(p)^\dagger a(p) \right] + E_0$$

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Particles are the **low-lying excitations** of **quantum fields**

A **particle** is characterized by a number of “**Casimirs**”:

*Poincaré group: $\left\{ \begin{array}{l} P_\mu P^\mu = m^2 \\ W_\mu W^\mu = -m^2 s(s+1) \end{array} \right.$ \rightarrow Mass
 \rightarrow Spin

$$W^\mu = \frac{1}{2} \epsilon^{\mu\nu\alpha\beta} J_{\nu\alpha} P_\beta$$

vector de Pauli-Lubański

*Internal symmetry groups: $\left\{ \begin{array}{l} \text{electric charge} \\ \vdots \end{array} \right.$

To do particle physics, we have to **choose** the appropriate **interpolating field**:

*It **transforms correctly** (i.e., the right value for the “Casimirs”)

*It **creates** the corresponding particle out of the **vacuum**:

$$\langle 0 | \phi(x) | p \rangle \neq 0$$

The x -dependence is **fixed** by the Poincaré invariance of the vacuum

$$\langle 0 | \phi(x) | p \rangle = \langle 0 | e^{iP \cdot x} \phi(0) e^{-iP \cdot x} | p \rangle = \langle 0 | \phi(0) | p \rangle e^{-ip \cdot x}$$

The fields can be **canonically normalized**, such that:

***Scalar field:** $\langle 0 | \phi(0) | p \rangle = 1$

***Dirac field:**
$$\begin{cases} \langle 0 | \psi_\alpha(0) | p, \sigma; 0 \rangle = u_\alpha^{(\sigma)}(p) \\ \langle 0 | \bar{\psi}_\alpha(0) | 0; p, \sigma \rangle = \bar{v}_\alpha^{(\sigma)}(p) \end{cases}$$

***Photon field:** $\langle 0 | A_\mu(0) | p, \lambda \rangle = \varepsilon_\mu^{(\lambda)}(p)$

Any properly normalized interpolating field **does the job**, provided it satisfies **microcausality**

$$[\phi(x), \phi(x')] = 0 \quad \text{when} \quad (x - x')^2 < 0$$

 Borchers classes

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To describe scalar particles, instead of $\phi(x)$ we can also use

$$\Phi(x) = -\frac{1}{m^2} \square \phi(x)$$



$$\langle 0 | \Phi(x) | p \rangle = -\frac{1}{m^2} \square \langle 0 | \phi(x) | p \rangle = -\frac{1}{m^2} \square e^{-ip \cdot x} = e^{-ip \cdot x}$$

and

$$[\Phi(x), \Phi(x')] = \frac{1}{m^4} \square_x \square_{x'} [\phi(x), \phi(x')] = 0$$

for $(x - x')^2 < 0$

$\sigma) (p)$

$\sigma) (p)$

, provided it satisfies

< 0

Borchers classes

Still, to study particle physics we need to **introduce interactions...**

In interacting field theories, **particles** still emerge as **weakly coupled excitations:**

$$\mathcal{L}(\phi, \partial\phi) = \mathcal{L}(\phi, \partial\phi)_{\text{free}} + \sum_i g_i \mathcal{O}_i(\phi, \partial\phi)$$

dimensionless effective couplings

nonquadratic terms

where g_i is **“small”** and

$$\langle \Omega | \phi(x) | p \rangle = \phi(\mathbf{p}) e^{-ip \cdot x}$$

ground state of the **full** theory

one-particle wave function

Thus:

- * Particles are identified by **quantizing** the **free theory**.
- * Interactions are treated in **perturbation theory**.

Still, to study particle physics we need to **introduce interactions...**

In interacting field theories, **particles** still emerge as **weakly coupled excitations:**

Examples

***QED:** electrons+photons

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} + \bar{\psi}(\gamma^\mu\partial_\mu - m)\psi - qA_\mu\bar{\psi}\gamma^\mu\psi$$

***QCD (high energies):** quarks+gluons

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}^a F^{a\mu\nu} + \sum_{f=1}^6 \bar{Q}_f(\gamma^\mu\partial_\mu - m_f)Q_f - gA_\mu^a \sum_{f=1}^6 \bar{Q}_f T^a \gamma^\mu Q_f$$

***QCD (low energies):** pions (+nucleons)

$$\mathcal{L} = \frac{1}{2}\text{tr}\left(\partial_\mu\pi\partial^\mu\pi\right) - \frac{1}{3f_\pi^2}\text{tr}\left(\partial_\mu\pi[\pi, [\pi, \partial^\mu\pi]]\right) + \dots$$

*** Interactions are treated in perturbation theory.**

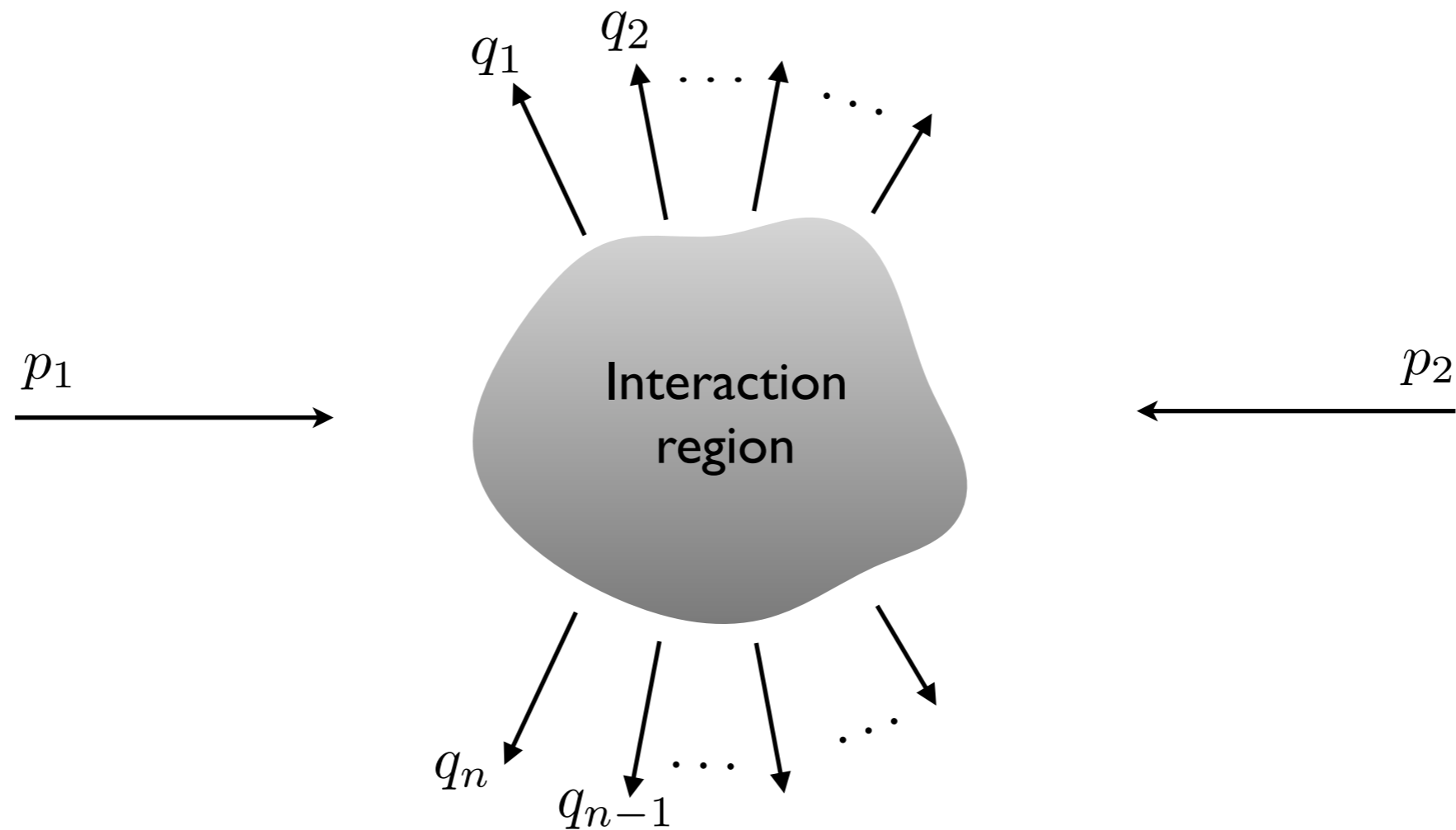
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A **scattering** experiment is characterized by its initial (*in*) and final (*out*) multiparticle state:

$$|p_1, p_2\rangle_{\text{in}} \qquad |q_1, q_2, \dots, q_{n-1}, q_n\rangle_{\text{out}}$$

Both are **Heisenberg-picture** (i.e., time-independent) **states** in a very complicated **interacting theory**.

Our aim is to **compute** the **probability amplitude**:

$$S(i \longrightarrow f) = {}_{\text{out}} \langle q_1, \dots, q_n | p_1, p_2 \rangle_{\text{in}}$$

$$|p_1, p_2\rangle_{\text{in}}$$

$$|q_1, q_2, \dots, q_{n-1}, q_n\rangle_{\text{out}}$$

These states can also be seen as belonging to the **free, multiparticle Fock space**

$$|p_1, p_2\rangle, |q_1, q_2, \dots, q_n\rangle \in \mathcal{F} \equiv \bigoplus_{n=0}^{\infty} \mathcal{H}_1 \otimes \dots \otimes \mathcal{H}_1$$

The scattering experiment is then described by the **S-matrix operator**

$$S : \mathcal{F} \longrightarrow \mathcal{F}$$



$$S(i \longrightarrow f) = {}_{\text{out}}\langle q_1, \dots, q_n | p_1, p_2 \rangle_{\text{in}} \equiv \langle q_1, \dots, q_n | S | p_1, p_2 \rangle$$

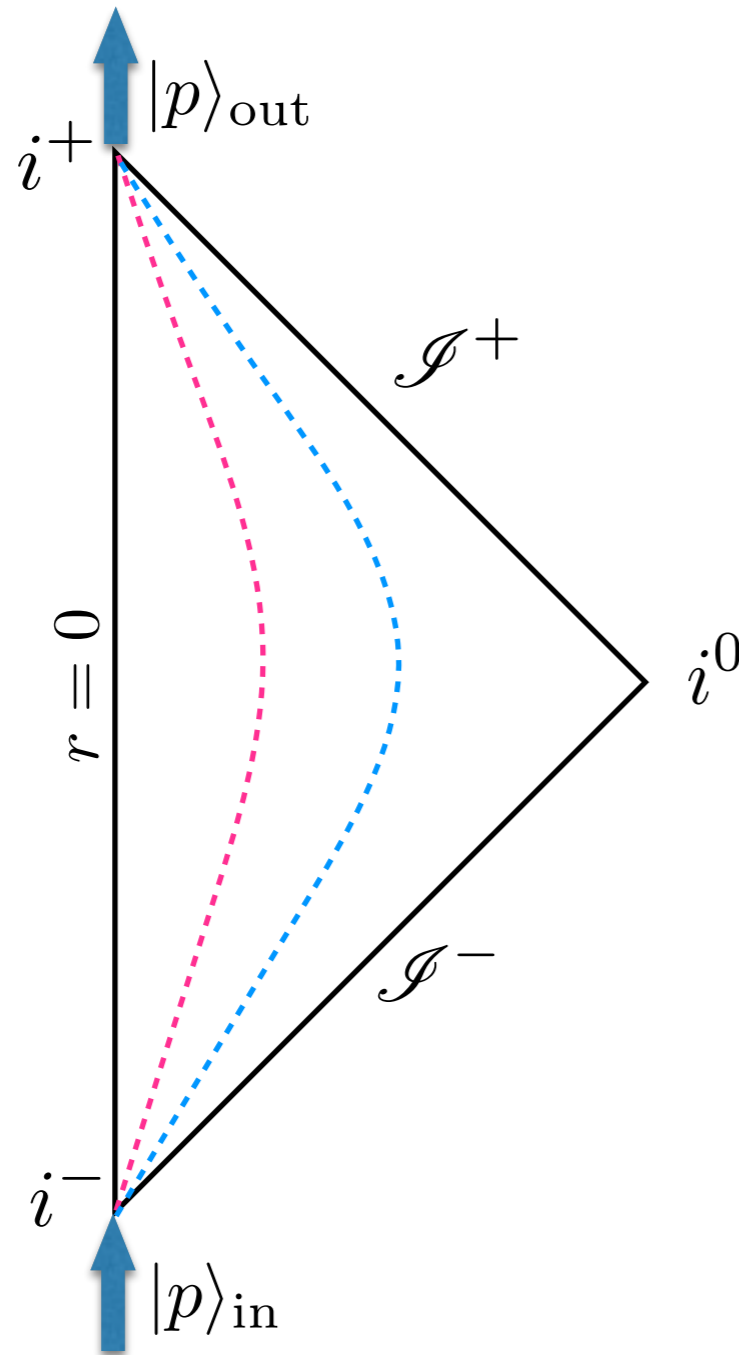
The **S-matrix** operator satisfies a number of **properties**:

* **Unitarity:** $S^\dagger = S^{-1}$

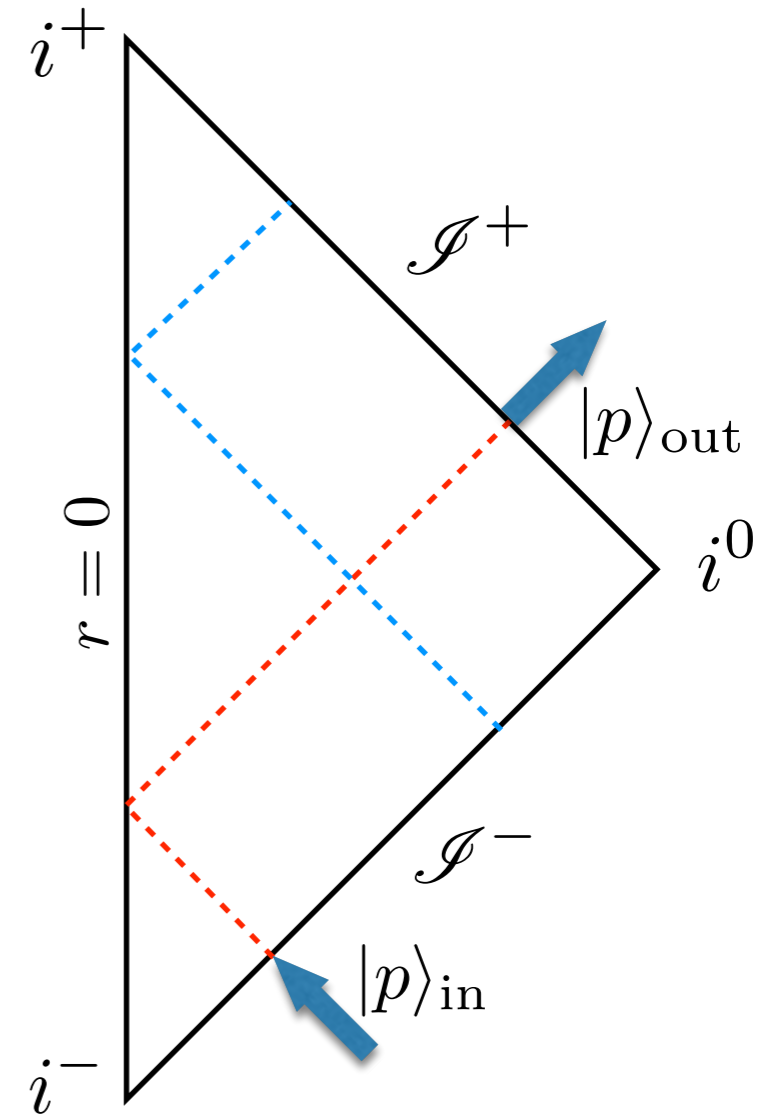
* **Lorentz invariance:** $\mathcal{U}(\Lambda) S \mathcal{U}(\Lambda)^\dagger = S$ with $\Lambda \in \text{SO}(1,3)$

* $\langle q_1, \dots, q_n | S | p_1, p_2 \rangle$ is **analytic** in the external momenta.

The **S-matrix** is a kind of **holographic** quantity in **Minkowski** space-time: *in*- and *out*-states **live** on its **boundary**.



massive states



massless states

The computation of the S-matrix in terms of the interacting field theory is done using the **Lehmann-Symanzik-Zimmermann (LSZ) reduction formula**.



Harry Lehmann
(1924-1998)



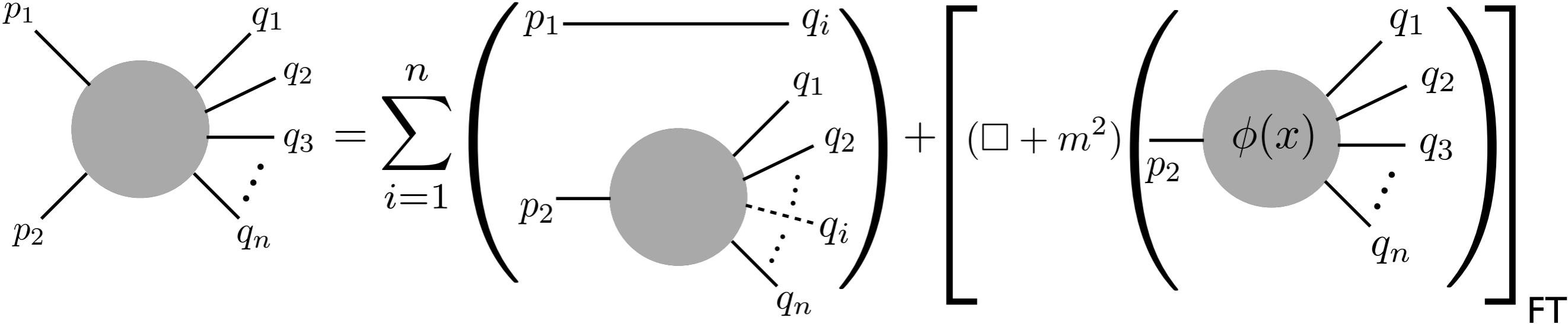
Kurt Symanzik
(1923-1983)



Wolfhart Zimmermann
(1928-2016)

$$\text{out} \langle q_1, \dots, q_n | p_1, p_2 \rangle_{\text{in}} = \sum_{i=1}^n (2\pi)^3 (2\omega_{\mathbf{q}_i}) \delta^{(3)}(\mathbf{q}_i - \mathbf{p}_1) \text{out} \langle q_1, \dots, \hat{q}_i, \dots, q_n | p_2 \rangle_{\text{in}} \\ + iZ^{-1/2} \int d^4x e^{-ip_1 \cdot x} (\square + m^2) \text{out} \langle q_1, \dots, q_n | \phi(x) | p_2 \rangle_{\text{in}}$$

Symbolically:

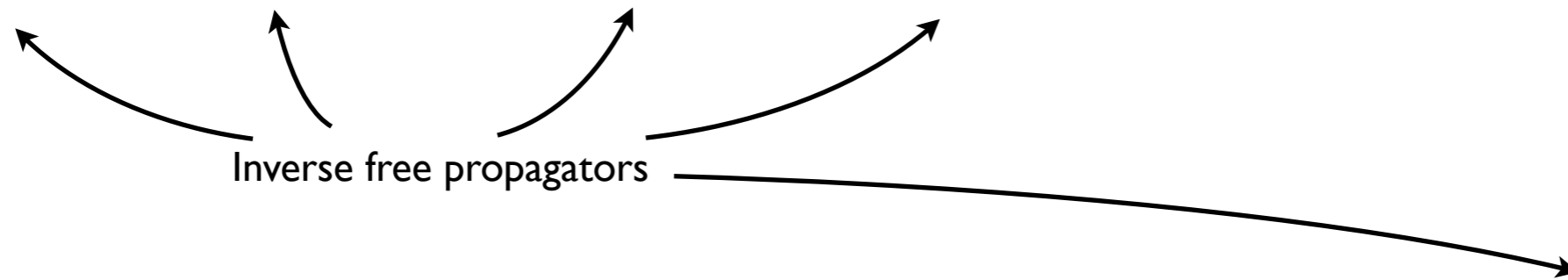


Iterating the procedure, we **trade** all incoming and outgoing particles by **time-ordered** field insertions:

$$\langle q_1, \dots, q_n | S | p_1, p_2 \rangle = \text{disconnected terms}$$

$$+ i(Z^{-1/2})^{n+2} \int d^4x_1 d^4x_2 e^{-ip_1x_1 - ip_2x_2} \int d^4y_1 \dots d^4y_n e^{iq_1y_1 + \dots + iq_ny_n}$$

$$\times (\square + m^2)_{x_1} (\square + m^2)_{x_2} (\square + m^2)_{y_1} \dots (\square + m^2)_{y_n} \langle \Omega | T[\phi(x_1)\phi(x_2)\phi(y_1) \dots \phi(y_n)] | \Omega \rangle$$



S-matrix amplitudes are computed in terms of **time-ordered** (amputated) **correlation functions**

$$G(x_1, \dots, x_n) = \langle \Omega | T[\phi(x_1) \dots \phi(x_n)] | \Omega \rangle$$

which can be computed in **perturbation theory**.

Feynman diagrammatics

Iterating the procedure, we **trade** all incoming field insertions:

$$T[\phi(x)\phi(y)] = \theta(x^0 - y^0)\phi(x)\phi(y) + \theta(y^0 - x^0)\phi(y)\phi(x)$$

field

$\langle q_1, \dots, q_n | S | p_1, p_2 \rangle =$ disconnected terms

$$+ i(Z^{-1/2})^{n+2} \int d^4x_1 d^4x_2 e^{-ip_1x_1 - ip_2x_2} \int d^4y_1 \dots d^4y_n e^{iq_1y_1 + \dots + iq_ny_n}$$

$$\times (\square + m^2)_{x_1} (\square + m^2)_{x_2} (\square + m^2)_{y_1} \dots (\square + m^2)_{y_n} \langle \Omega | T[\phi(x_1)\phi(x_2)\phi(y_1) \dots \phi(y_n)] | \Omega \rangle$$



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Feynman diagrammatics


We can isolate **nontrivial scattering** in the S-matrix by writing

$$S = \mathbf{1} + iT$$

so the matrix elements have the structure

$$\begin{aligned} \langle q_1, \dots, q_n | S | p_1, p_2 \rangle &= \langle q_1, \dots, q_n | p_1, p_2 \rangle + \langle q_1, \dots, q_n | iT | p_1, p_2 \rangle \\ &= \langle q_1, \dots, q_n | p_1, p_2 \rangle + (2\pi)^4 \delta^{(4)} \left(p_1 + p_2 - \sum_{i=1}^n q_i \right) i\mathcal{M}_{i \rightarrow f} \end{aligned}$$

invariant
amplitude



In terms of the **invariant amplitude**, the **differential cross section** is given by

$$d\sigma = \underbrace{\frac{|i\mathcal{M}_{i \rightarrow f}|^2}{4\omega_{\mathbf{p}_1} \omega_{\mathbf{p}_2} |\mathbf{v}_1 - \mathbf{v}_2|}}_{\text{observer dependent}} (2\pi)^4 \delta^{(4)} \left(p_1 + p_2 - \sum_{i=1}^n q_i \right) \underbrace{\prod_{k=1}^n \frac{d^3 q_k}{(2\pi)^3} \frac{1}{2\omega_{\mathbf{q}_k}}}_{\text{phase space factor}}$$

We can isolate **nontrivial scattering** in the S-matrix by writing

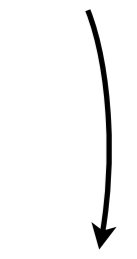
In the case of **particle decay**,

$$|p\rangle \longrightarrow |q_1, \dots, q_n\rangle$$

the **decay width** is given by

$$d\Gamma = \frac{|i\mathcal{M}_{i \rightarrow f}|^2}{2\omega_{\mathbf{p}}} (2\pi)^4 \delta^{(4)}\left(p_1 + p_2 - \sum_{i=1}^n q_i\right) \prod_{k=1}^n \frac{d^3 q_k}{(2\pi)^3} \frac{1}{2\omega_{\mathbf{q}_k}}$$

invariant
amplitude



$$\left(\sum_{i=1}^n q_i \right) i\mathcal{M}_{i \rightarrow f}$$

given by

$$d\sigma = \underbrace{\frac{|i\mathcal{M}_{i \rightarrow f}|^2}{4\omega_{\mathbf{p}_1} \omega_{\mathbf{p}_2} |\mathbf{v}_1 - \mathbf{v}_2|}}_{\text{observer dependent}} (2\pi)^4 \delta^{(4)}\left(p_1 + p_2 - \sum_{i=1}^n q_i\right) \underbrace{\prod_{k=1}^n \frac{d^3 q_k}{(2\pi)^3} \frac{1}{2\omega_{\mathbf{q}_k}}}_{\text{phase space factor}}$$

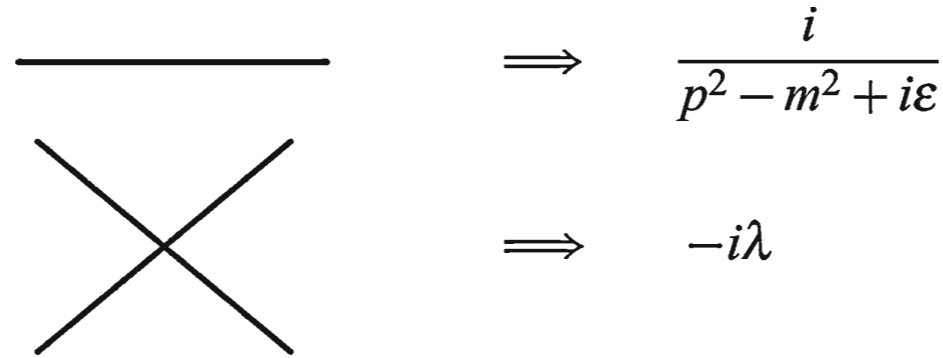
so t

$\langle q_1,$

In ter

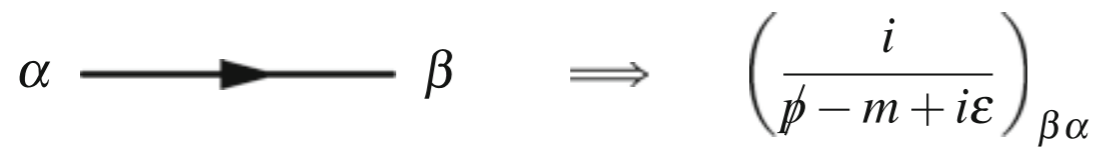
The **perturbative computation** of correlation functions in **momentum space** is carried out using **Feynman diagrammatics**

For a ϕ^4 scalar theory, the **Feynman rules** are



Richard Feynman
(1981-1988)

whereas for **QED**



Incoming fermion:

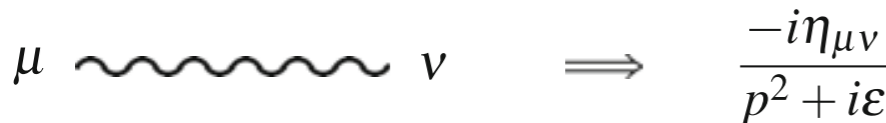


$$\Rightarrow u_{\alpha}(\mathbf{p}, s)$$

Incoming antifermion:



$$\Rightarrow \bar{v}_{\alpha}(\mathbf{p}, s)$$



Outgoing fermion:



$$\Rightarrow \bar{u}_{\alpha}(\mathbf{p}, s)$$

Outgoing antifermion:



$$\Rightarrow v_{\alpha}(\mathbf{p}, s)$$

Incoming photon:

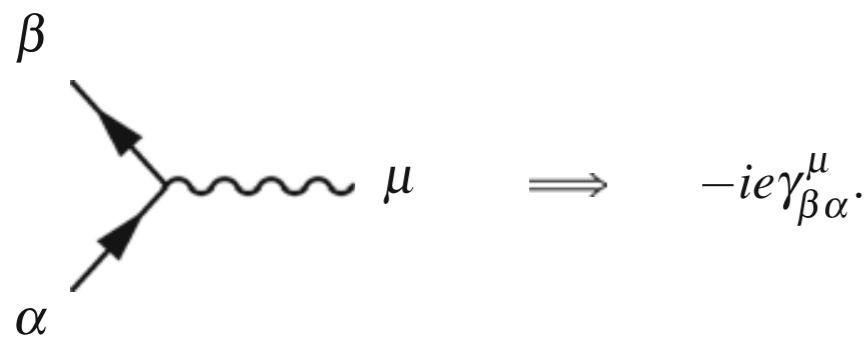


$$\Rightarrow \epsilon_{\mu}(\mathbf{p})$$

Outgoing photon:



$$\Rightarrow \epsilon_{\mu}(\mathbf{p})^*$$



+ **integration** over internal momenta, a **delta function momentum conservation** at each **vertex**, a **factor of -1** for each fermion loop, and a **combinatorial factor**.

As an example, for **Compton scattering**

$$\gamma(k, \varepsilon) + e^-(p, s) \longrightarrow \gamma(k', \varepsilon') + e^-(p', s')$$

the invariant amplitude at leading $\mathcal{O}(e^2)$ is given by

$$\begin{aligned}
 i\mathcal{M}_{i \rightarrow f} = & \quad \text{[Feynman diagrams: s-channel and u-channel Compton scattering]} \\
 = & u_\alpha(\mathbf{p}, s) \varepsilon_\mu(\mathbf{k}) (-ie\gamma_{\beta\alpha}^\mu) \left(\frac{i}{\not{p} + \not{k} - m} \right)_{\sigma\beta} (-ie\gamma_{\lambda\sigma}^\nu) \varepsilon'_\nu(\mathbf{k}')^* \bar{u}_\lambda(\mathbf{p}', s') \\
 & + u_\alpha(\mathbf{p}, s) \varepsilon'_\mu(\mathbf{k}')^* (-ie\gamma_{\beta\alpha}^\mu) \left(\frac{i}{\not{p} - \not{k} - m} \right)_{\sigma\beta} (-ie\gamma_{\lambda\sigma}^\nu) \varepsilon_\nu(\mathbf{k}) \bar{u}_\lambda(\mathbf{p}', s') \\
 = & -ie^2 \bar{u}(\mathbf{p}', s') \not{\varepsilon}'(\mathbf{k}')^* \frac{\not{p} + \not{k} + m}{(p+k)^2 - m^2} \not{\varepsilon}(\mathbf{k}) u(\mathbf{p}, s) \\
 & -ie^2 \bar{u}(\mathbf{p}', s') \not{\varepsilon}(\mathbf{k}) \frac{\not{p} - \not{k} + m}{(p-k)^2 - m^2} \not{\varepsilon}'(\mathbf{k}')^* u(\mathbf{p}, s)
 \end{aligned}$$

Remember:

$$\cancel{A} \equiv A_\mu \gamma^\mu$$

$$p^2 = p'^2 = m^2$$

$$k^2 = k'^2 = 0$$

$$k \cdot \varepsilon(\mathbf{k}) = k' \cdot \varepsilon(\mathbf{k}') = 0$$

In the **low energy limit** $p^2, p'^2, k^2, k'^2 \ll m^2$ the invariant amplitude is

$$i\mathcal{M}_{i \rightarrow f} = \frac{ie^2}{m} \left[\varepsilon(\mathbf{k}) \cdot \varepsilon'(\mathbf{k}') \right] \bar{u}(\mathbf{p}', s') \frac{\not{k}}{|\mathbf{k}|} u(\mathbf{p}, s) \quad (\text{exercise})$$

If our experiment is **blind** to the **electron spin**, we have to **average** over the **incoming electron spin** and **sum** over the **spin of the outgoing electron**

$$\overline{|i\mathcal{M}_{i \rightarrow f}|^2} = \frac{1}{2} \left(\frac{e^2}{m|\mathbf{k}|} \right)^2 |\varepsilon(\mathbf{k}) \cdot \varepsilon'(\mathbf{k}')^*|^2 \sum_{s=\pm\frac{1}{2}} \sum_{s'=\pm\frac{1}{2}} |\bar{u}(\mathbf{p}', s') \not{k} u(\mathbf{p}, s)|^2 \quad (\text{exercise})$$
$$= 4e^4 |\varepsilon(\mathbf{k}) \cdot \varepsilon'(\mathbf{k}')^*|^2$$

For an electron at **rest**, the differential cross section is

$$\frac{d\sigma}{d\Omega} = \frac{3e^4}{48\pi m^2} |\varepsilon(\mathbf{k}) \cdot \varepsilon'(\mathbf{k}')^*|^2$$