

NYU

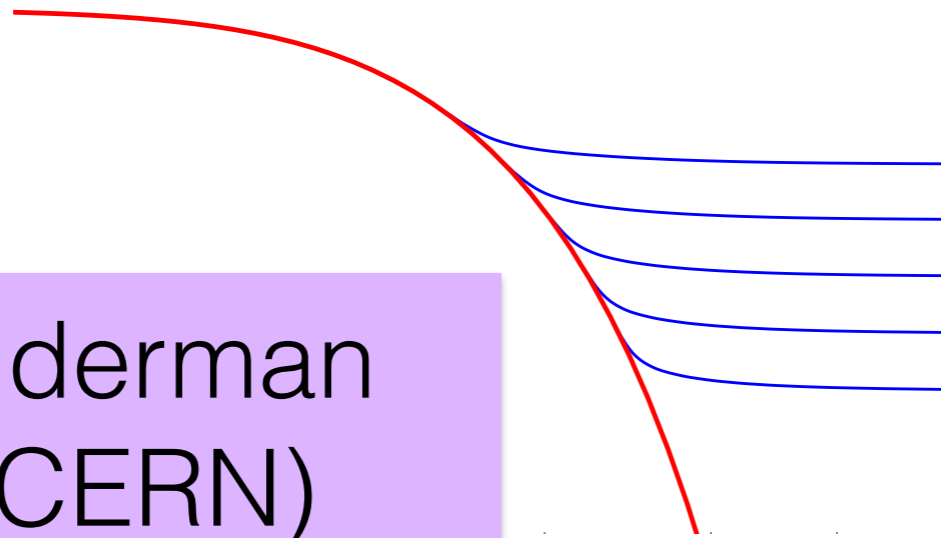


Center for Cosmology  
and Particle Physics



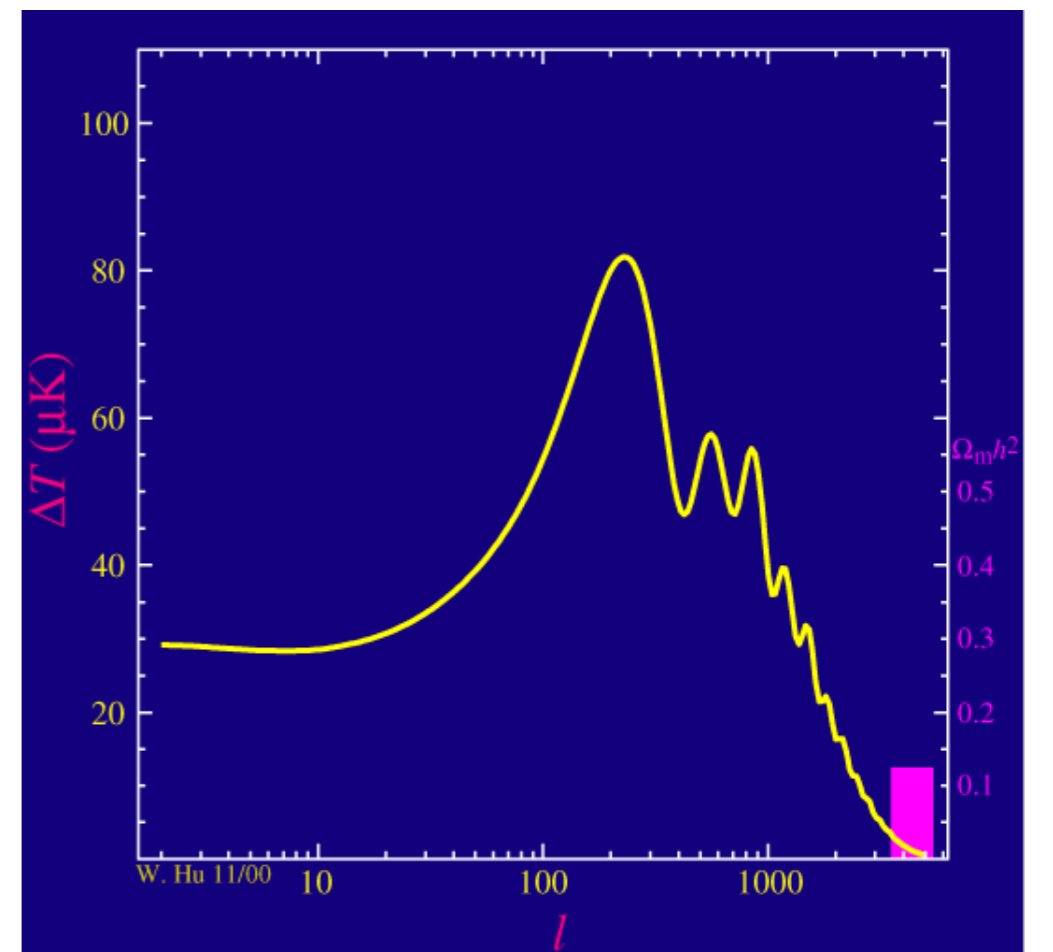
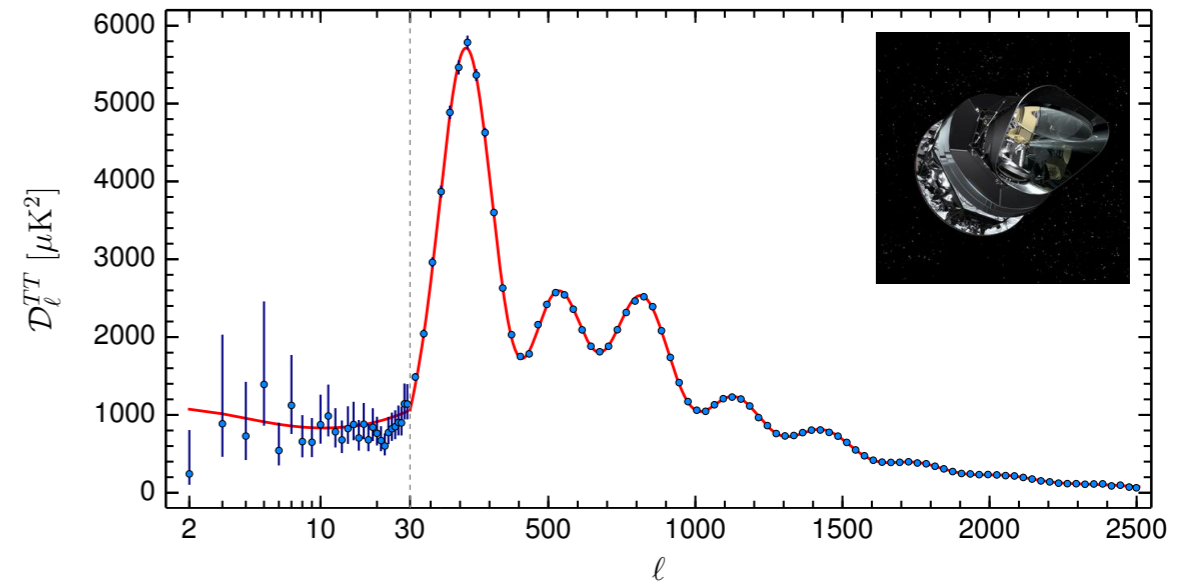
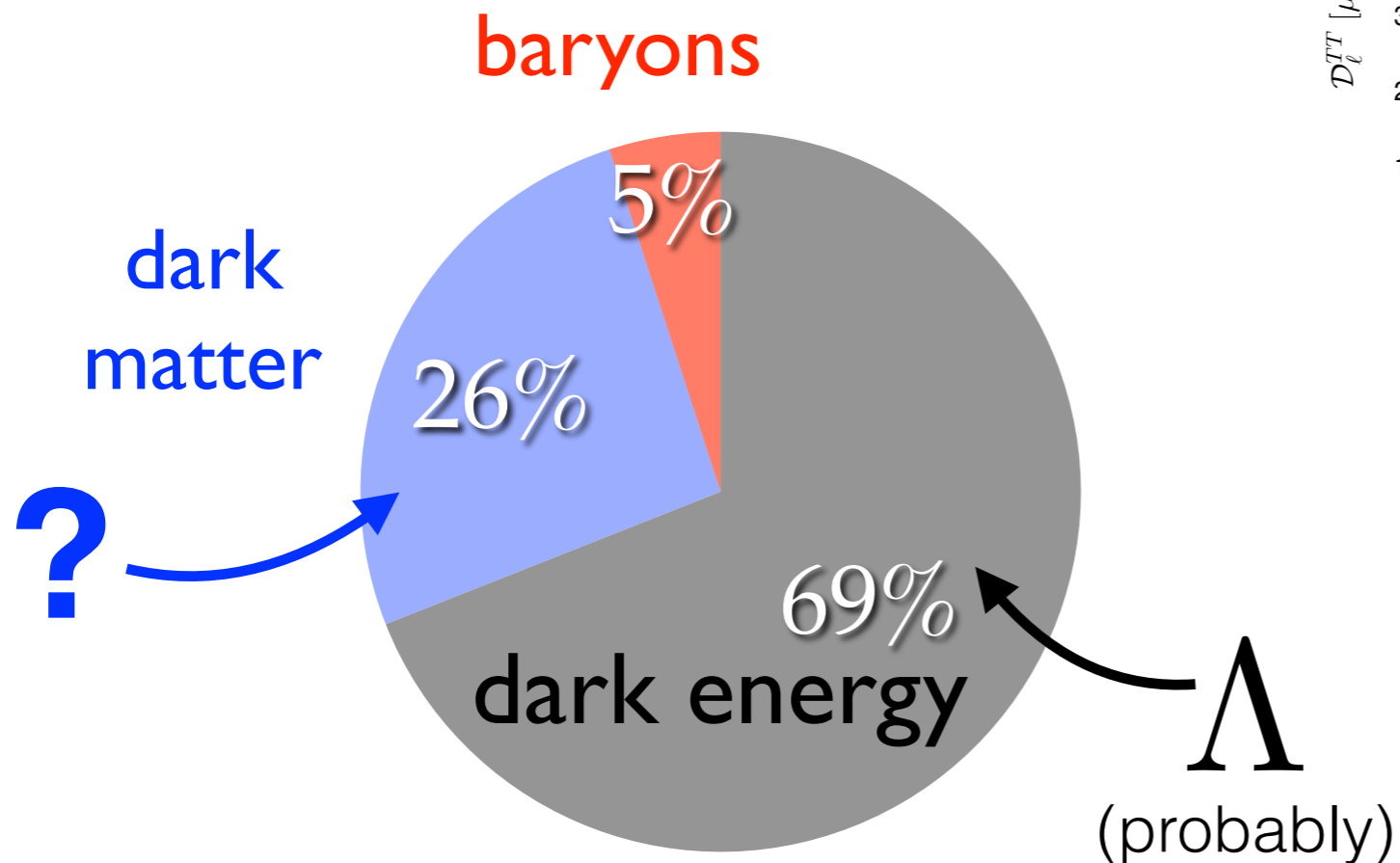
# Thermal Relic Exceptions

Josh Ruderman  
(NYU, CERN)  
@IFAE, 6/8/2018

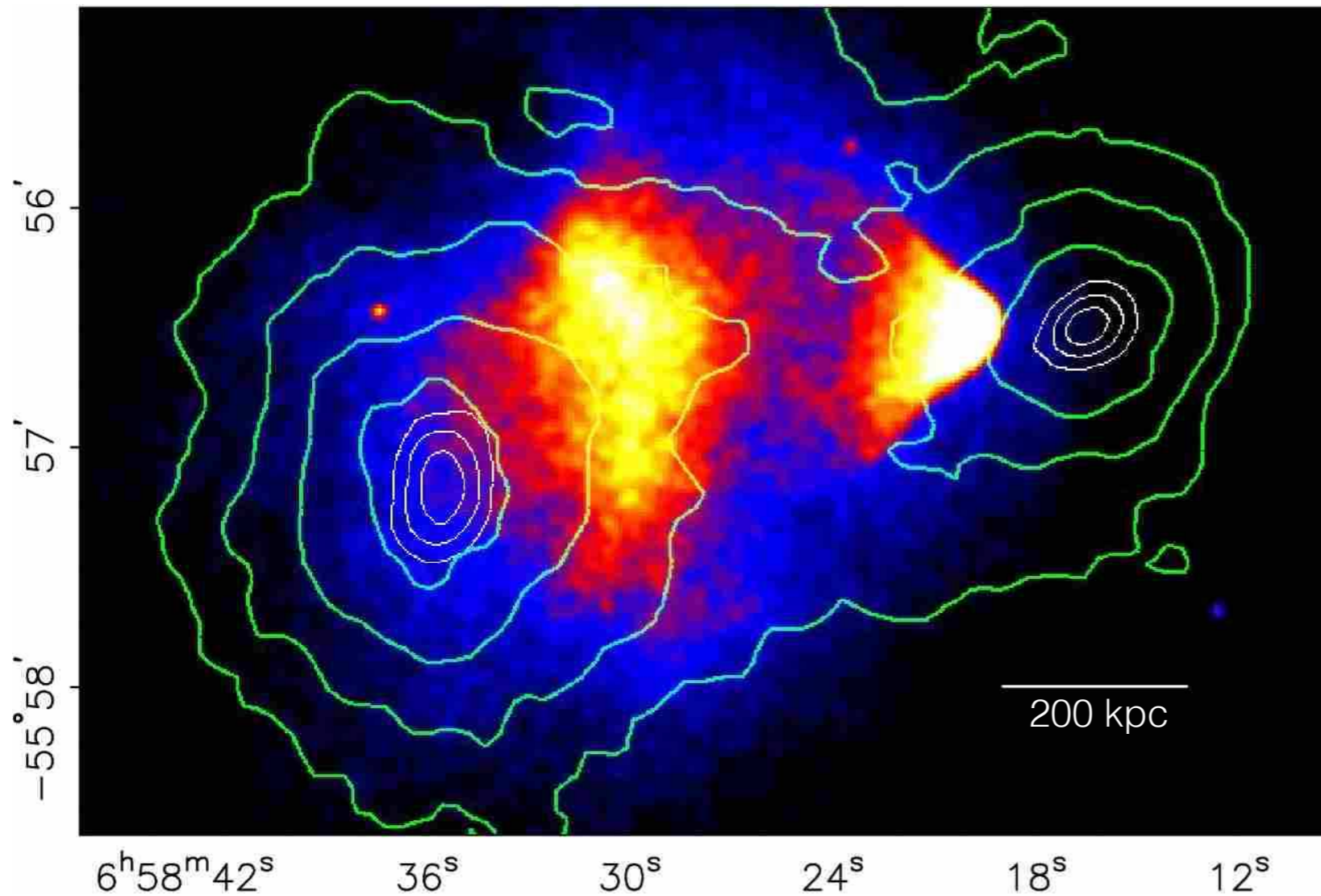


Credit: J. Wise, M. Bradač (Stanford/KIPAC)

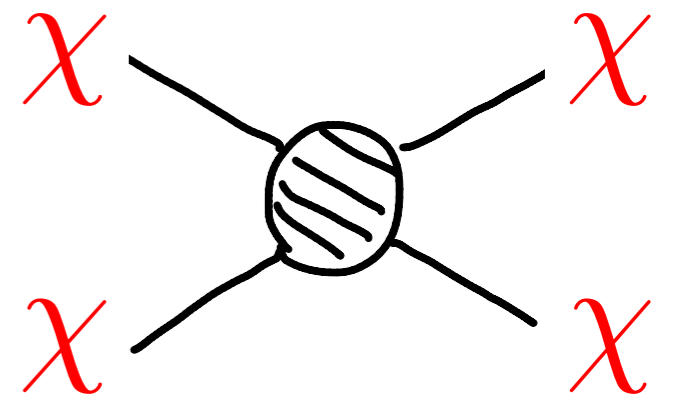
# Energy Budget



# Bullet Cluster



self-interactions:

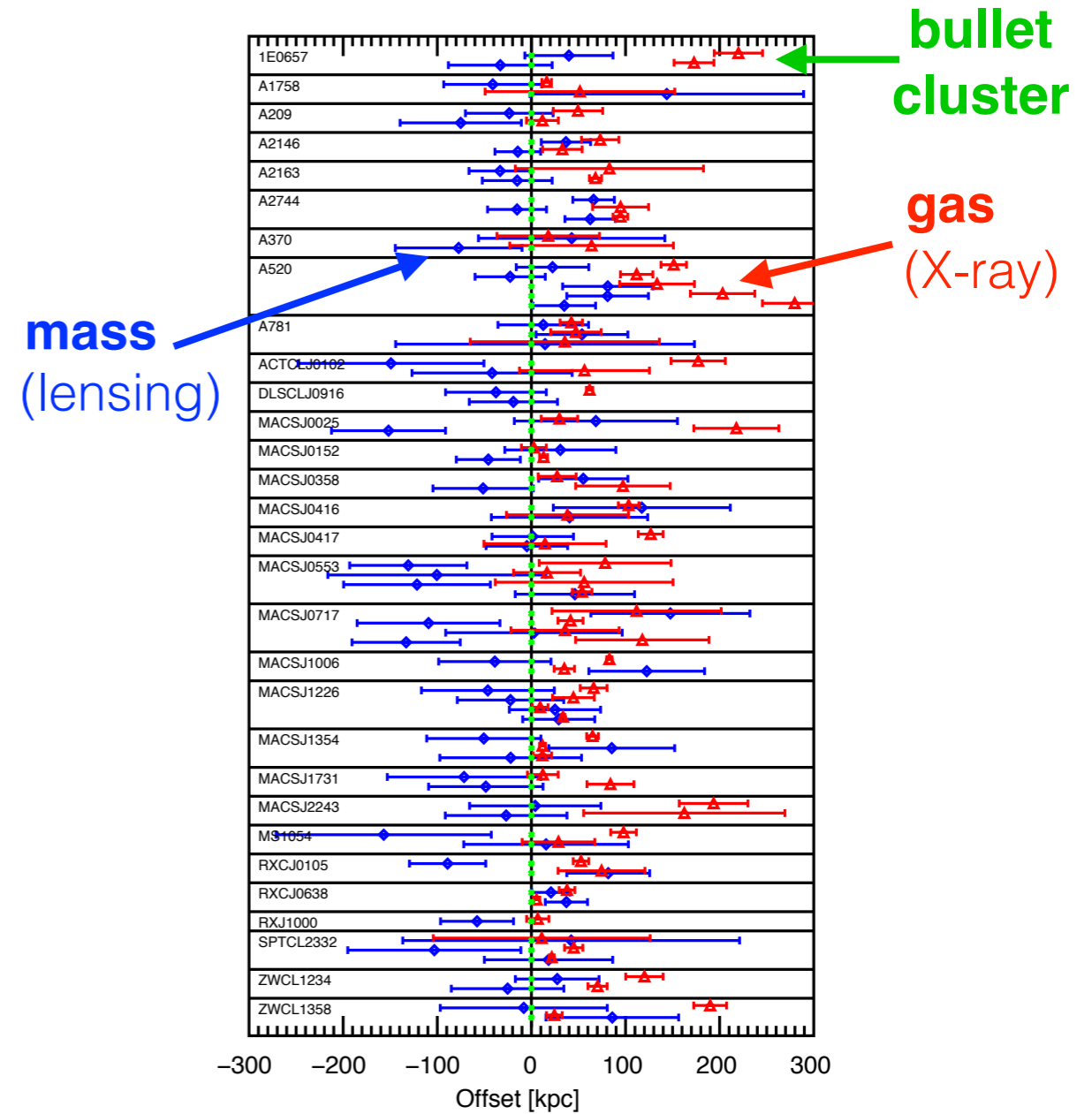
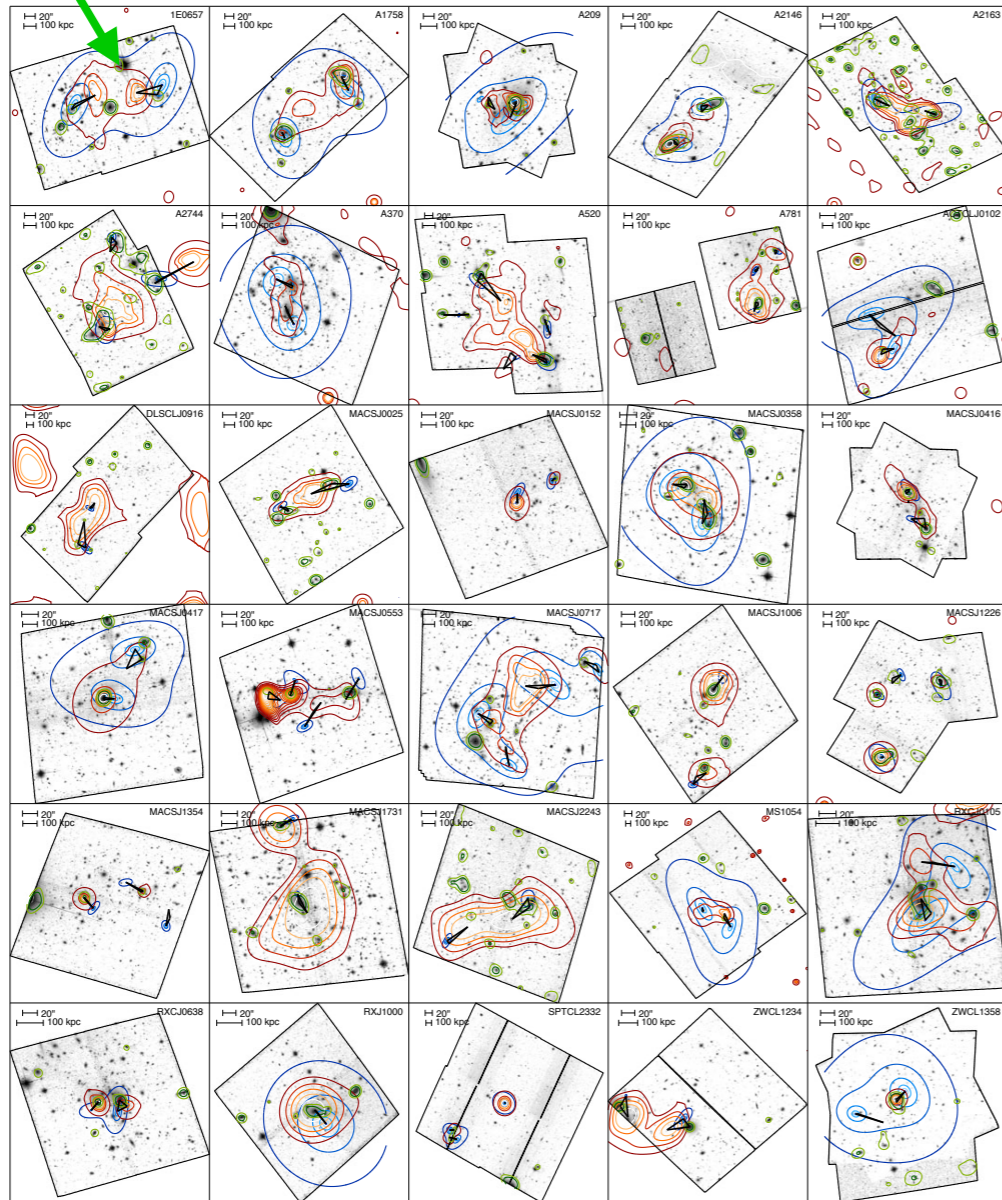


$$\frac{\sigma_{SI}}{m_{\chi}} \lesssim 1 \text{ cm}^2/\text{g}$$
$$\sim 2 \text{ barn}/\text{GeV}$$

- Clowe *et al.*, **astro-ph/0608407**
- Randall *et al.*, **0704.0261**

# Cluster Merger Zoo

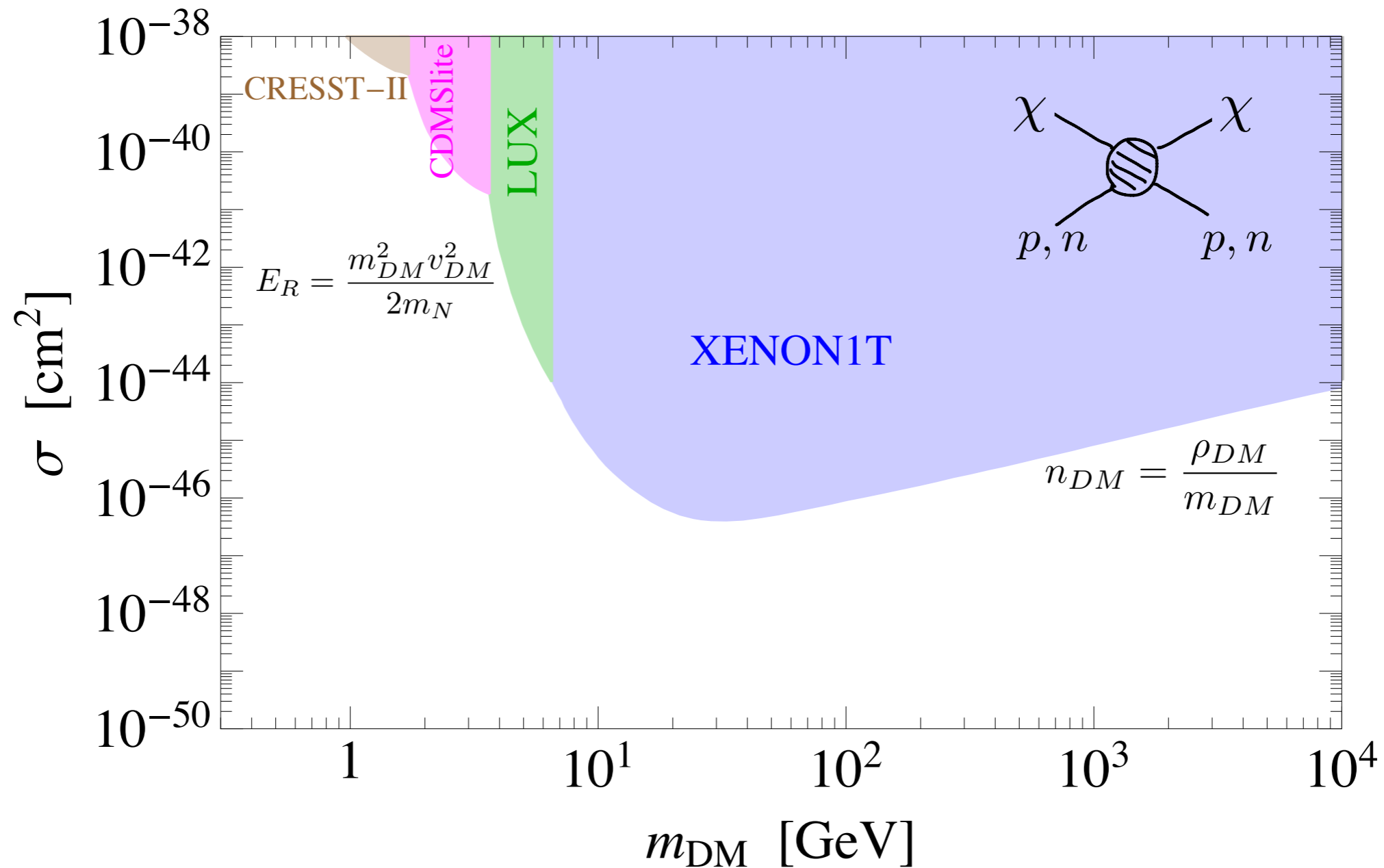
bullet cluster



• Harvey *et al.*, **1503.07675**

$$\frac{\sigma_{SI}}{m_{\chi}} \lesssim 1 \text{ cm}^2/\text{g}$$

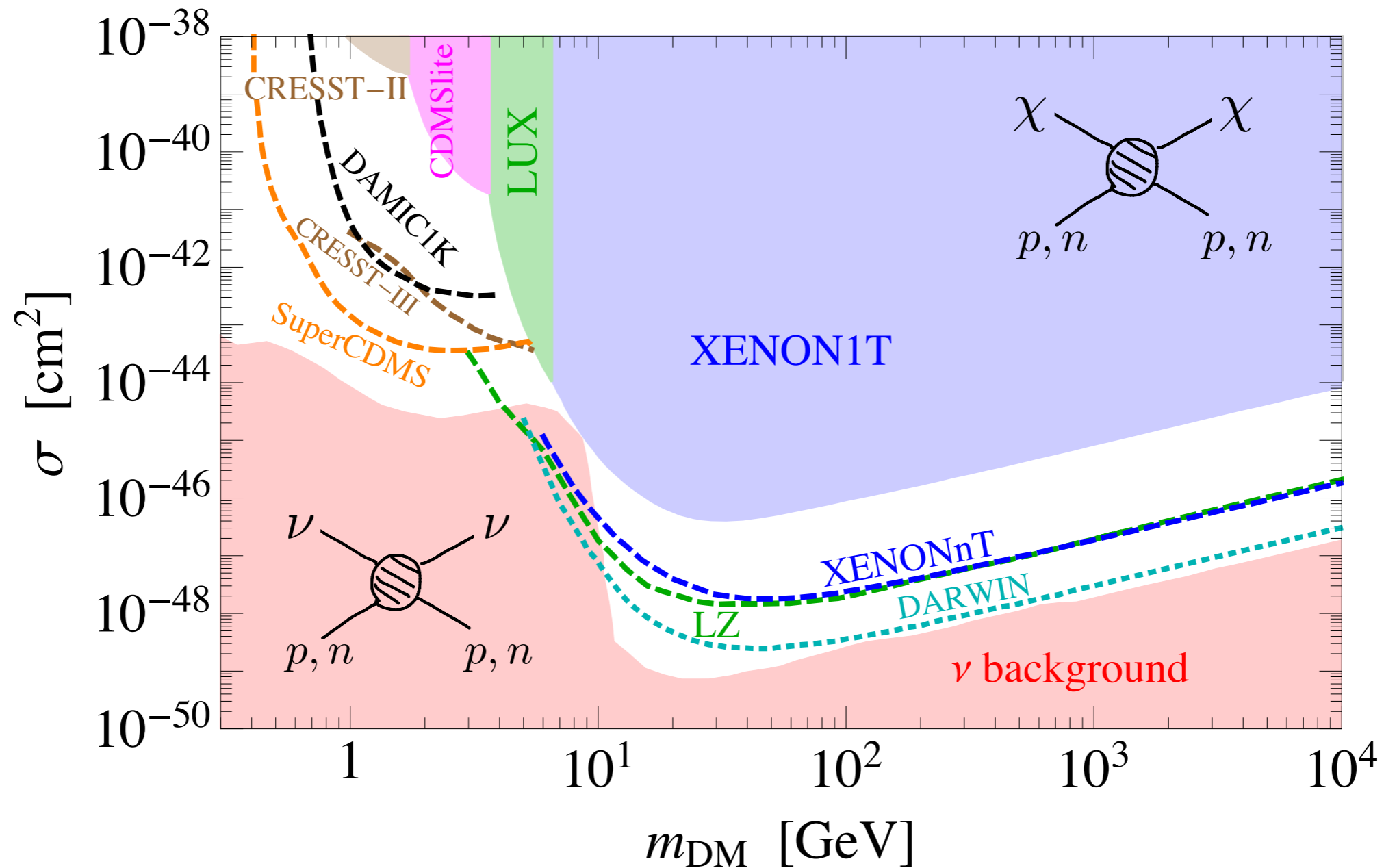
# Direct Detection vs. Dark Matter



- CRESST-II, **1509.01515**
- SuperCDMS, **1509.02448**

- LUX, **1512.03506**
- XENON1T, Monday talks

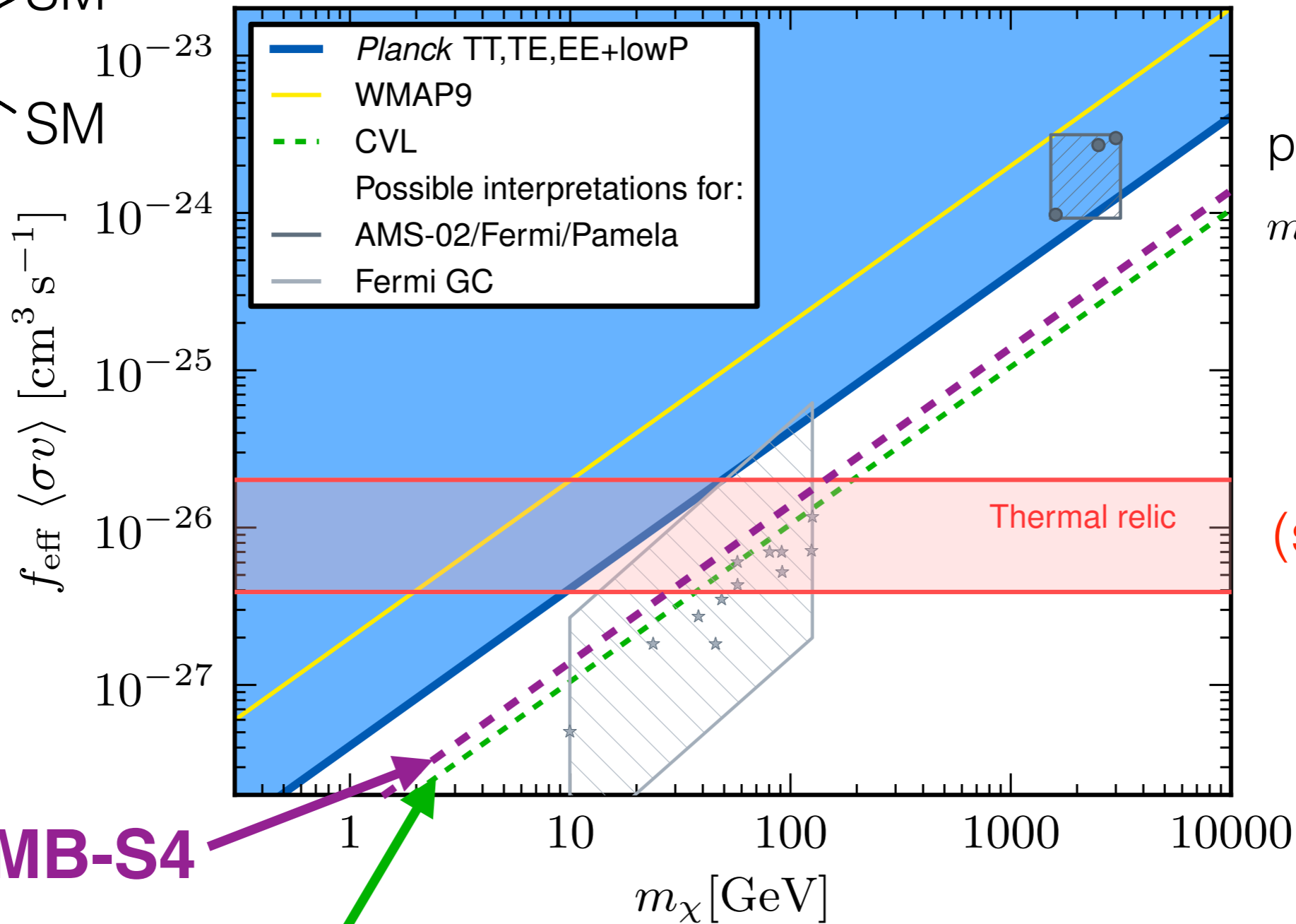
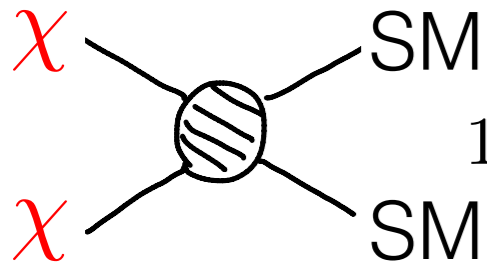
# Direct Detection vs. Dark Matter



- CRESST-III, **1503.08065**
- DAMIC1K, **1707.04591**
- SuperCDMS, **1610.00006**

- XENONnT, **1512.07501**
- LZ, **1703.09144**
- DARWIN, **1606.07001**

# CMB vs. Dark Matter



power:  

$$m_{DM} n_{DM}^2 \propto \frac{1}{m_{DM}}$$

(s-wave)

**CMB-S4**

**cosmic variance**

- Planck, **1502.01589**
- CMS-S4, **1610.02743**

# plan

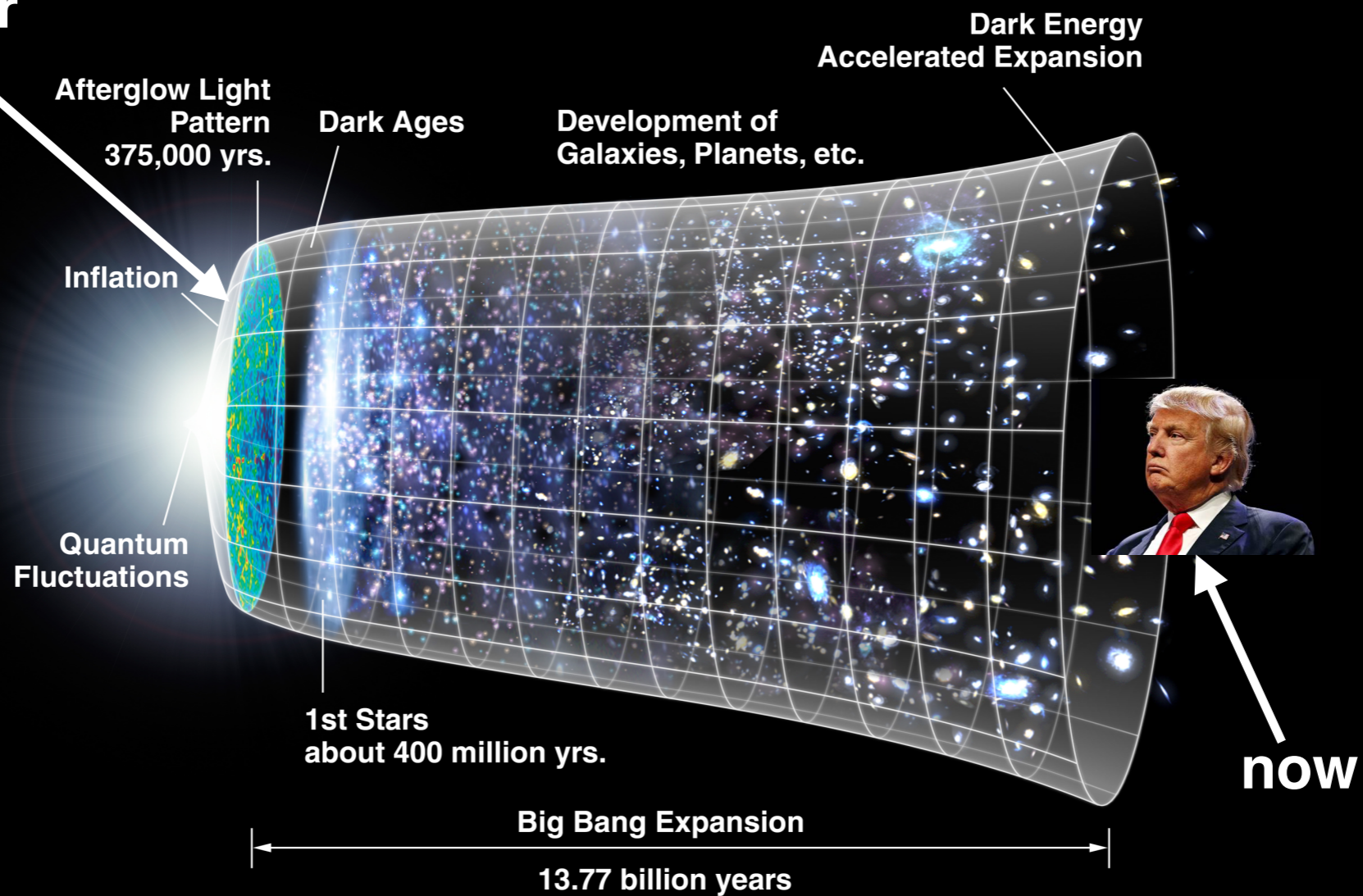
- I. WIMP Warmup
- II. Forbidden DM
- III. Light DM from Coannihilation
- IV. Coscattering





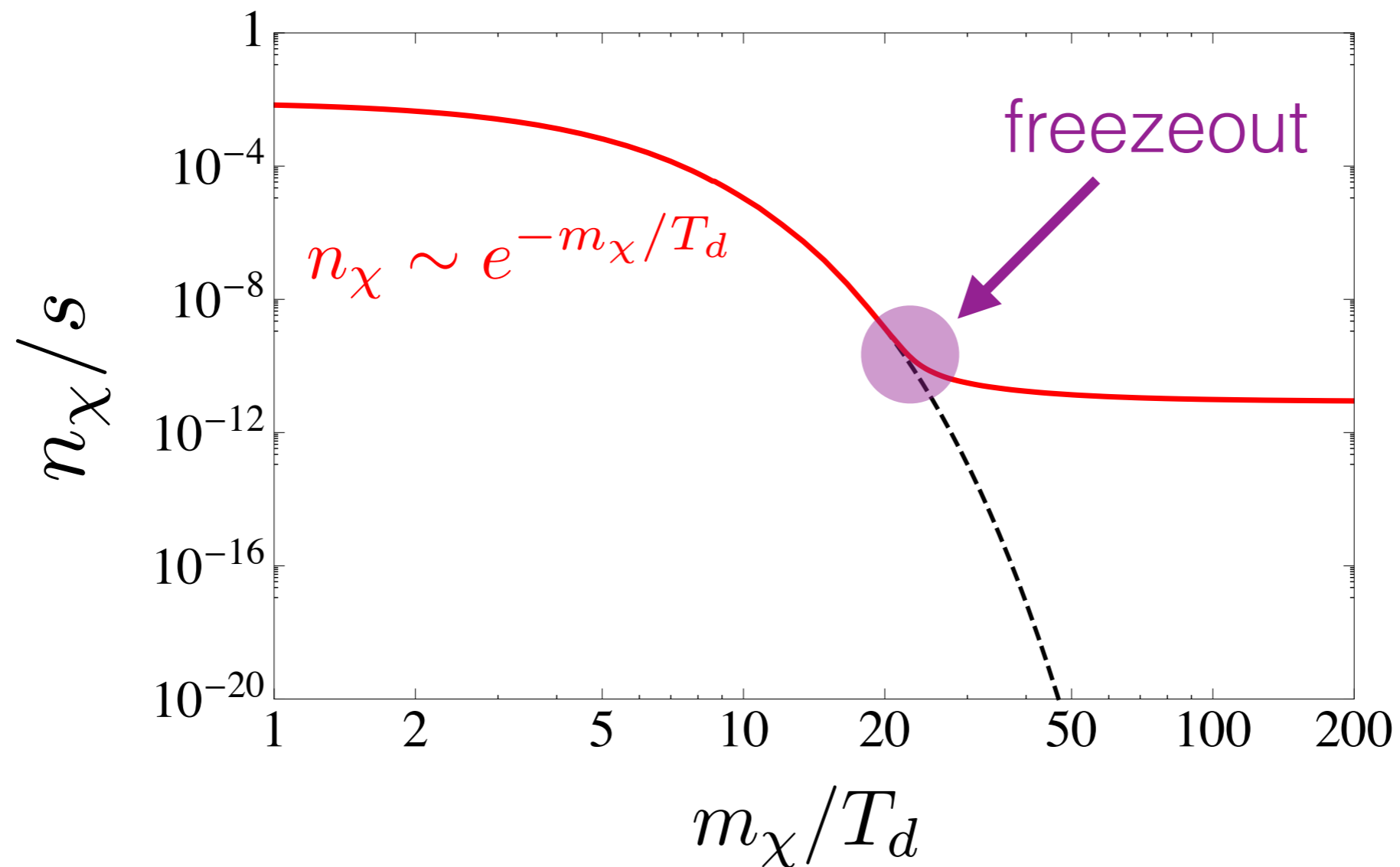
# I. WIMP Warmup

## Dark Matter Genesis

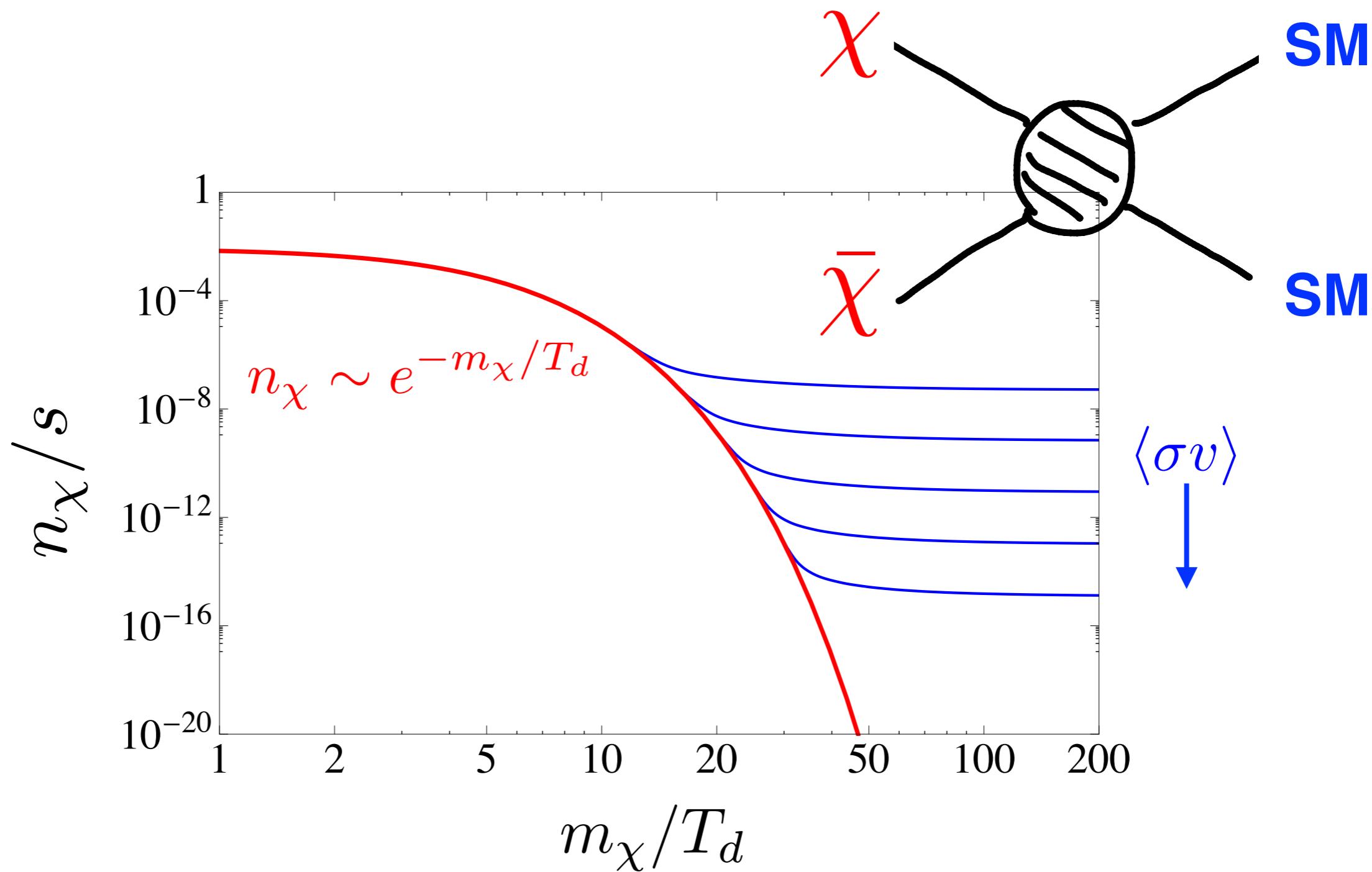


# Thermal Relic Dark Matter

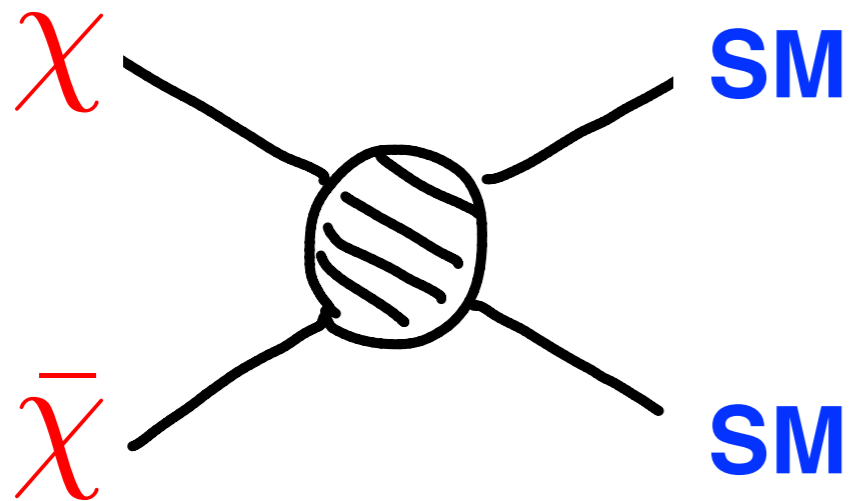
- thermal relic:  $T_d \gtrsim m_\chi$



# WIMP



# WIMP “Miracle”



- freezeout:  $n_\chi \langle \sigma v \rangle \sim H$

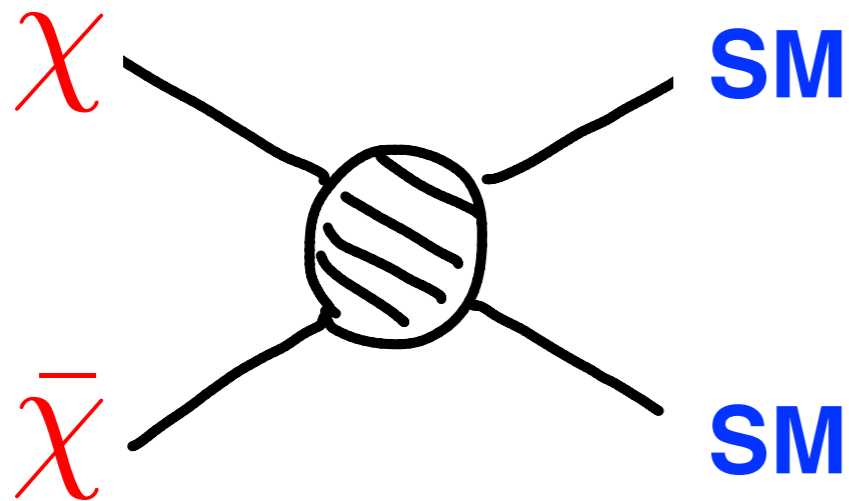
- abundance:

$$\Omega_\chi h^2 \sim \frac{m_\chi n_\chi}{s T_{eq}} \sim \frac{m_\chi H}{\langle \sigma v \rangle s T_{eq}} \sim \frac{1}{\langle \sigma v \rangle T_{eq} M_{pl}}$$

$n_\chi \langle \sigma v \rangle \sim H$  (indicated by a purple arrow pointing up to the  $n_\chi$  term)

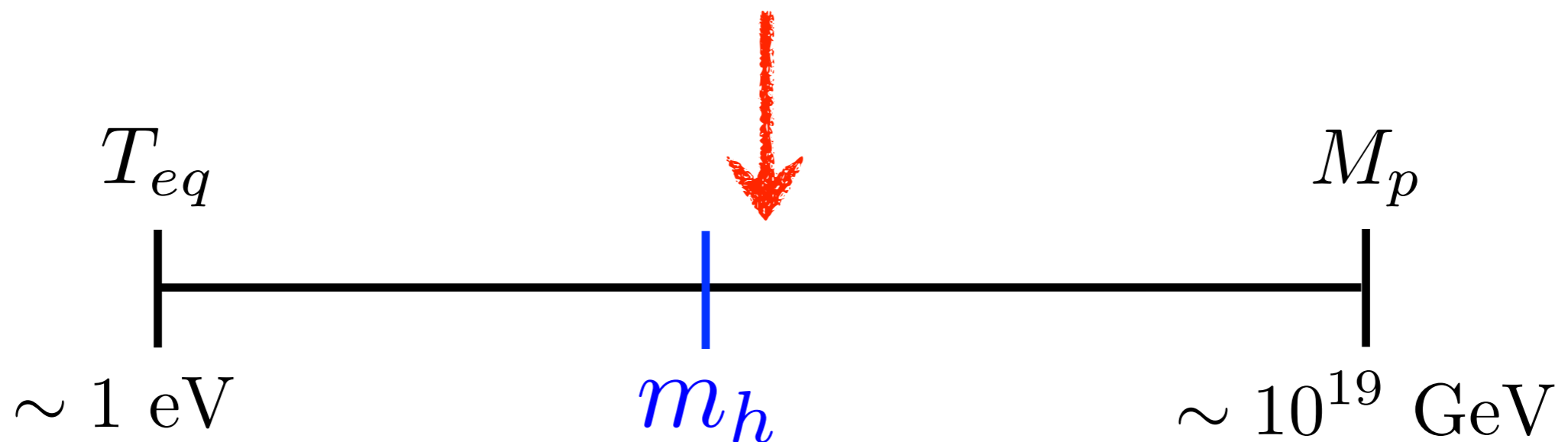
$H \sim T^2 / M_{pl}$   
 $s \sim T^3$   
 $T \sim m$  (indicated by a purple arrow pointing down to the  $T_{eq}$  term)

# WIMP “Miracle”

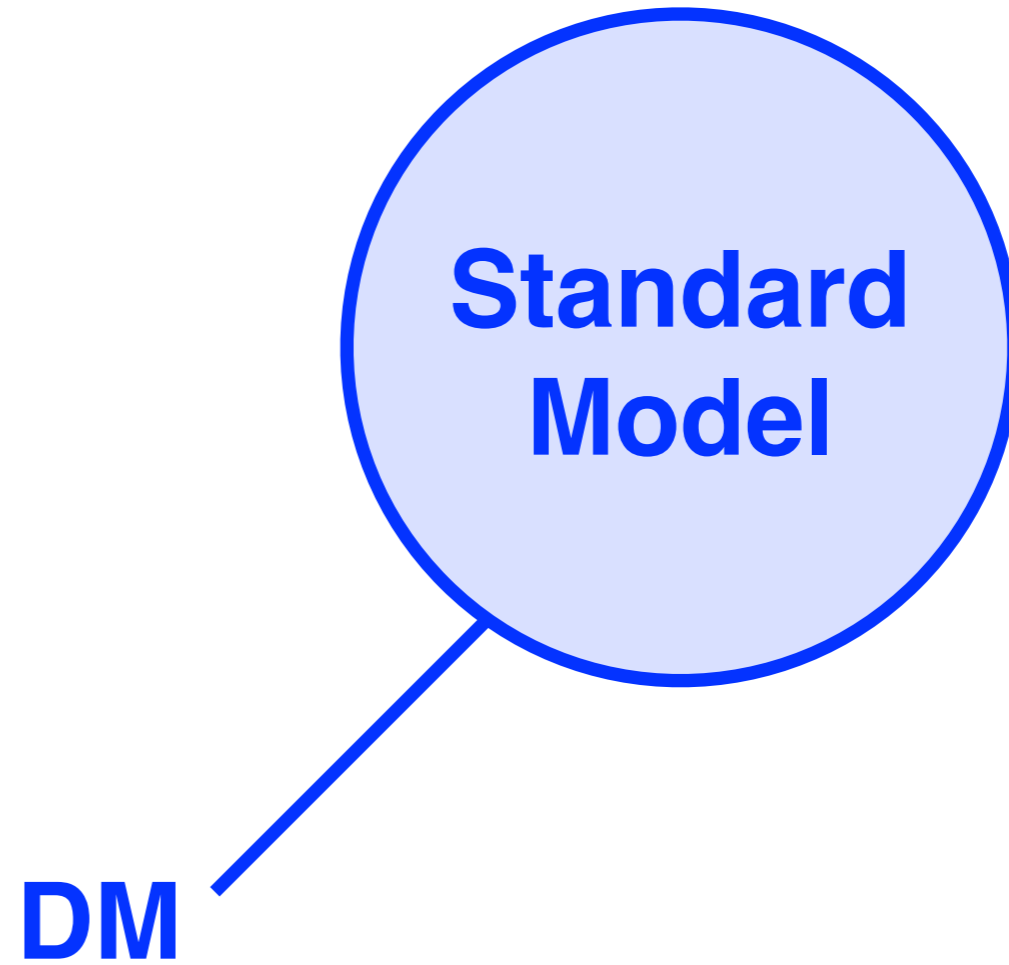


$$\langle \sigma v \rangle \sim \frac{\alpha_W^2}{m_\chi^2} \sim \frac{1}{T_{eq} M_{pl}}$$

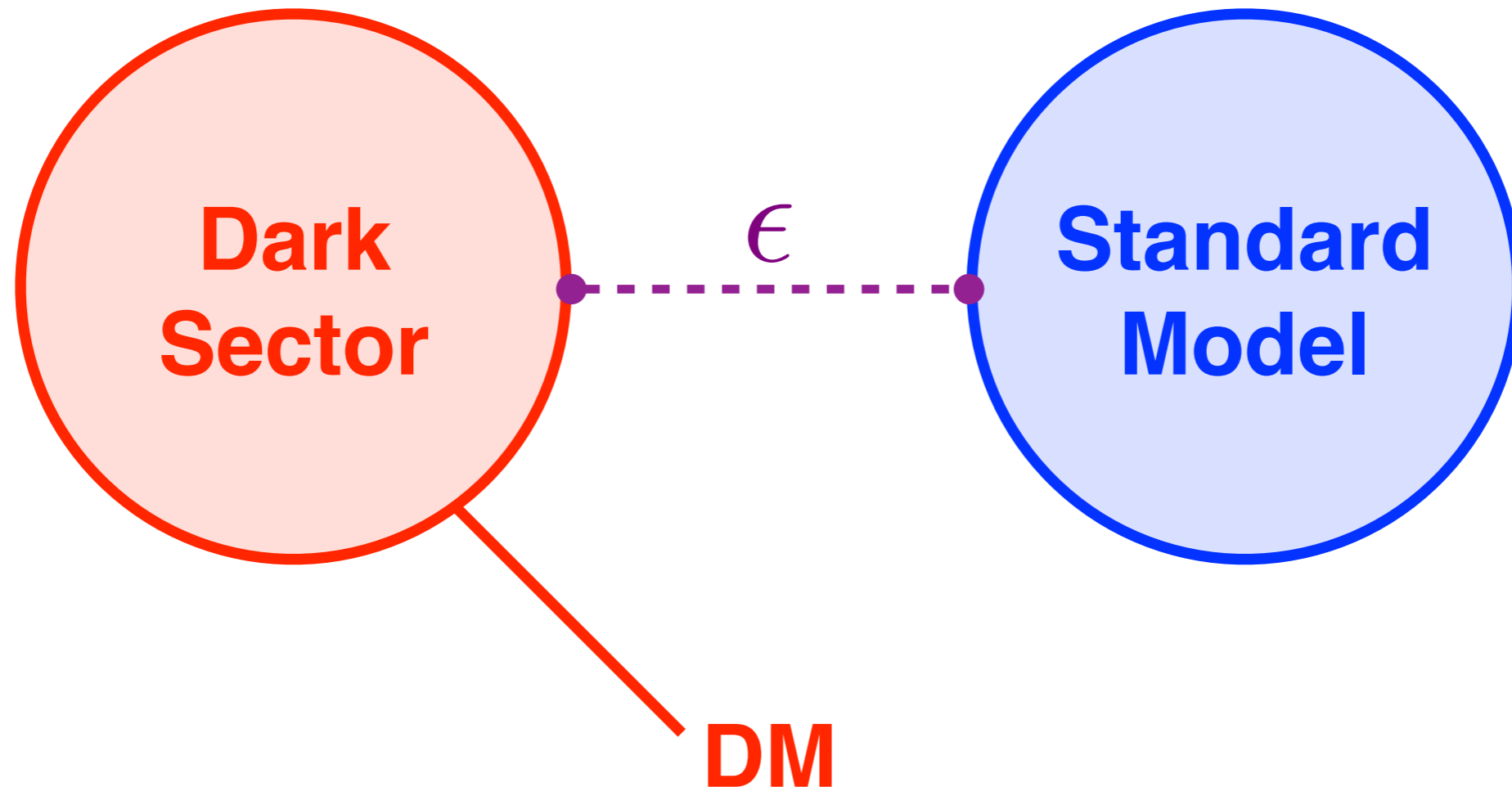
$$m_\chi \sim \alpha_W \sqrt{T_{eq} M_{pl}}$$



# WIMP



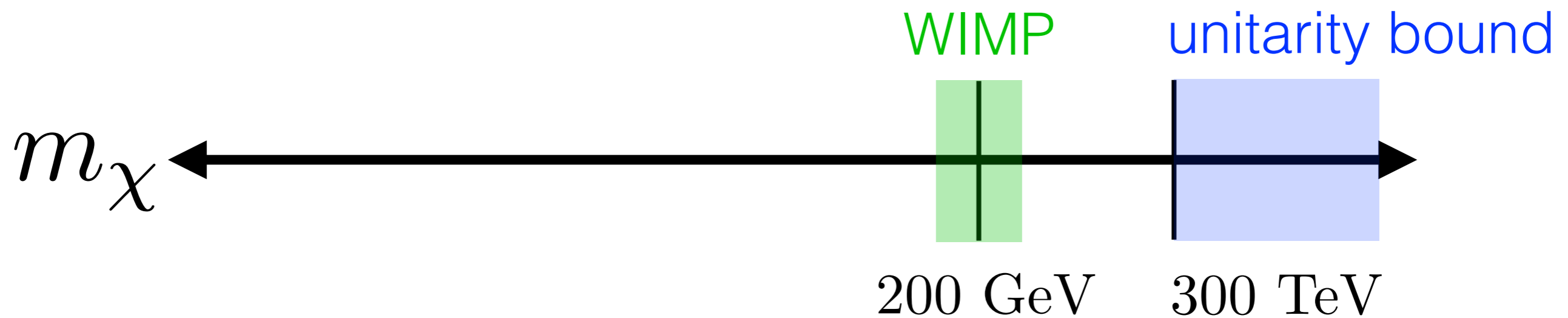
# Hidden Sector Dark Matter



What are the possible cosmologies for thermal relics in hidden sectors?

- Goldberg, Hall, **1986**
- Strassler, Zurek, **hep-ph/0604261**
- Arkani-Hamed, Finkbeiner, Slatyer, Weiner, **0810.0713**

# Thermal Relic Spectrum



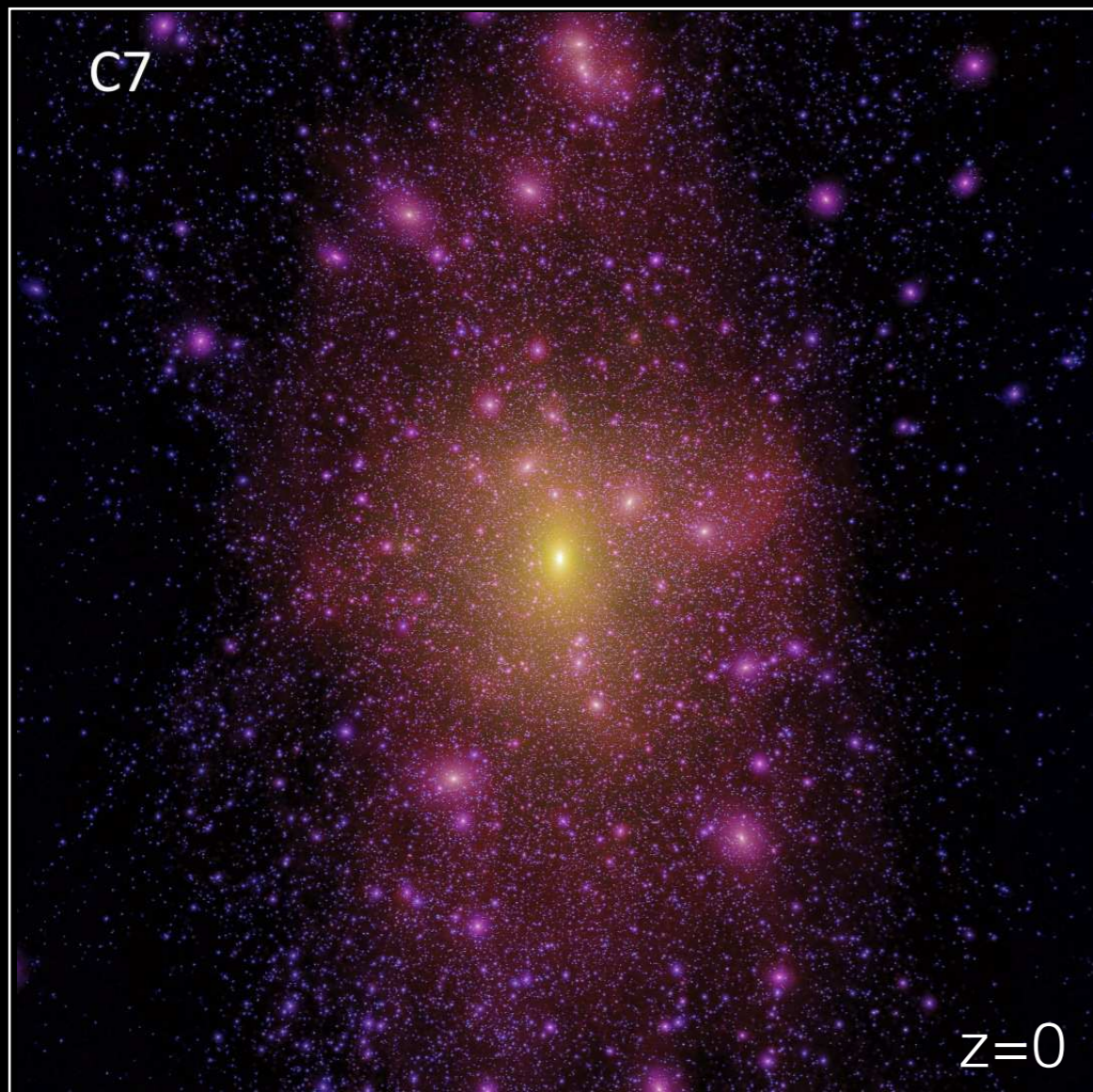
- Griest, Kamionkowski, Phys. Rev. Lett. **64**, 615 (1990).



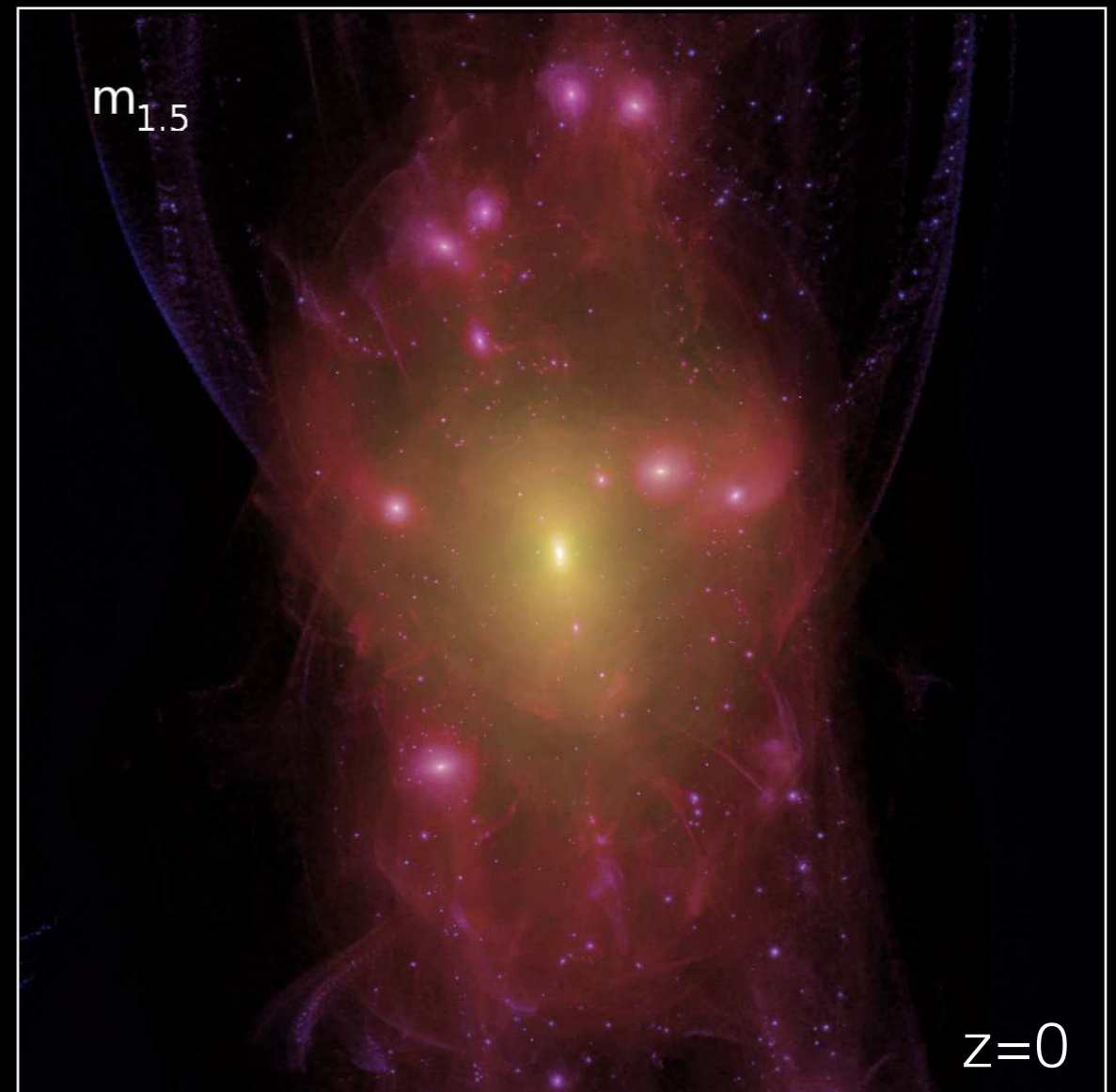
# Cold vs. Warm Dark Matter

**Cold:**  $m_{DM} \gg 1 \text{ keV}$

**Warm:**  $m_{DM} = 1.5 \text{ keV}$



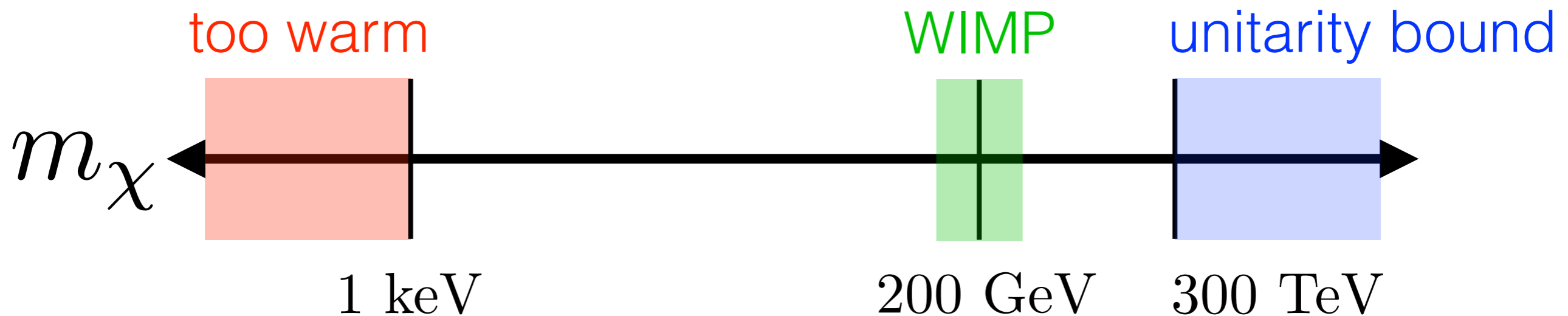
1.5 Mpc



1.5 Mpc

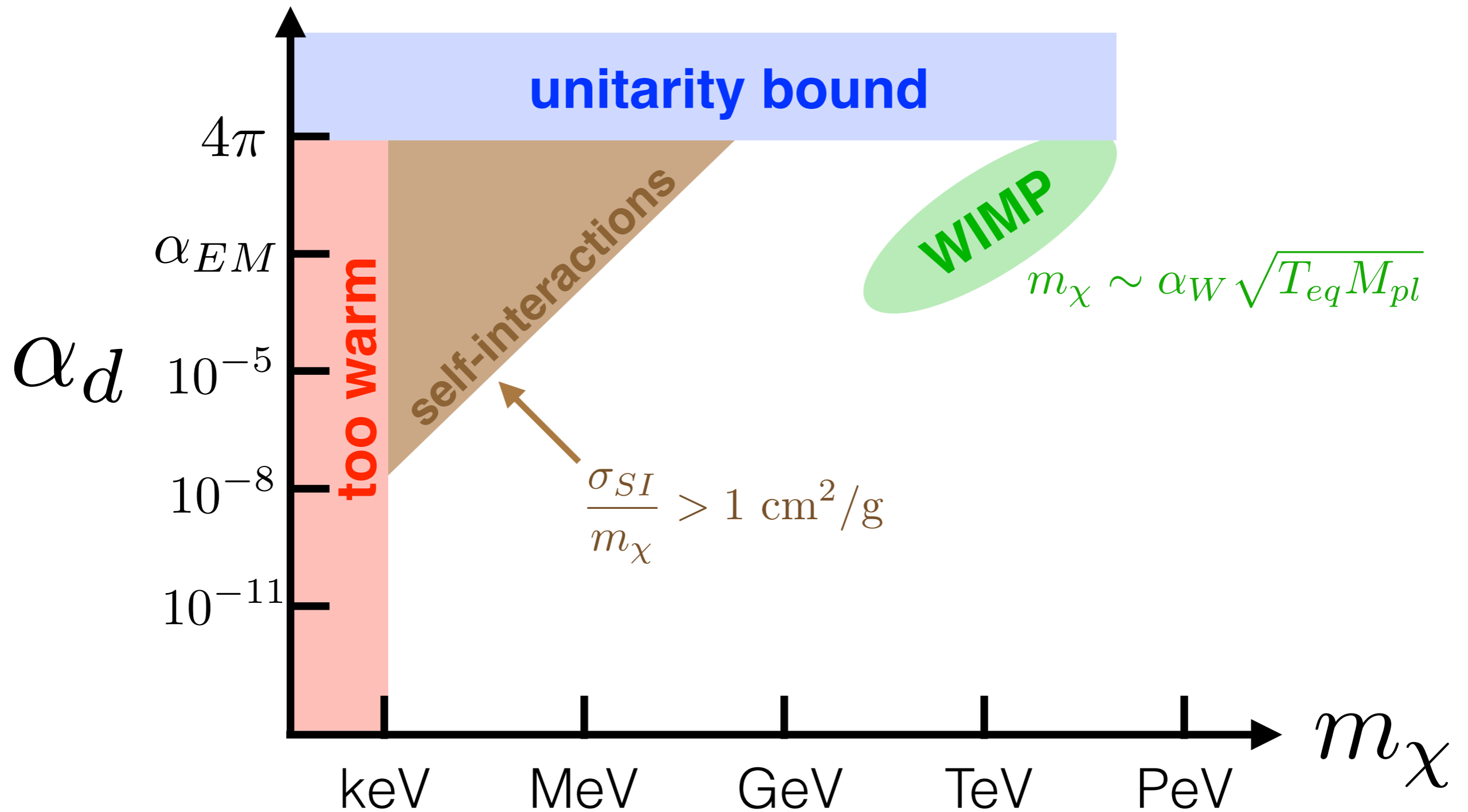
- Lovell *et al.*, **1308.1399**

# Thermal Relic Spectrum

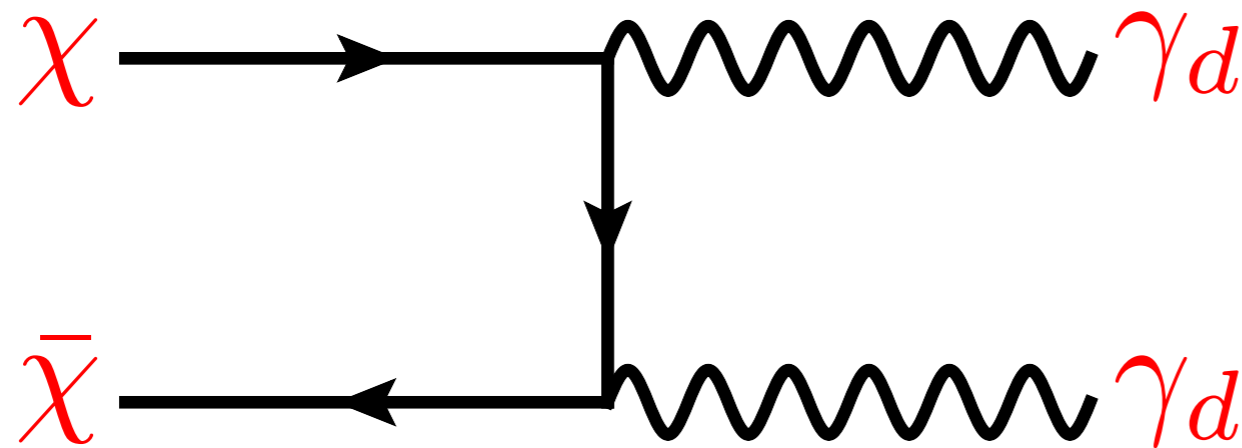


- Lovell *et al.*, **1308.1399**
- Irsic *et al.*, **1702.01764**
- Griest, Kamionkowski, Phys. Rev. Lett. **64**, 615 (1990).

# Dark Freezeout Map



# Ultraweakly Coupled Dark Matter

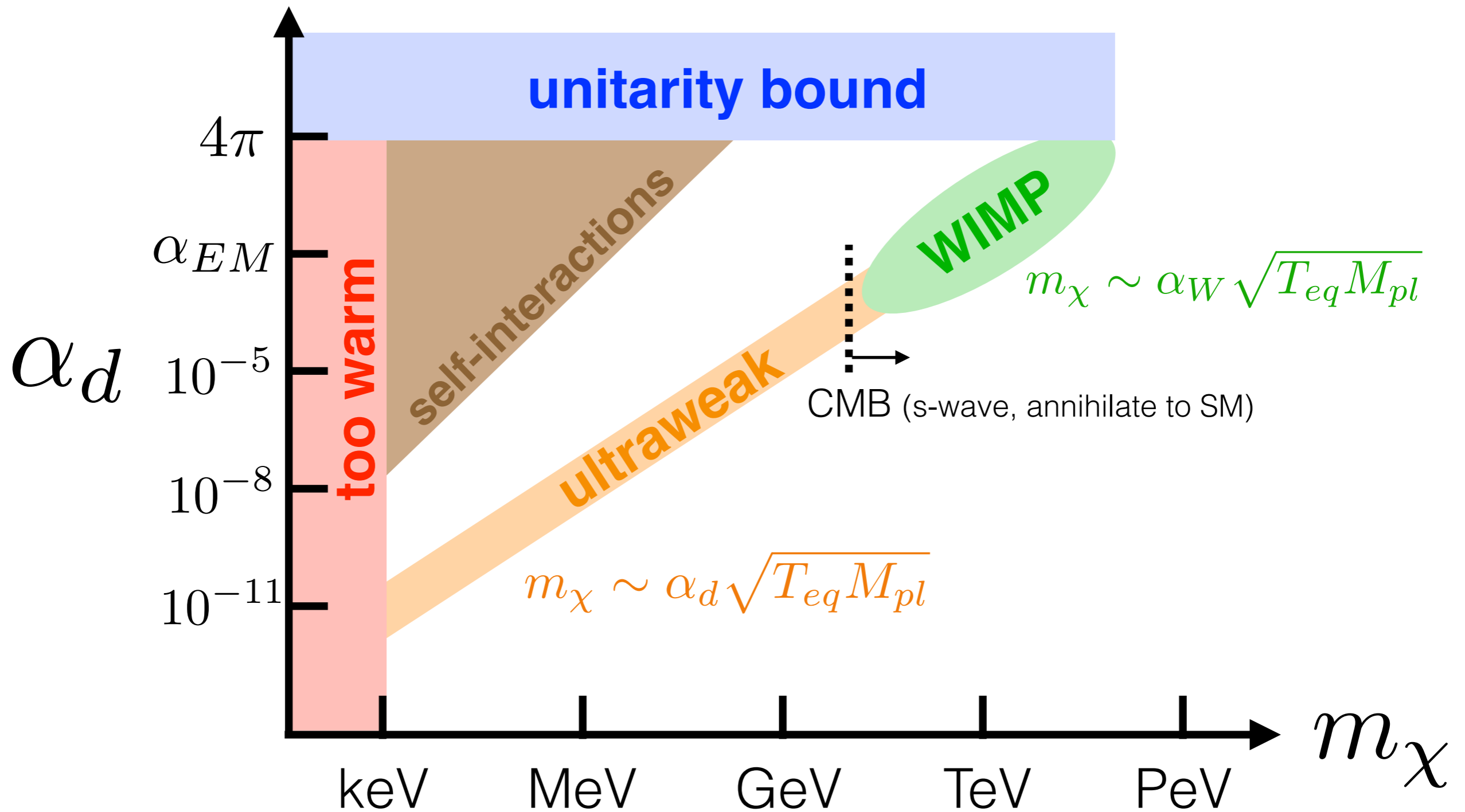


$$\langle \sigma v \rangle \sim \frac{\alpha_d^2}{m_\chi^2}$$

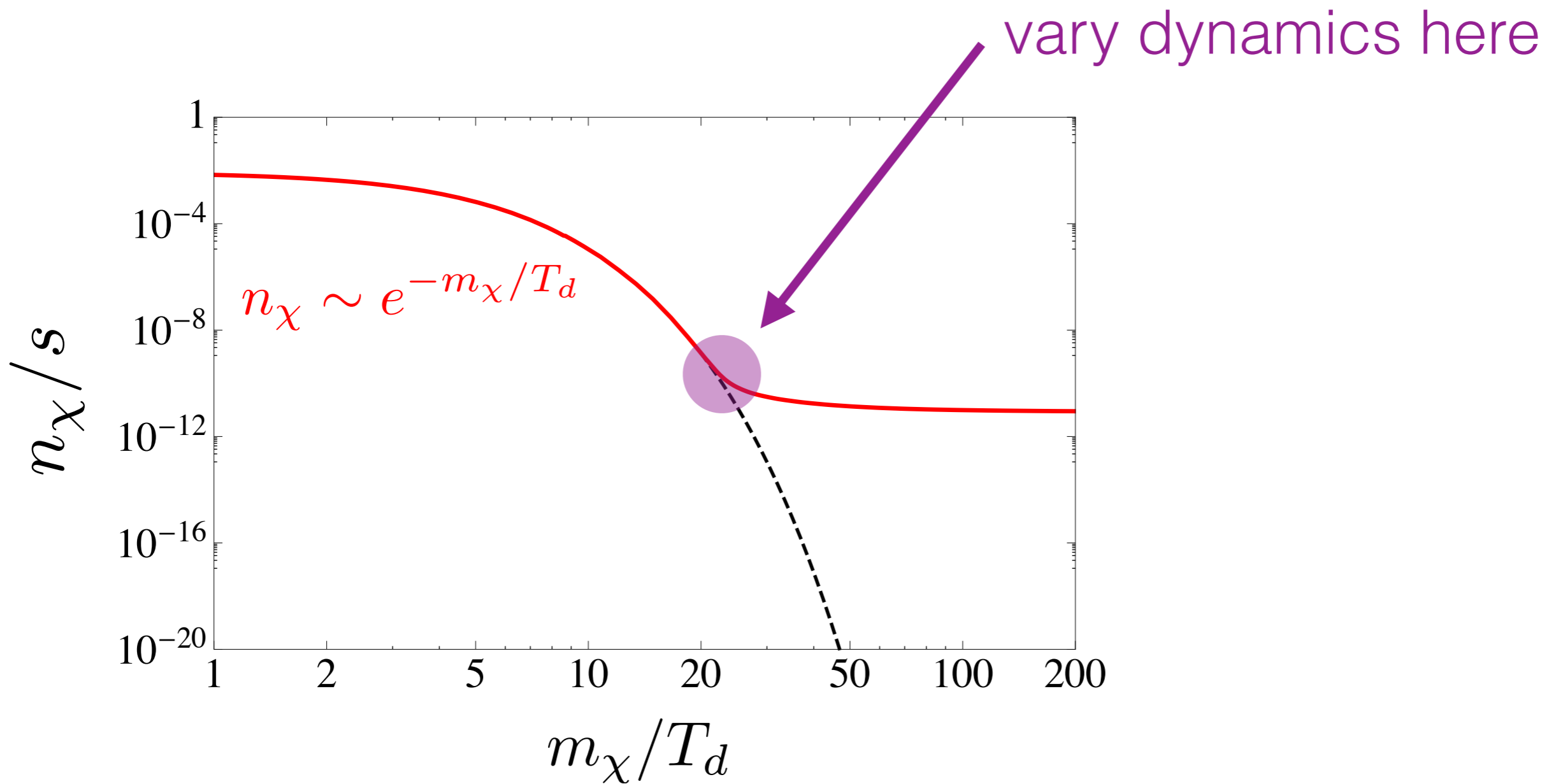
$$m_\chi \sim \alpha_d \sqrt{T_{eq} M_{pl}} \sim \left( \frac{\alpha_d}{\alpha_W} \right) \times 1 \text{ TeV}$$

- Boehm, Fayet, **hep-ph/0305261**
- Finkbeiner, Weiner, **astro-ph/0702587**
- Pospelov, Ritz, Voloshin, **0711.4866**
- Feng, Kumar **0803.4196**

# Dark Freezeout Map



# Thermal Relics Beyond the WIMP



# WIMP-like Assumptions

- thermal relic with:

**A)**  $2 \leftrightarrow 2$  annihilations\*

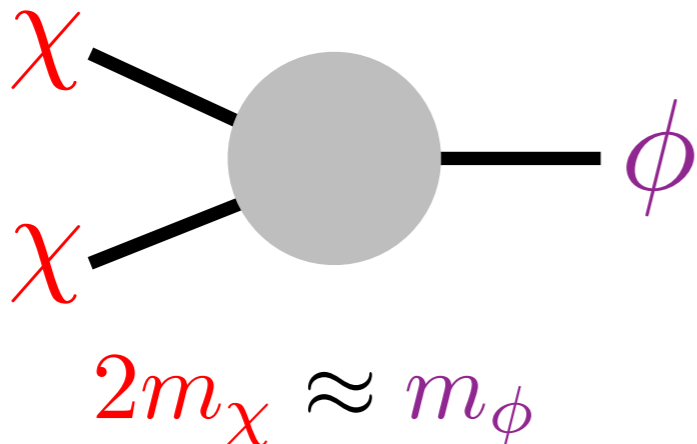
**B)**  $T_d \sim T_\gamma$

**C)** no entropy dump

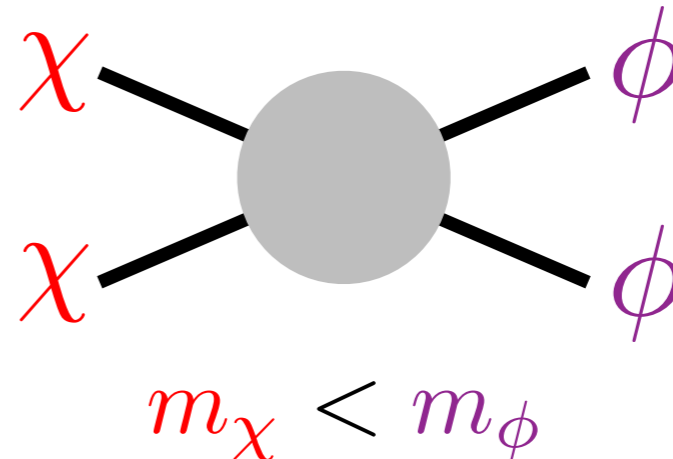
\* $Z_2$ , symmetric

# Dark Exceptions

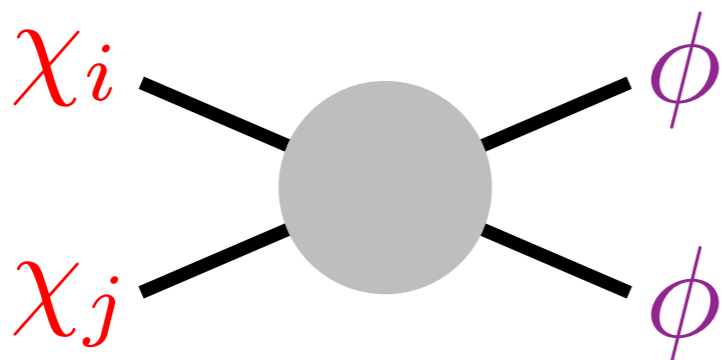
1) pole



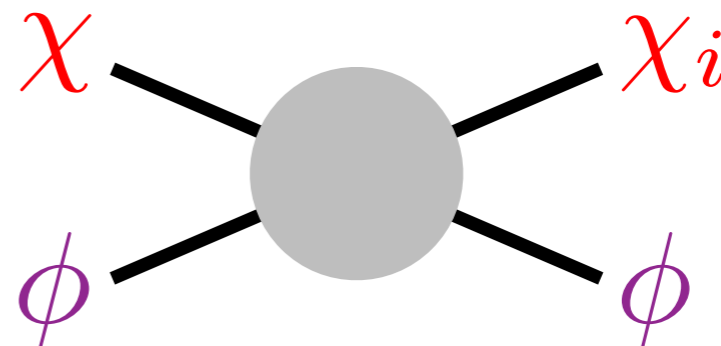
2) forbidden channels



3) coannihilations



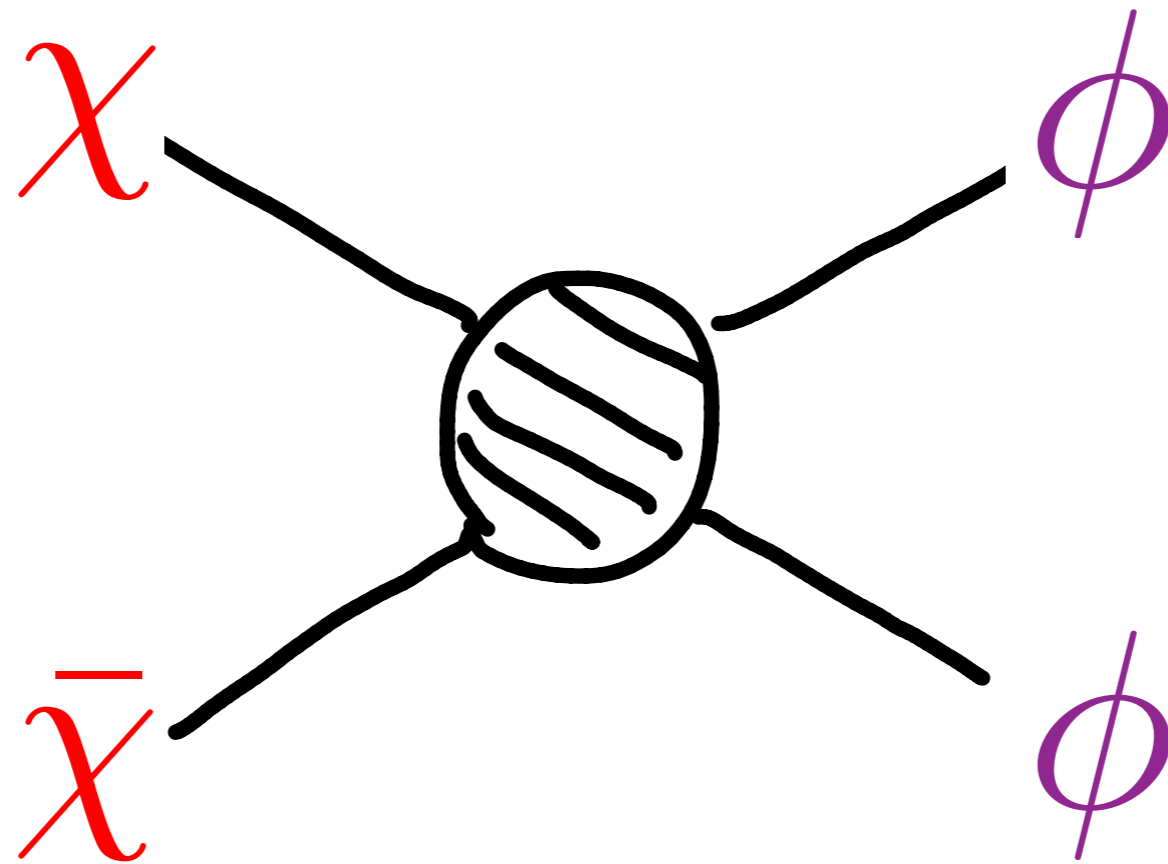
4) cospattering



- Griest, Seckel, **1991**
- D'Agnolo, Pappadopulo, JTR, **1705.08450**



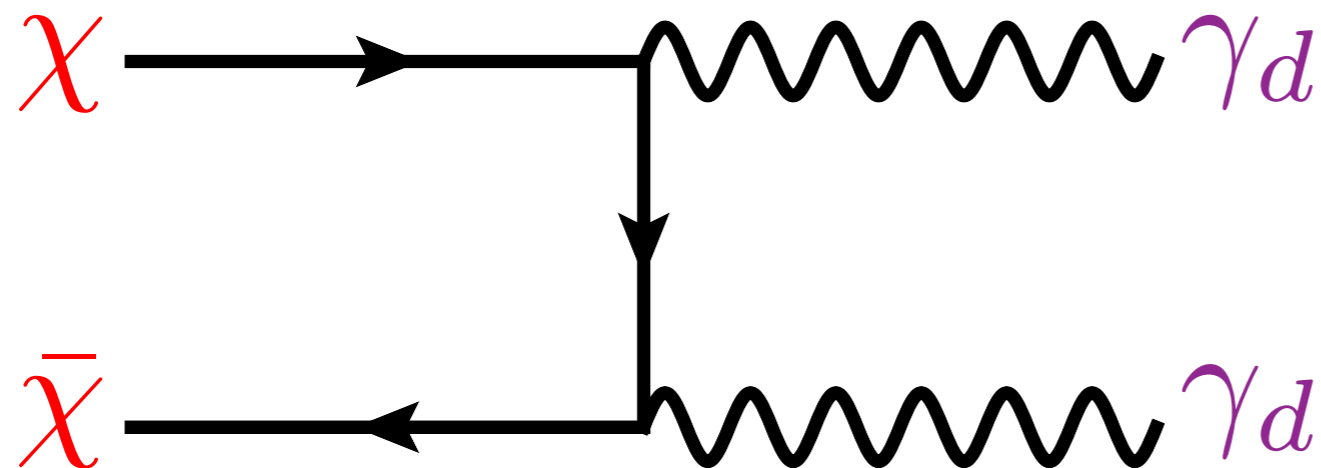
# II. Forbidden Dark Matter



$$m_{\chi} < m_{\phi}$$

- Raffaele D'Agnolo, JTR, **1505.07107** (PRL)

# Forbidden Dark Matter



$$\langle \sigma v \rangle_{\chi\bar{\chi}}$$

- detailed balance:

$$(n_{\chi}^{eq})^2 \times \begin{array}{c} \chi \longrightarrow \\ \bar{\chi} \longleftarrow \end{array} \begin{array}{c} \text{---} \gamma_d \\ \text{---} \gamma_d \end{array} = (n_{\gamma_d}^{eq})^2 \times \begin{array}{c} \gamma_d \text{---} \\ \gamma_d \text{---} \end{array} \begin{array}{c} \longleftarrow \chi \\ \longrightarrow \bar{\chi} \end{array} \sim \frac{\alpha_d^2}{m_{\chi}^2}$$

- thermally averaged cross:

$$\langle \sigma v \rangle_{\chi\bar{\chi}} = \frac{(n_{\gamma_d}^{eq})^2}{(n_{\chi}^{eq})^2} \langle \sigma v \rangle_{\gamma_d\gamma_d} \sim \frac{\alpha_d^2}{m_{\chi}^2} e^{-2\Delta m/T_{FO}} \sim \frac{1}{T_{eq} M_{pl}}$$

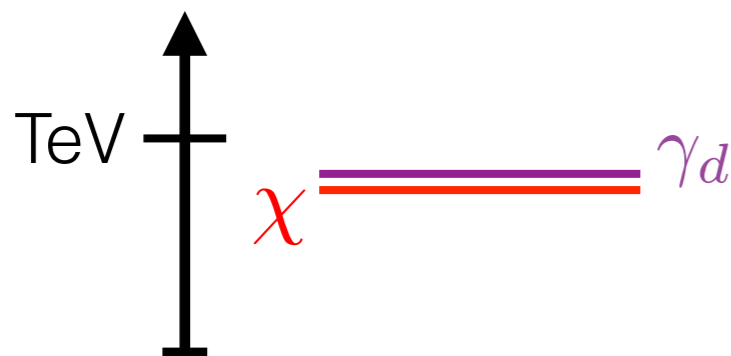
# Forbidden Mass Spectrum

$$m_\chi \sim e^{-\Delta m/T_{FO}} \alpha_d \sqrt{T_{eq} M_{pl}}$$

**Griest, Seckel, 1991**

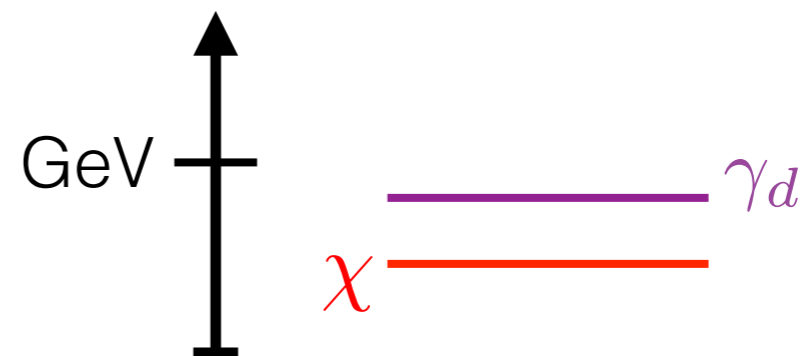
- assume:  $m_\chi \sim \text{TeV}$
- forbidden relevant when:

$$\frac{\Delta m}{m_\chi} \lesssim 0.1$$



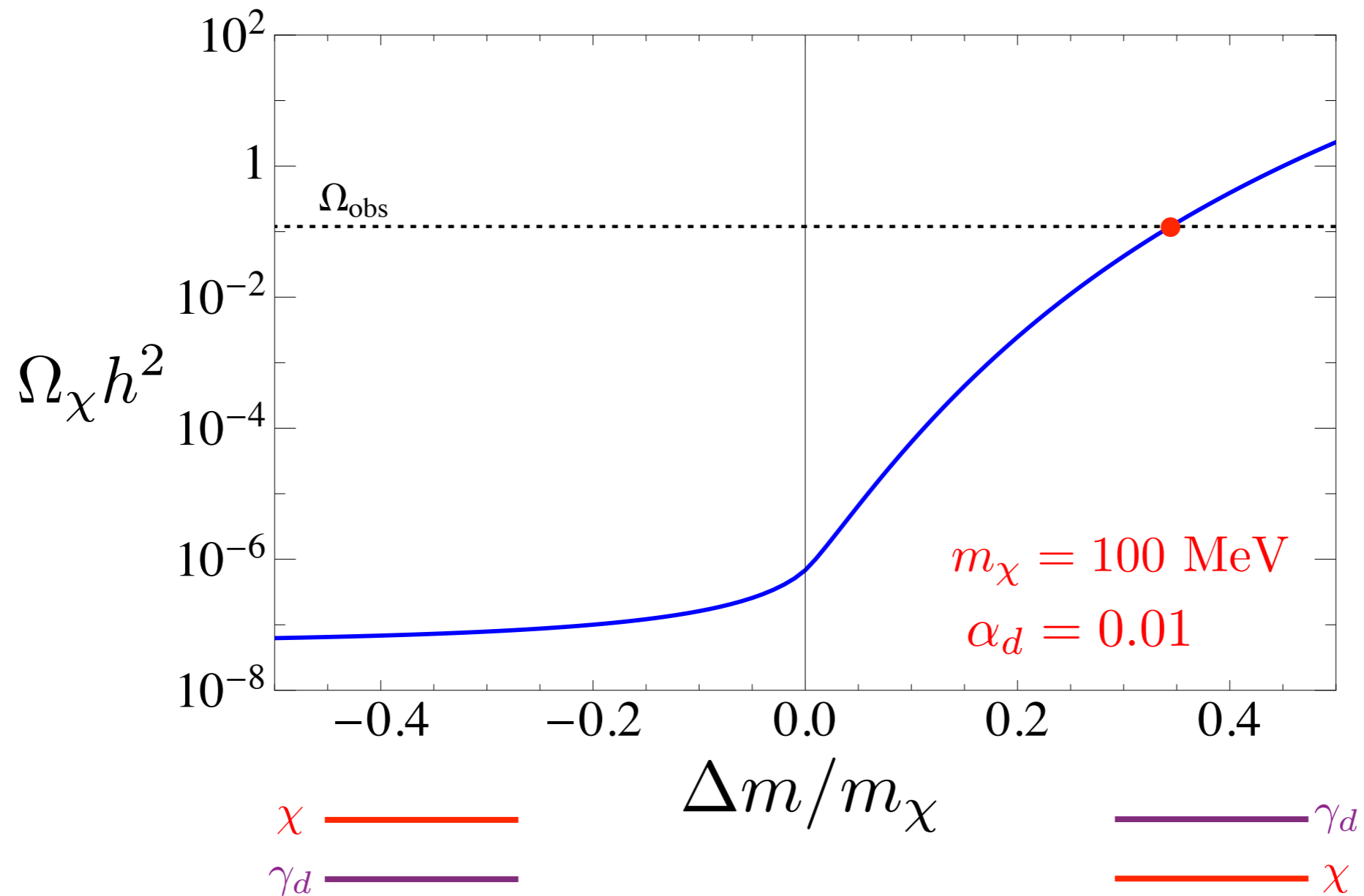
**us**

- allow:  $\frac{\Delta m}{m_\chi} \sim \mathcal{O}(1)$
- predict:  $m_\chi \ll \text{TeV}$
- evades CMB:  $\Delta m \gg T_{\text{rec}}$



# Forbidden Relic Density

$$\Omega_\chi h^2 \sim \frac{m_\chi^2}{\alpha_d^2 T_{eq} M_{pl}} e^{2\Delta m/T_{FO}}$$



# Relic Density with Exponential Sensitivity



## WIMP

$$\Omega_\chi h^2 \sim \frac{m_\chi^2}{\alpha_d^2 T_{eq} M_{pl}}$$

## forbidden

$$\Omega_\chi h^2 \sim \frac{m_\chi^2}{\alpha_d^2 T_{eq} M_{pl}} e^{2\Delta m/T_{FO}}$$

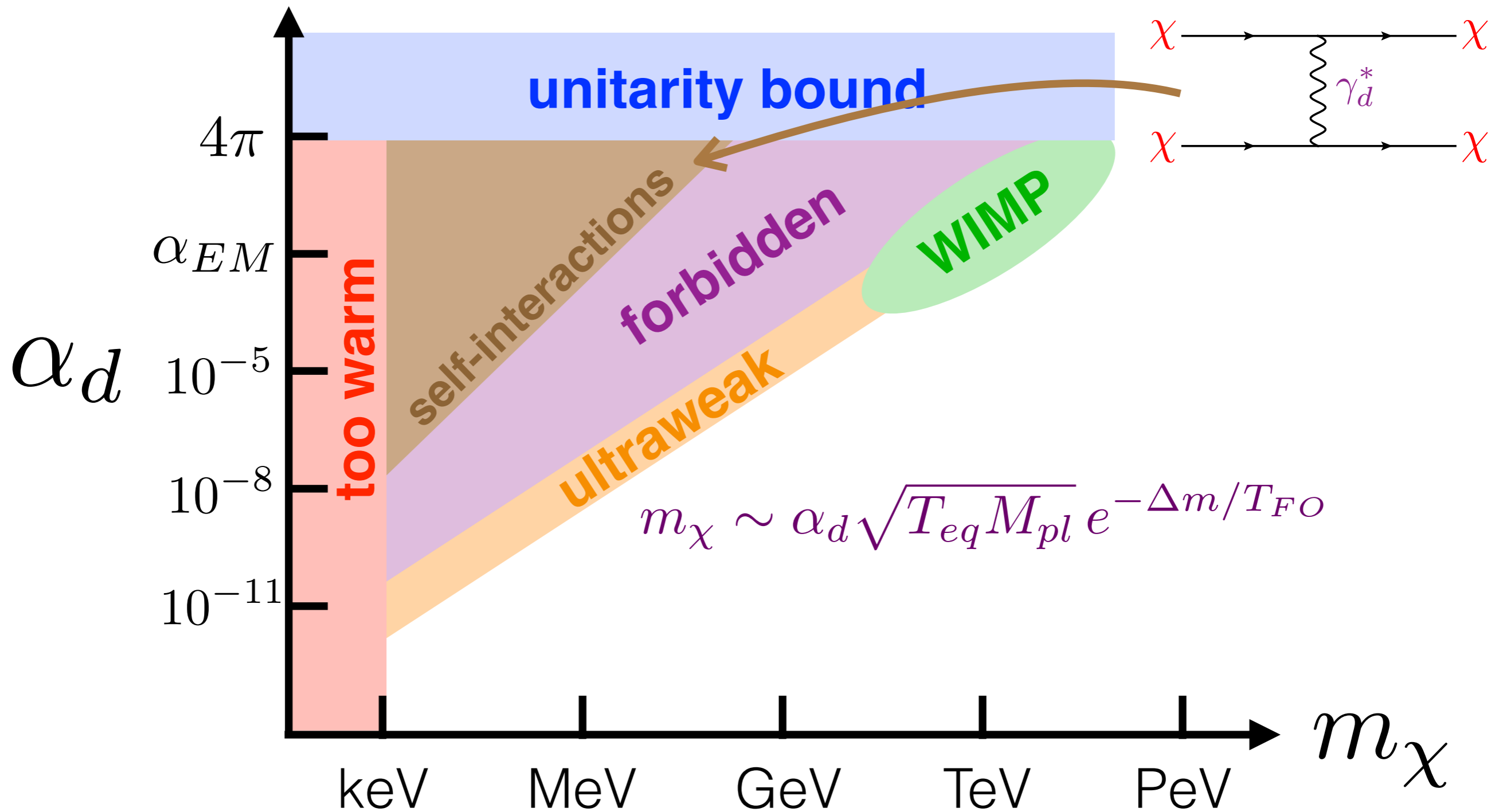
↑  
is this bad?

## QCD

$$\Lambda_{QCD} \approx \Lambda_{UV} e^{-\frac{2\pi}{b_3 \alpha_3(\Lambda_{UV})}} \quad m_p, \Omega_b h^2 \propto e^{-2\pi/(b_3 \alpha_3(\Lambda_{UV}))}$$

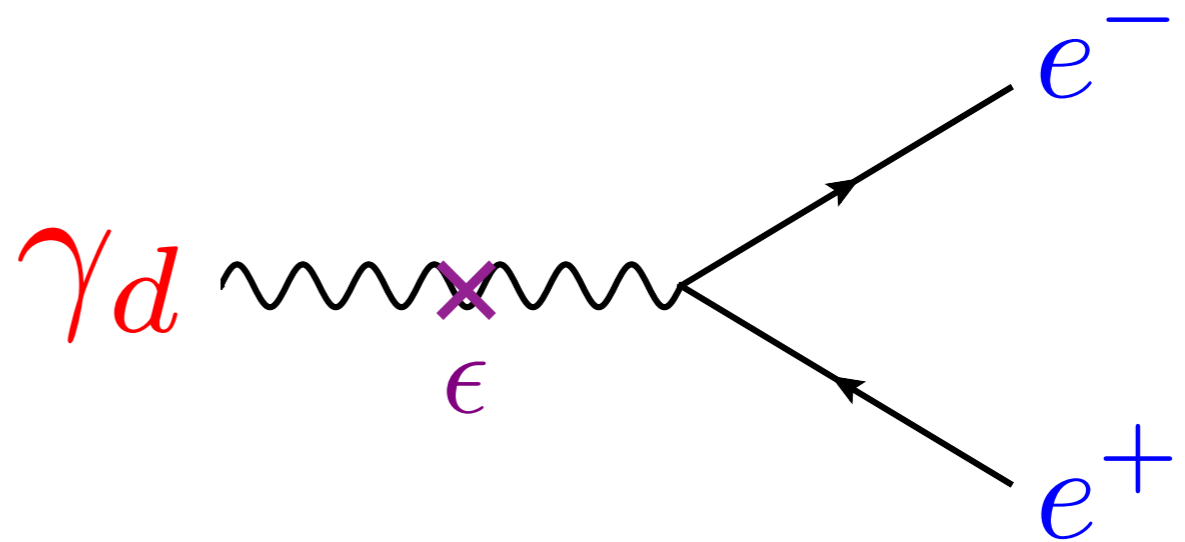
- Anderson, Castano **hep-ph/9409419**

# Dark Freezeout Map



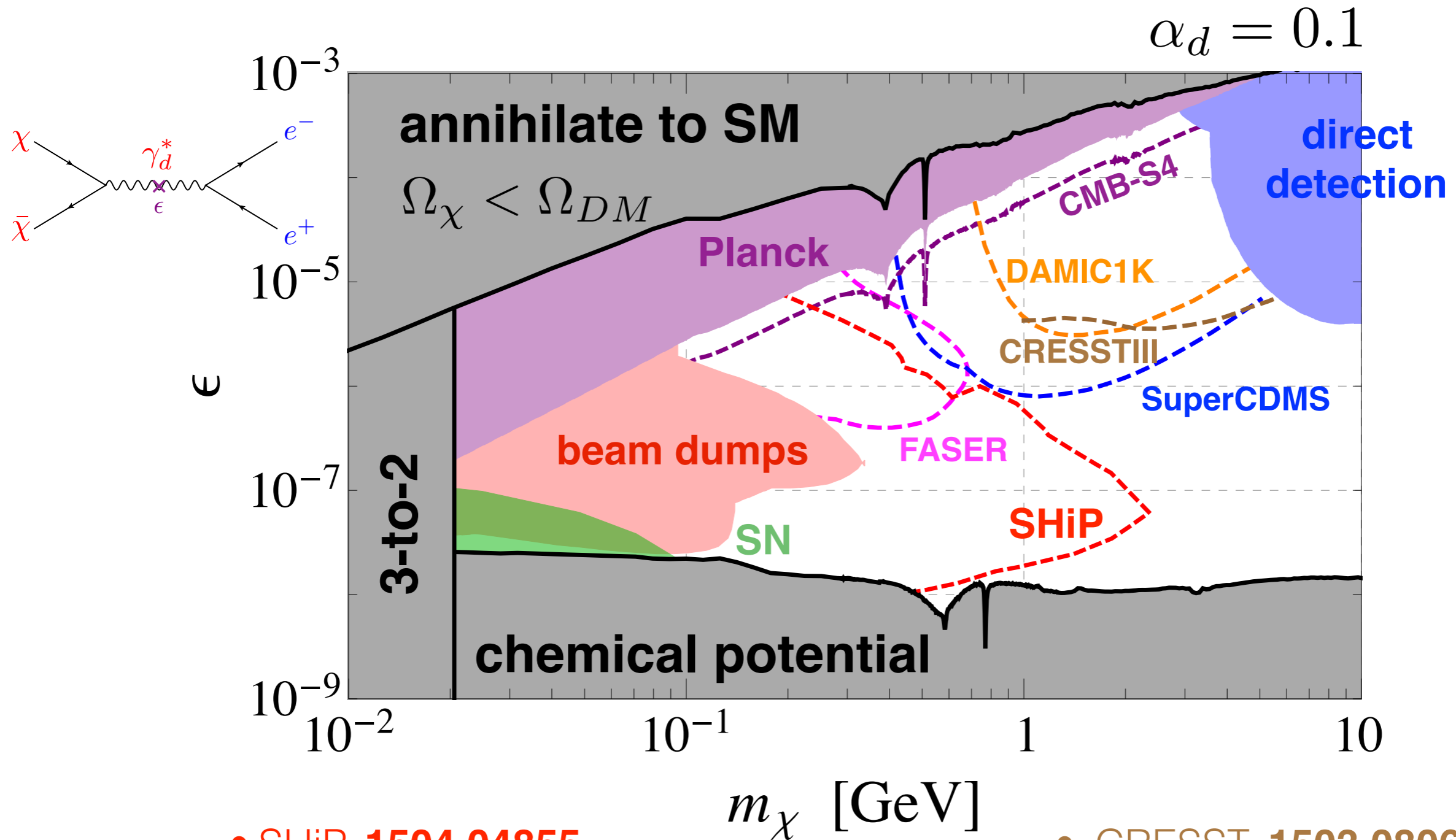
# Forbidden DM with Kinetic Mixing Portal

$$\mathcal{L} \supset \frac{\epsilon}{2} F_{\mu\nu}^d F^{\mu\nu}$$



- Holdom, **1986**

# Forbidden Parameter Space

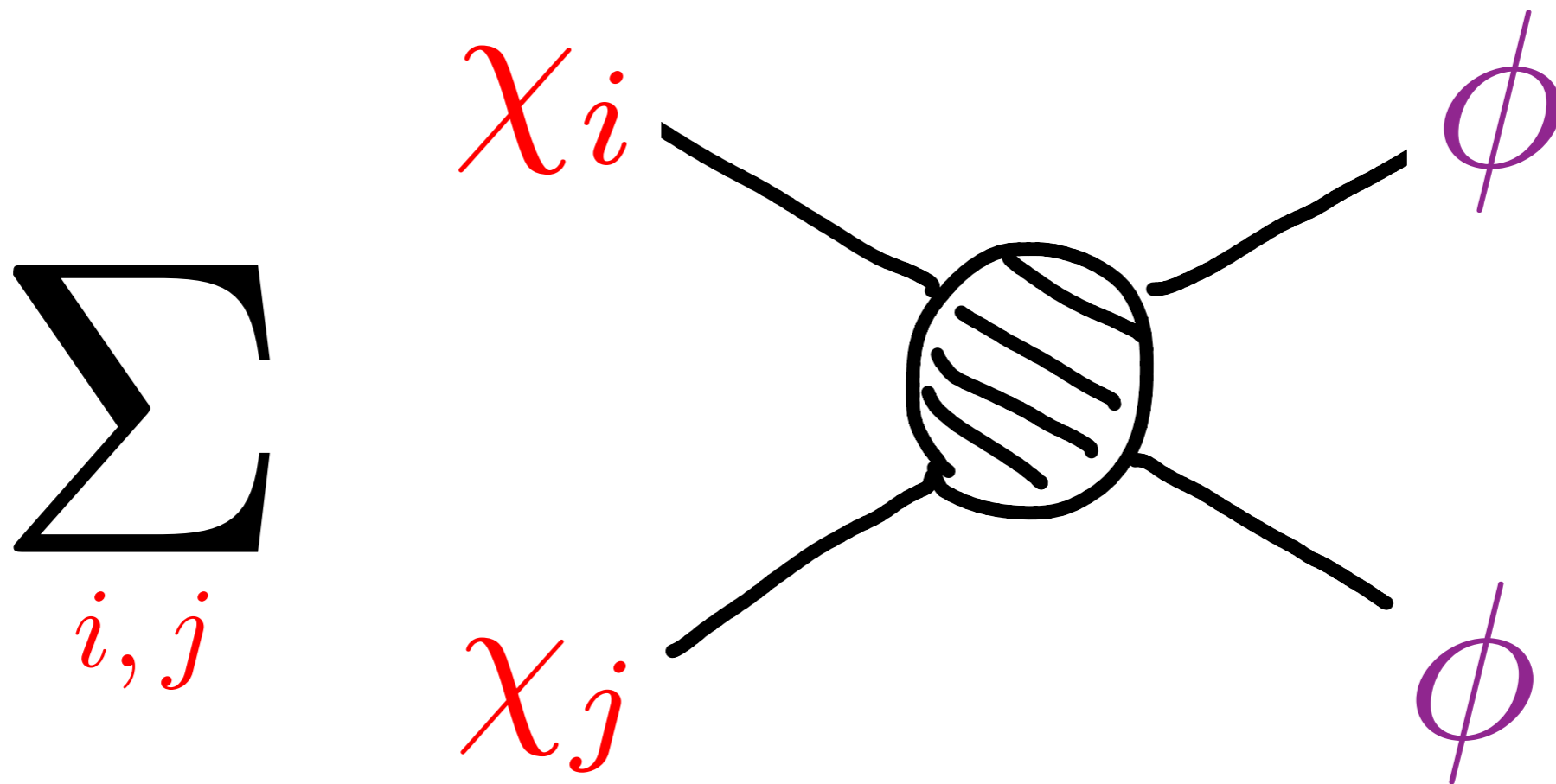


- SHiP, **1504.04855**
- CMB-S4, **1610.02743**
- FASER, **1708.09389**

- CRESST, **1503.08065**
- SuperCDMS, **1610.00006**
- DAMIC1K, **1707.04591**

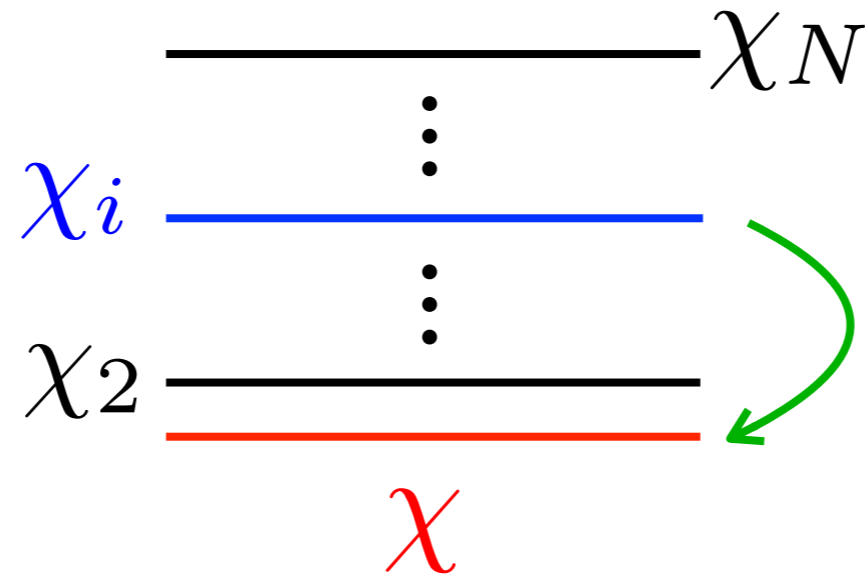


# III. Light DM from Coannihilation



- Raffaele D'Agnolo, Cristina Mondino, JTR, Po-Jen Wang, **1803.02901**

# Coannihilation



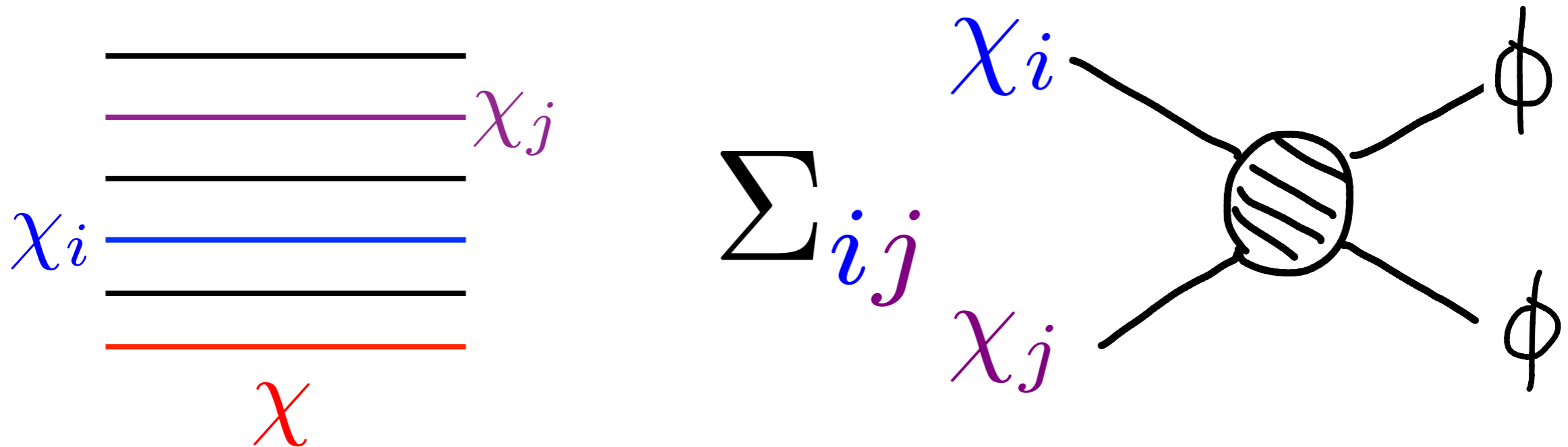
- heavier states decay to dark matter at late times

- *assume* equilibrium among dark states:  $\chi \leftrightarrow \chi_i$

↑  
(foreshadowing)

- Griest, Seckel, **1991**

# Coannihilation

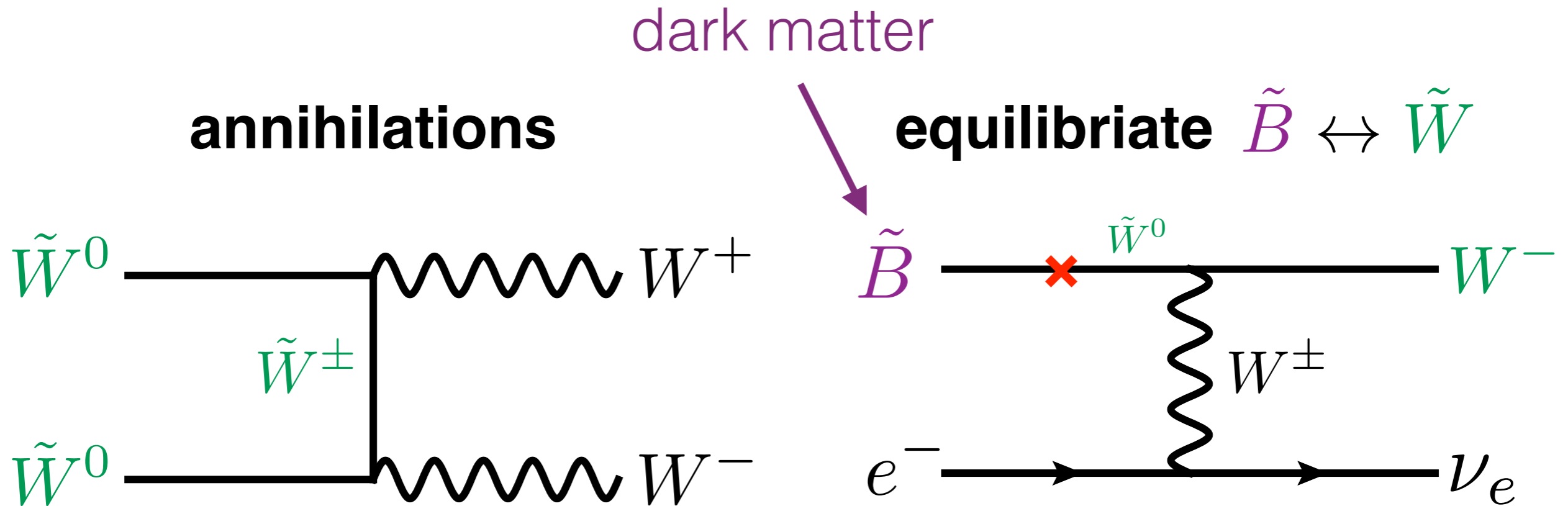


$$\Omega_\chi h^2 \sim \frac{1}{\langle \sigma v \rangle_{eff} T_{eq} M_{pl}}$$

$$\langle \sigma v \rangle_{eff} = \sum_{i,j} \frac{n_i^{eq} n_j^{eq}}{(n_\chi^{eq})^2} \langle \sigma_{ij} v \rangle$$

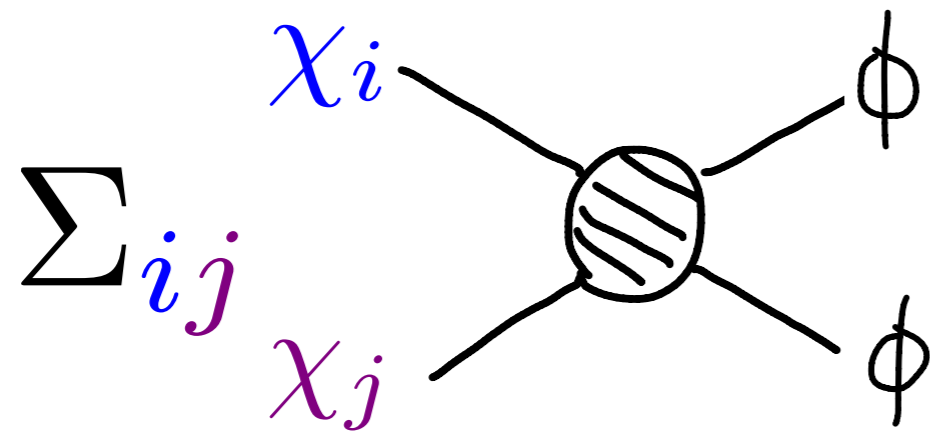
- Griest, Seckel, **1991**

# ex) Well-Tempered Neutralino



- Arkani-Hamed, Delgado, Giudice **hep-ph/0601041**
- Cheung, Hall, Pinner, JTR **1211.4873**

# Degenerate Coannihilations



$$\langle \sigma v \rangle_{eff} = \sum_{i,j} \frac{n_i^{eq} n_j^{eq}}{(n_\chi^{eq})^2} \langle \sigma_{ij} v \rangle$$

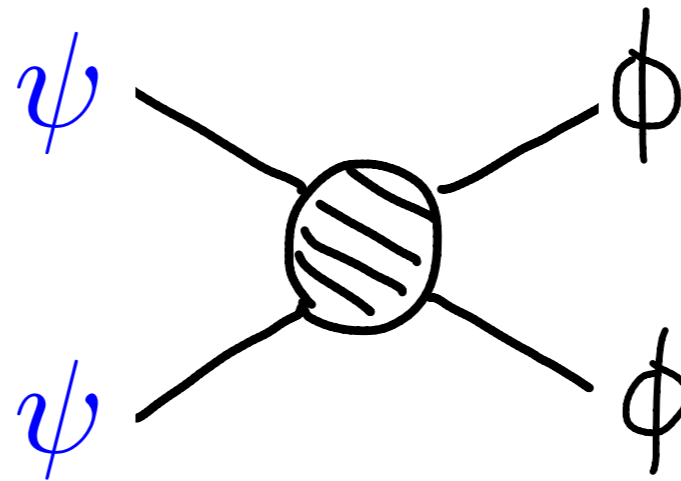
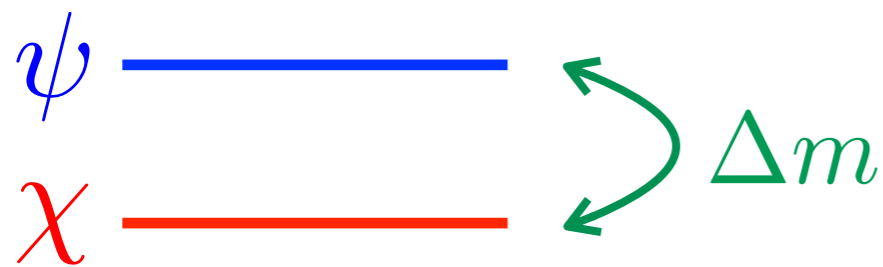
$$\frac{n_i^{eq}}{n_\chi^{eq}} \sim e^{-\Delta m/T}$$

- lore: coannihilations is only relevant for degenerate states

$$\frac{\Delta m}{m_\chi} \lesssim 0.1$$

# Non-Degenerate Coannihilations

dominate annihilation:



$$\langle \sigma v \rangle \sim \frac{\alpha_d^2}{m_\psi^2}$$

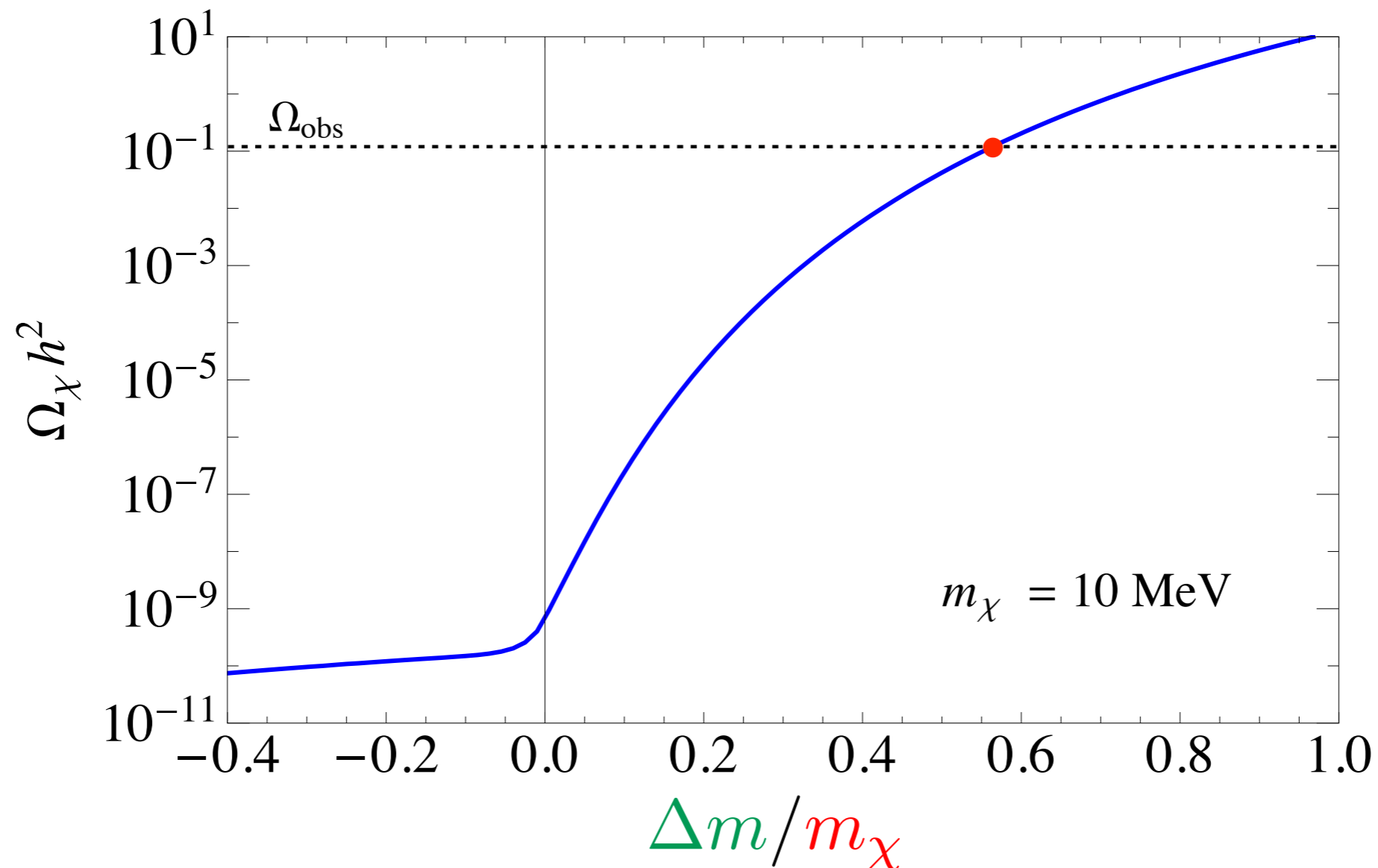
$$\langle \sigma v \rangle_{eff} \sim \frac{\alpha_d^2}{m_\psi^2} e^{-2\Delta m/T}$$

- DM mass:  $m_\chi \sim \alpha_d \sqrt{T_{eq} M_{pl}} e^{-2\Delta m/T_{FO}}$

- evades CMB:  $\sigma_{\chi\chi} \ll \sigma_{\psi\psi}$

# Non-Degenerate Coannihilations

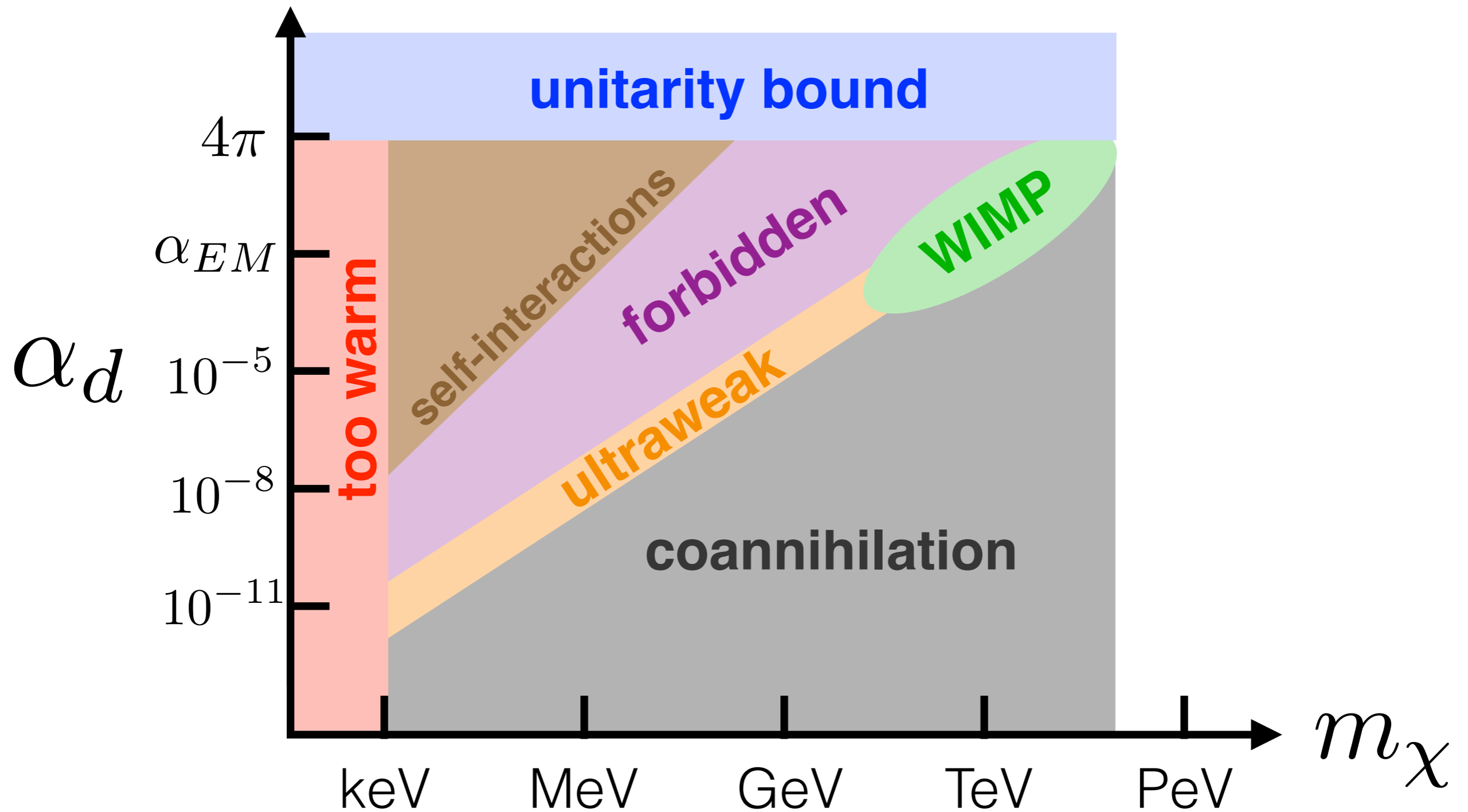
$$\Omega_\chi h^2 \sim \frac{m_\chi^2}{\alpha_d^2 T_{eq} M_{pl}} e^{2\Delta m/T_{FO}}$$



$\chi$  ———  
 $\psi$  ———

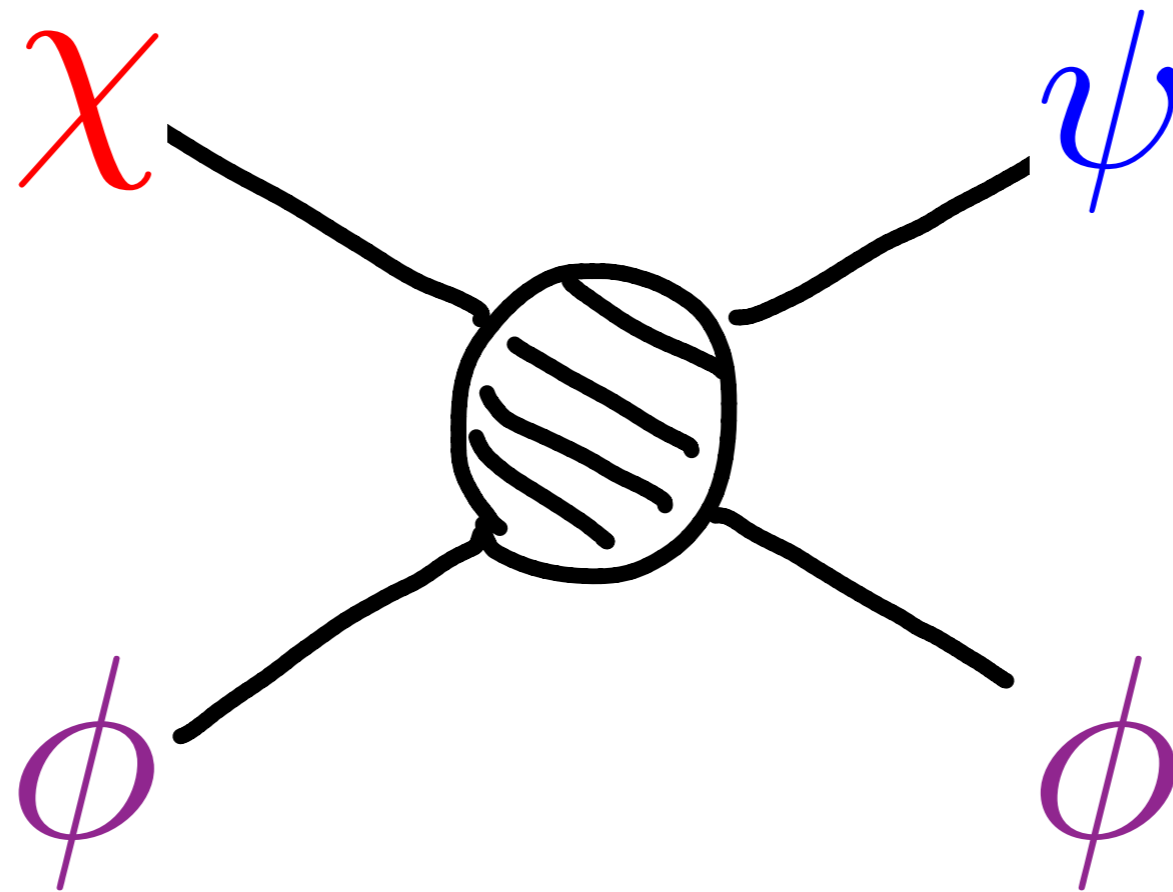
—————  $\psi$   
 —————  $\chi$

# Dark Freezeout Map



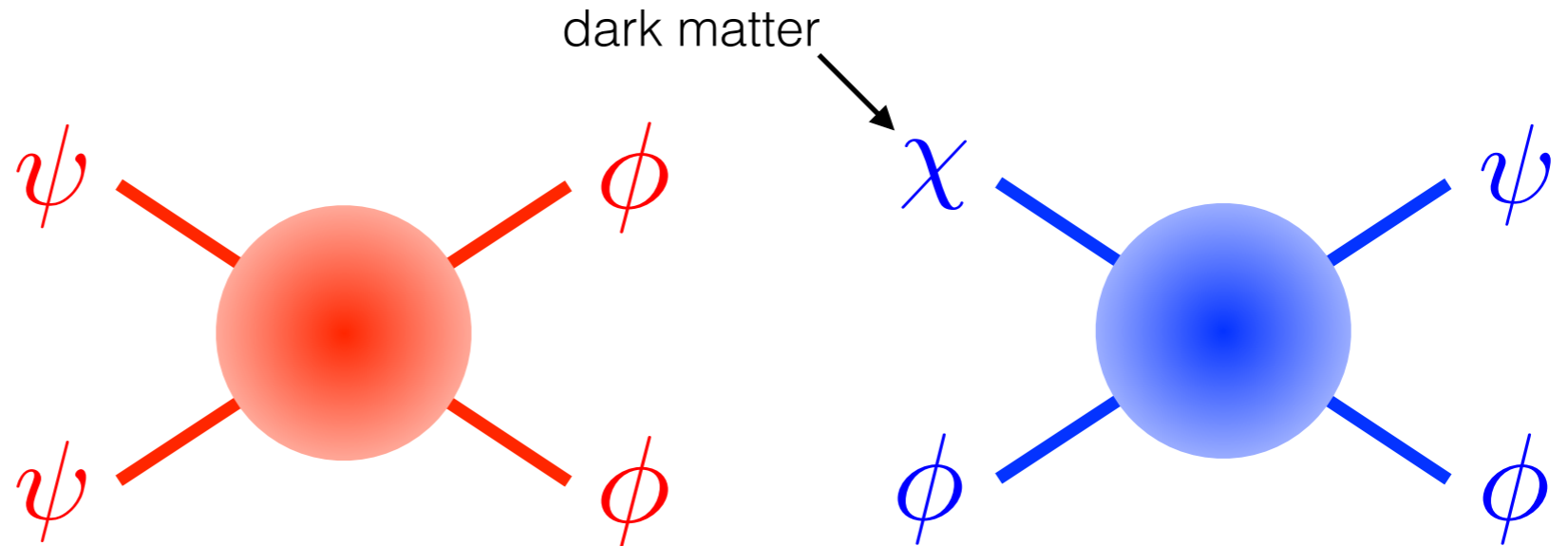
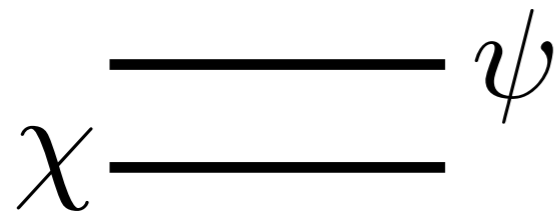


# IV. Coscattering

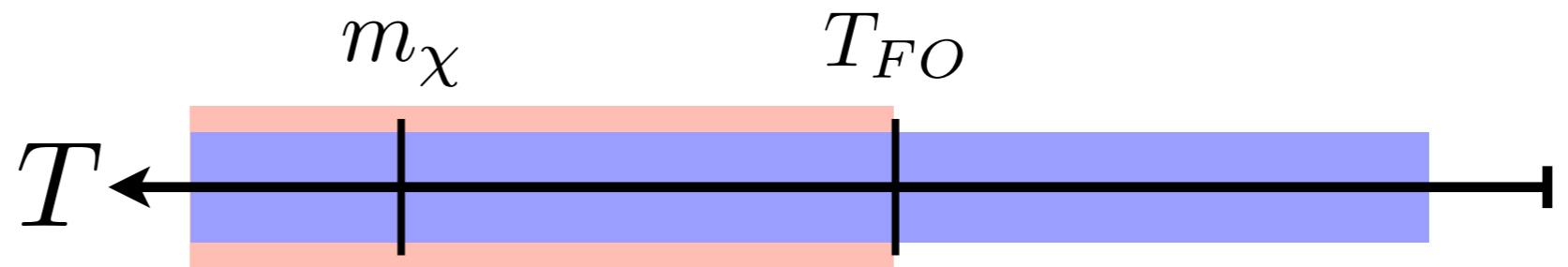


- Raffaele D'Agnolo, Duccio Pappadopulo, JTR, **1705.08450** (PRL)

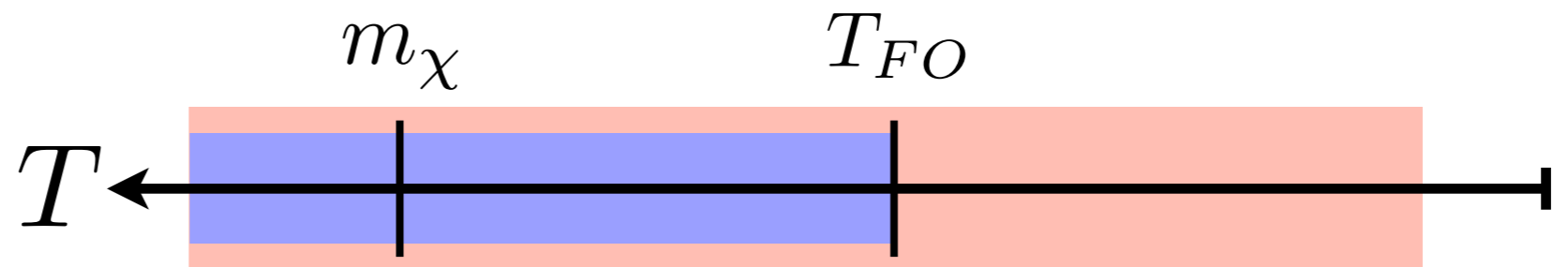
# Coannihilation vs. Coscattering



- coannihilation



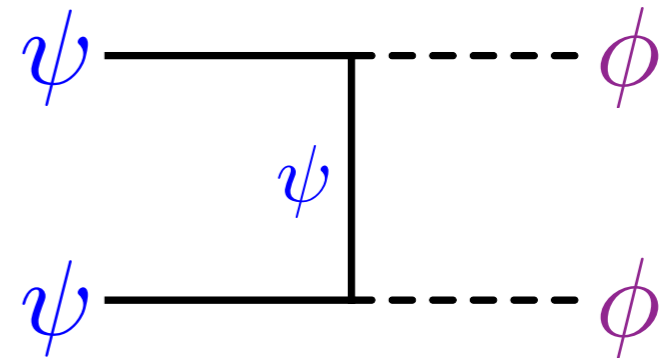
- coscattering



# Mix-In Dark Matter

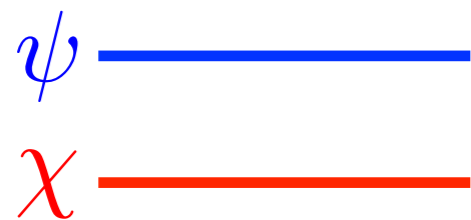
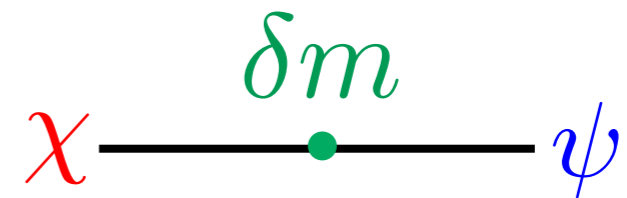
- active state

$$\frac{y}{2} \phi \psi^2$$



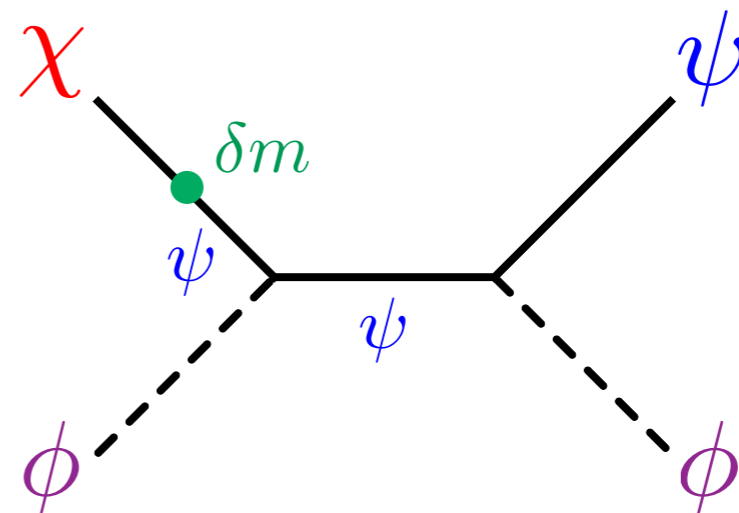
- sterile state (DM)

$$\delta m \chi \psi$$



$$\delta m \ll m_\chi$$

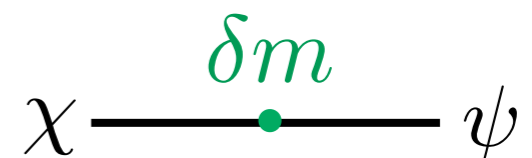
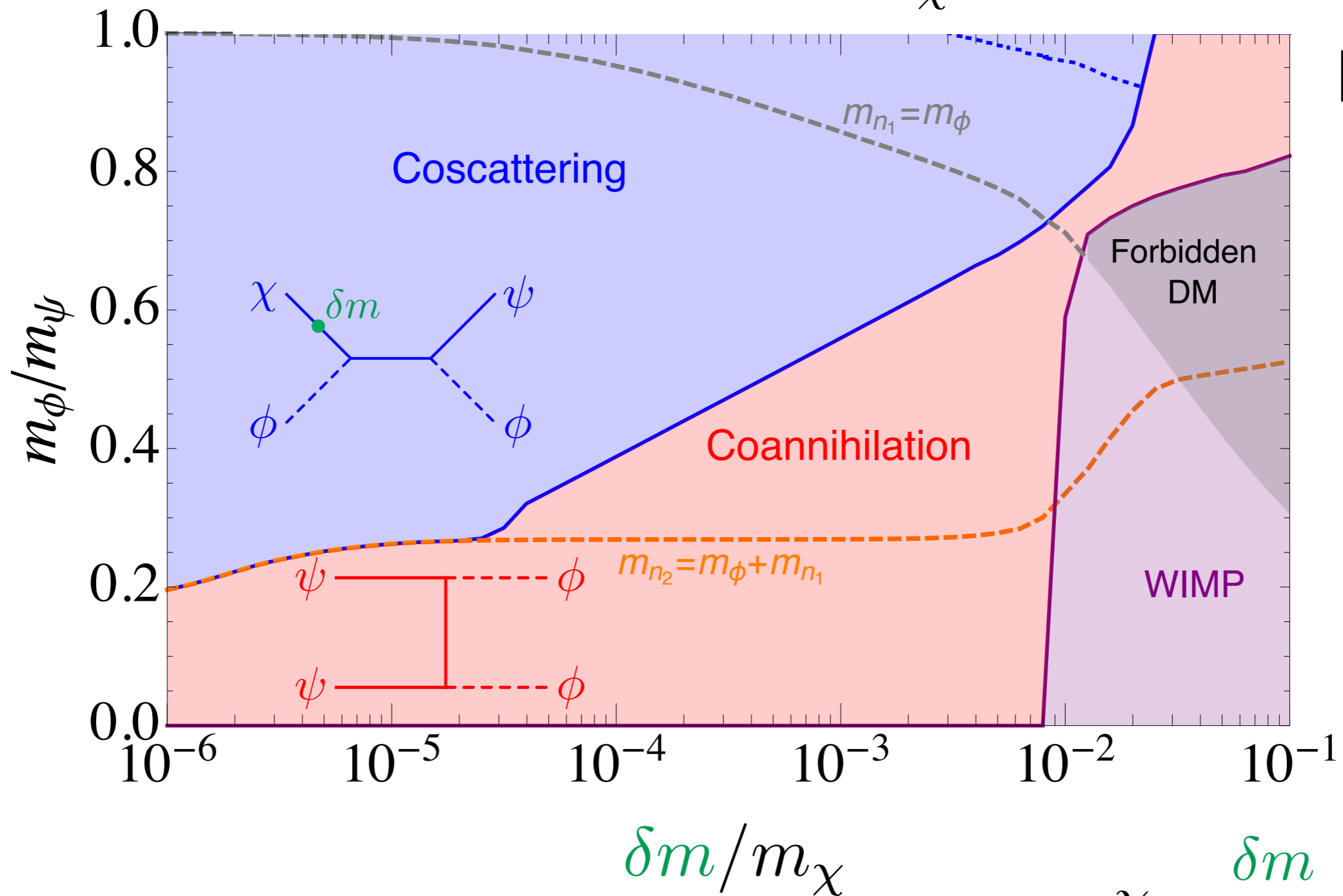
- equilibriate:  $\chi \leftrightarrow \psi$



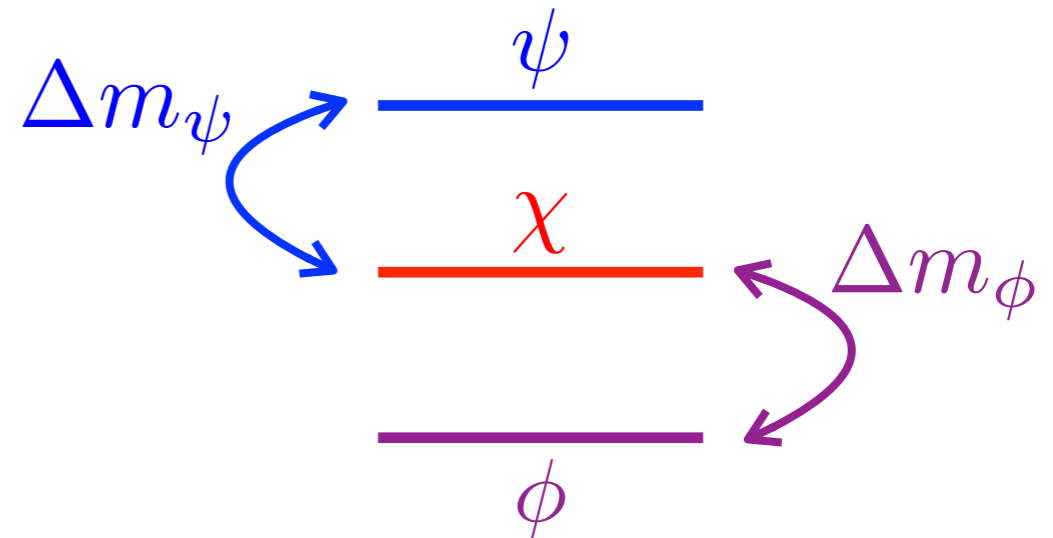
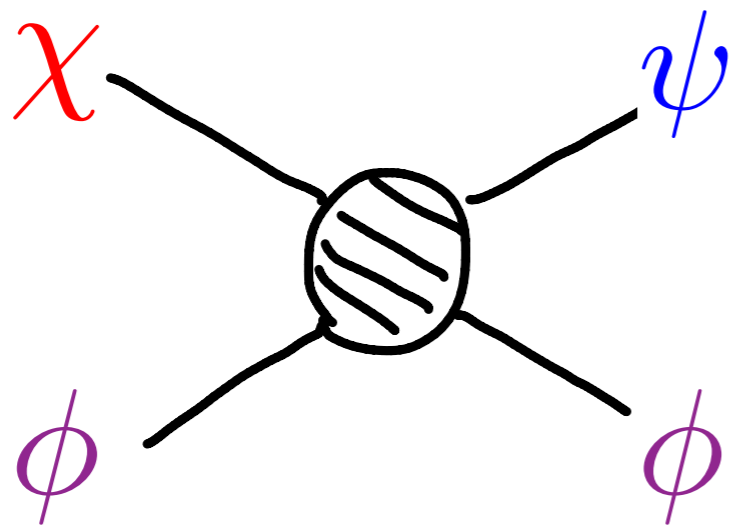
# Freeze-Out Phase Diagram

$$m_\chi = 100 \text{ MeV}$$

$$|y| = 0.5$$



# Coscattering Abundance



$$n_{\phi}^{eq} \langle \sigma v \rangle = H$$

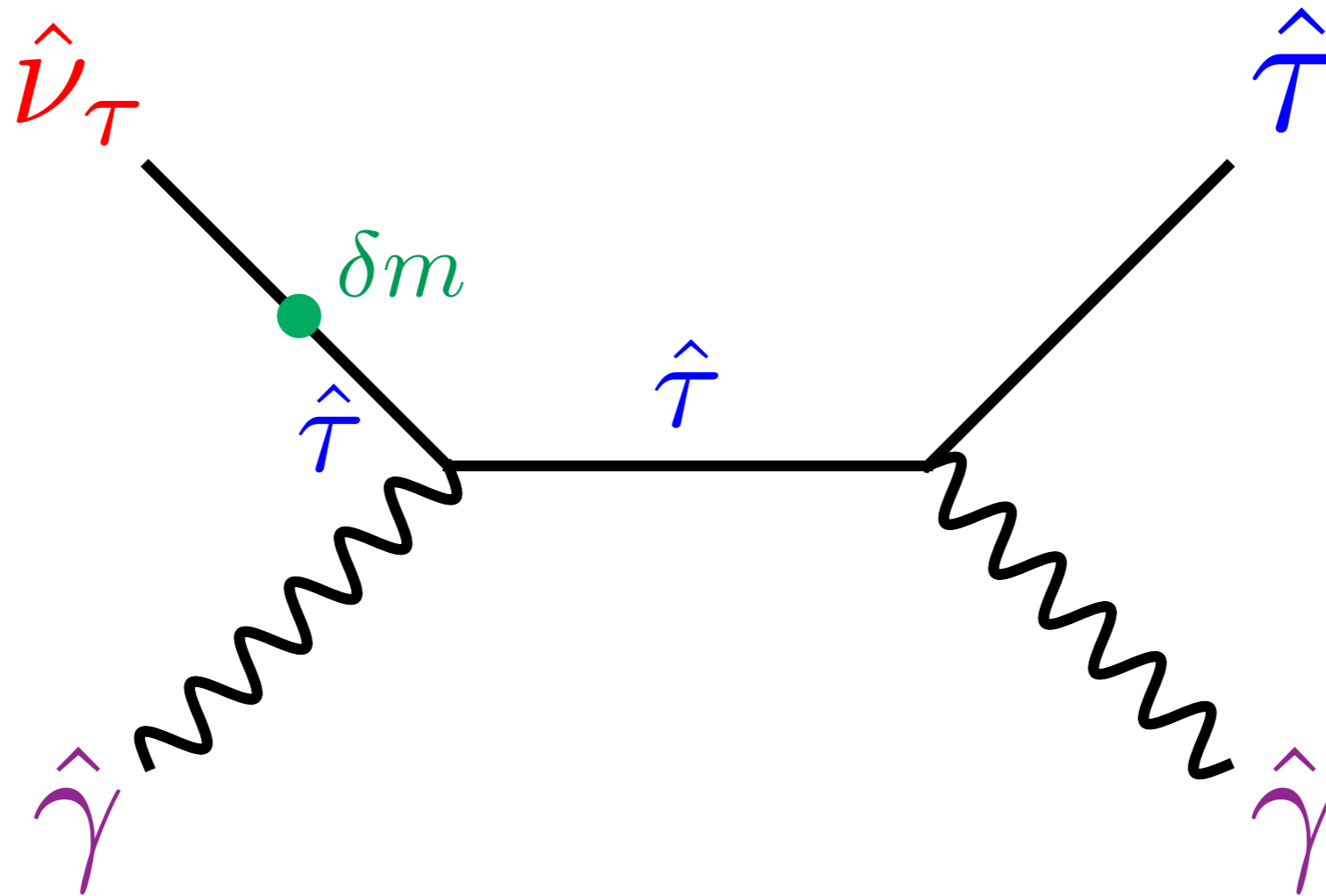
freeze-out

$$m_{\chi} \sim \alpha_d \sqrt{T_{eq} M_{pl}} e^{-(\Delta m_{\psi} - \Delta m_{\phi}) / 2T_{FO}}$$

thermal average

$$\langle \sigma v \rangle \sim \frac{\alpha_d^2}{m_{\chi}^2} e^{-\Delta m_{\psi} / T}$$

# Coscattering in Fraternal Twin Higgs

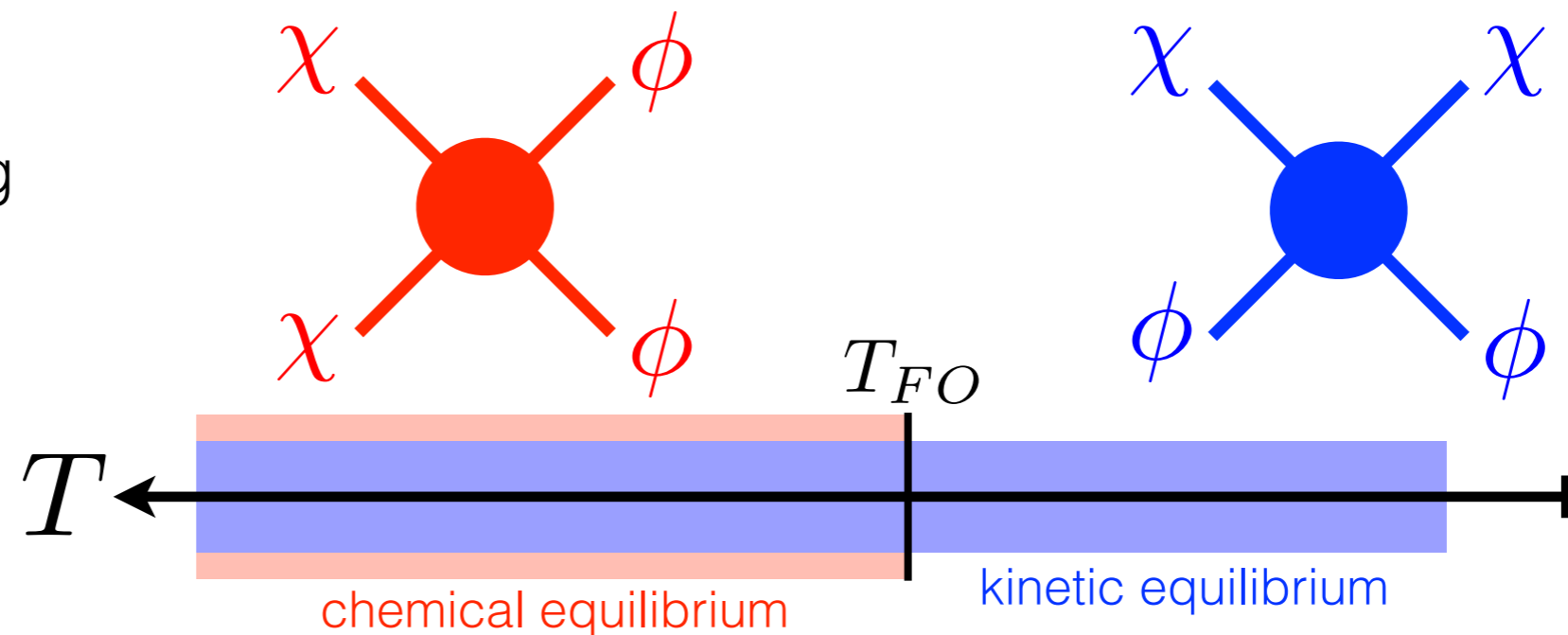


$$m_{\hat{\gamma}}, m_{\hat{\nu}_\tau}, m_{\hat{\tau}} \sim 0.1 - 10 \text{ GeV}$$

# Chemical vs. Kinetic Decoupling

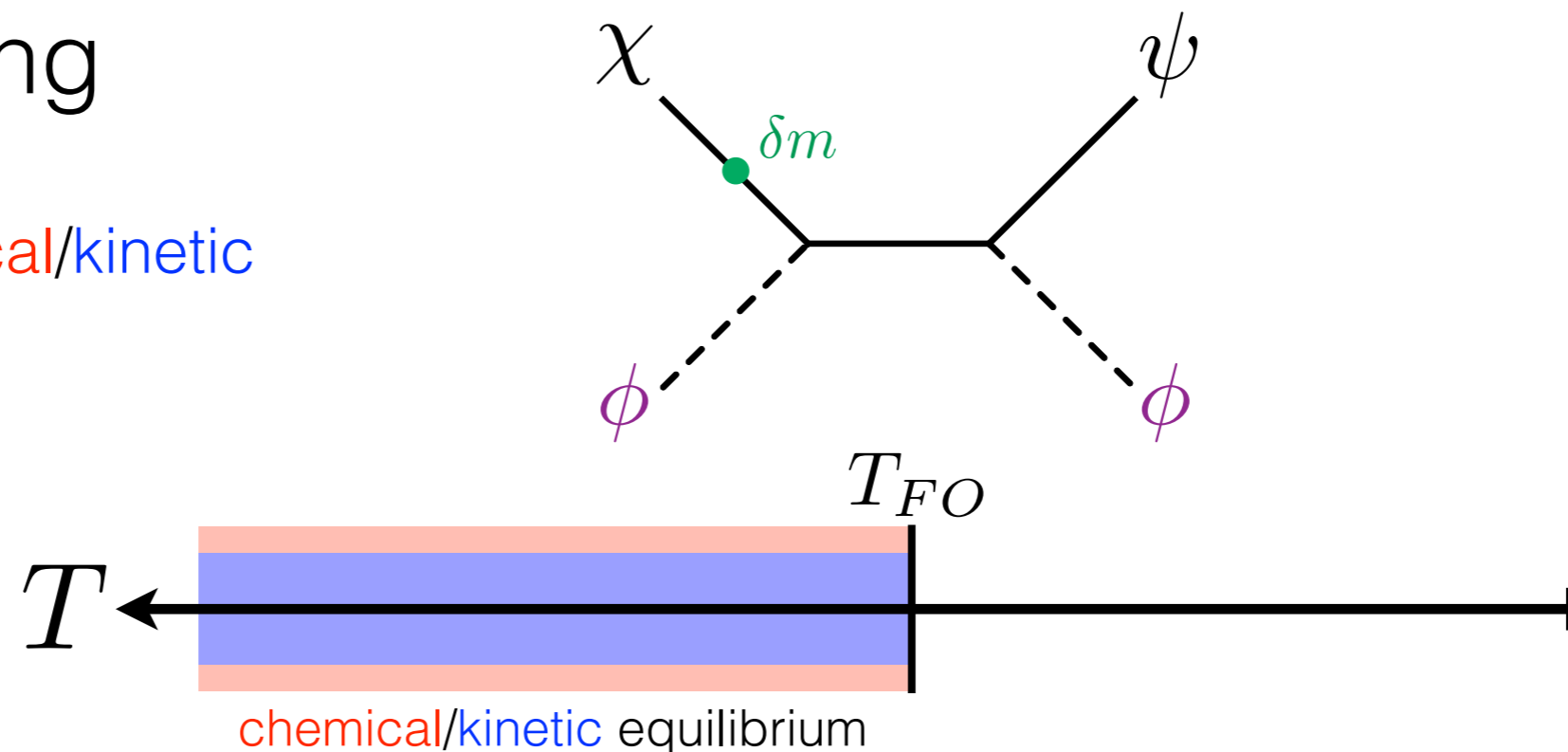
- WIMP/forbidden/coannihilations

chemical decoupling  
before kinetic



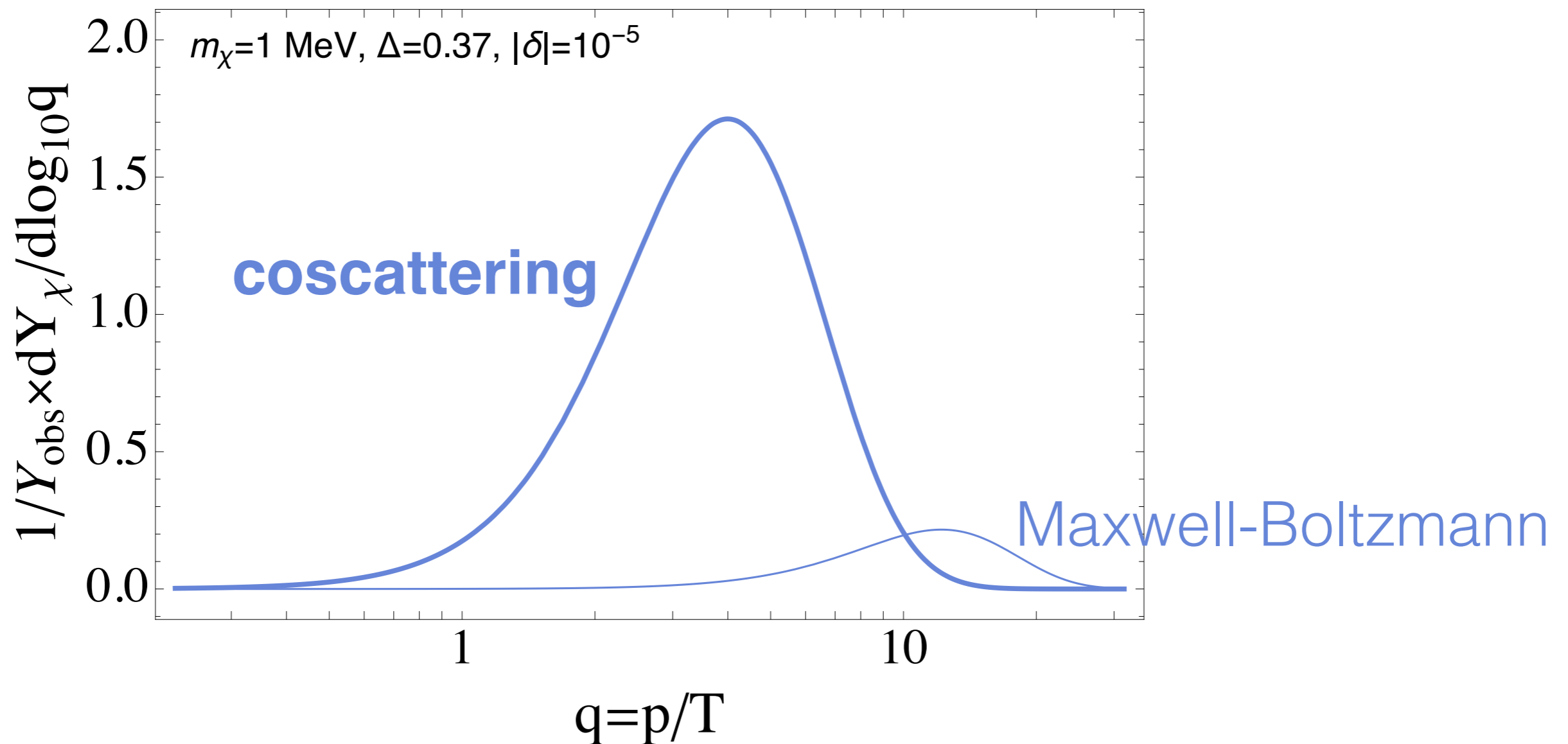
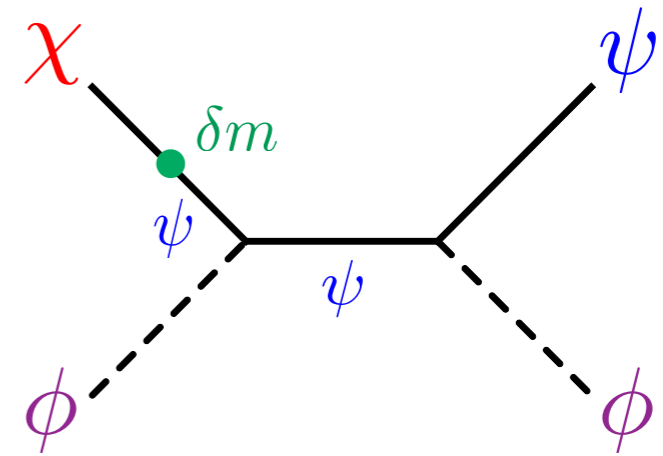
- cospattering

simultaneous chemical/kinetic  
decoupling



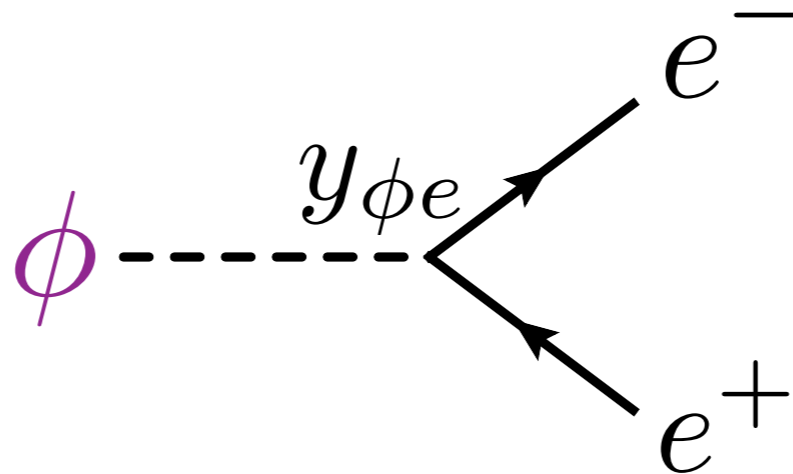
# Distorted Dark Matter Phase Space

low-momentum modes decouple before high-momentum modes

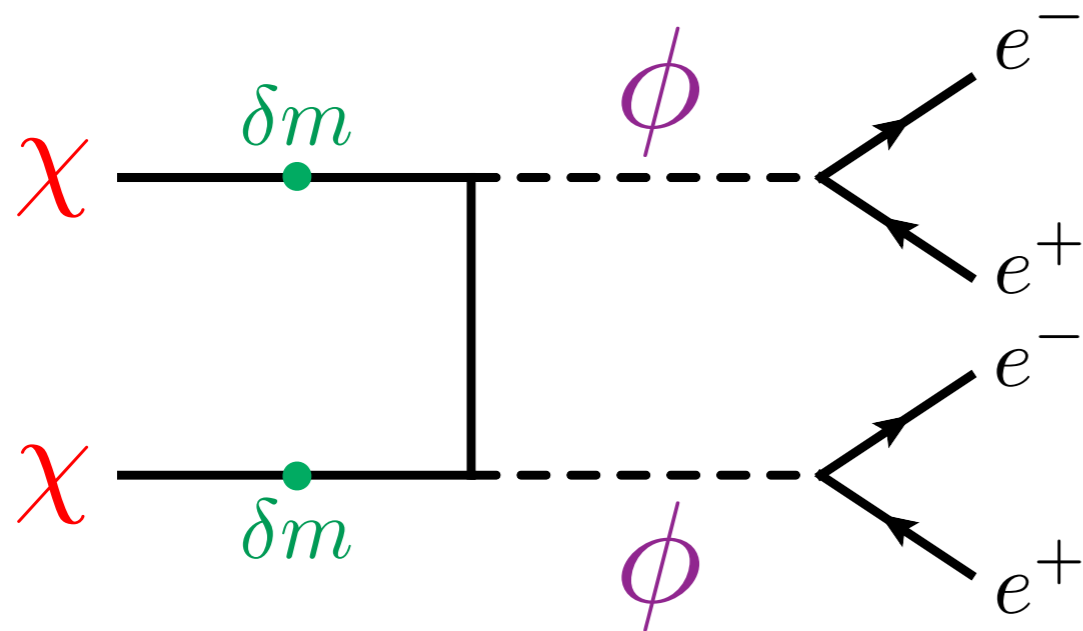




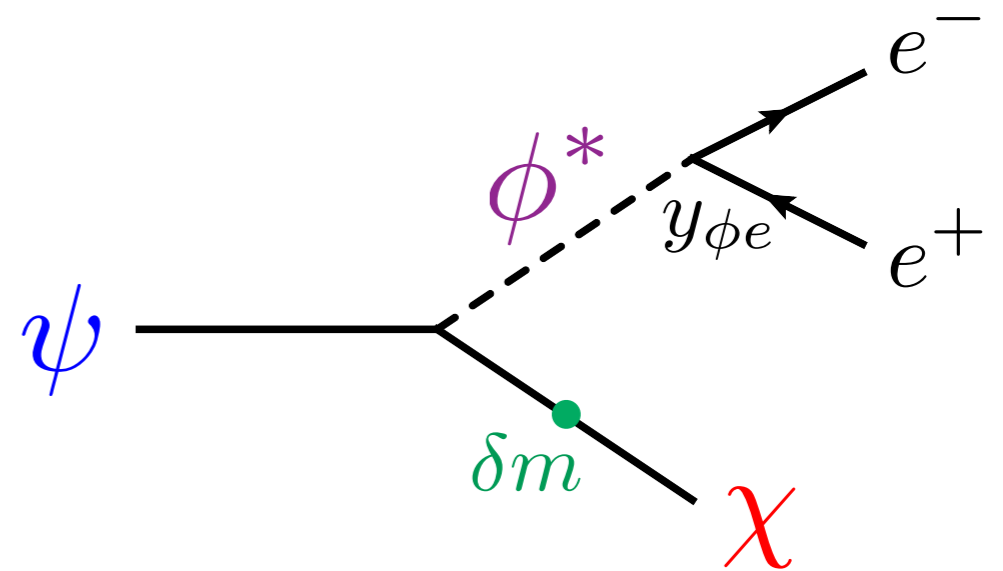
# Coscaterring Coupled to SM



**annihilations**

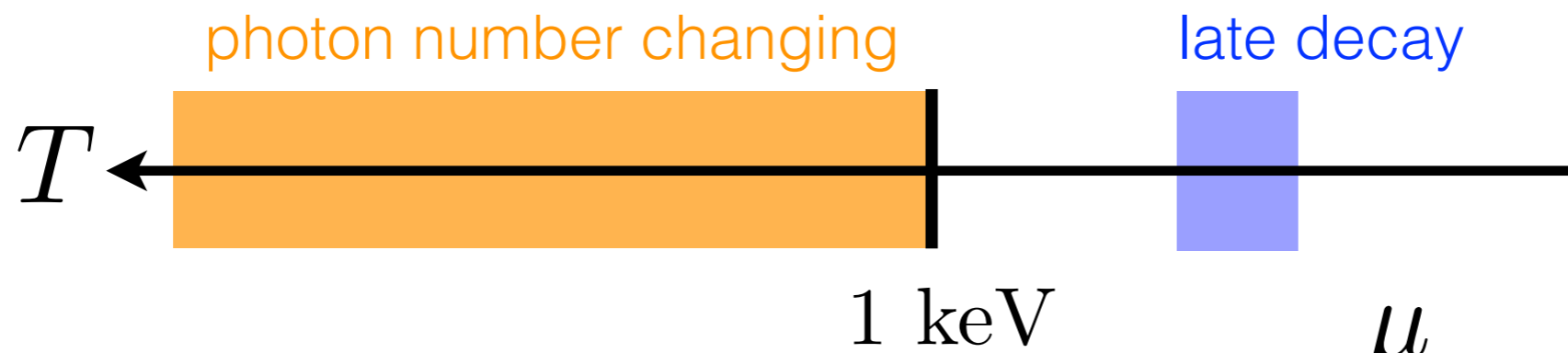
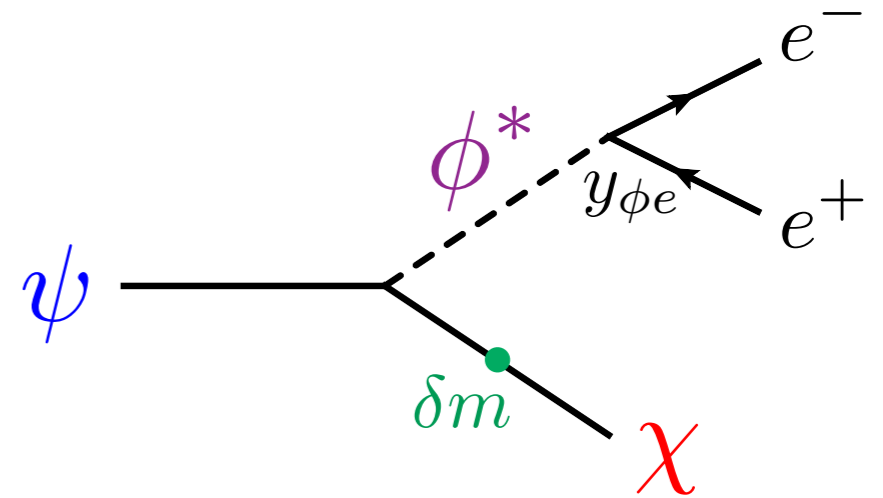
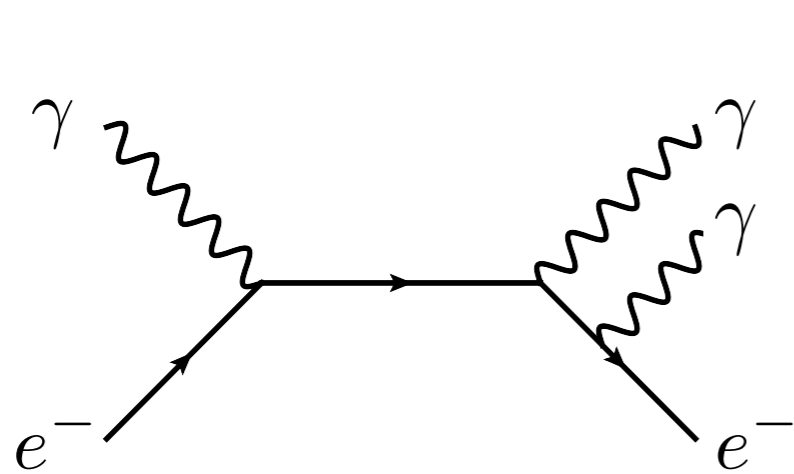


**decay**



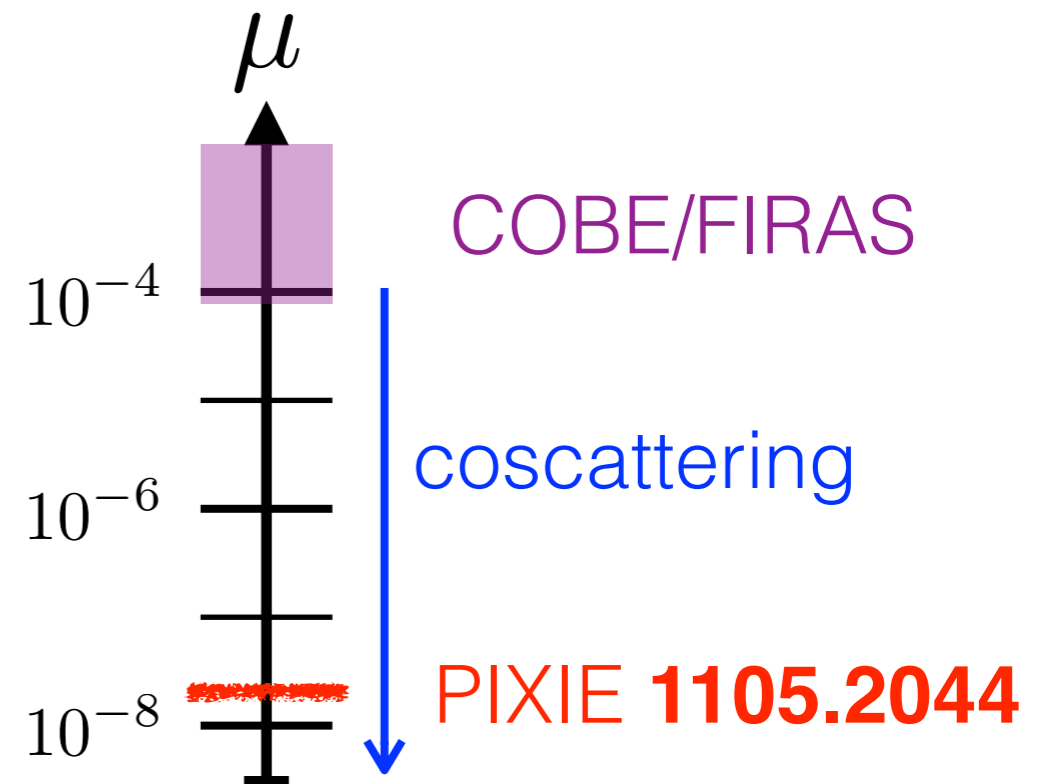
$$\langle \sigma v \rangle_{\chi\chi} \ll \langle \sigma v \rangle_{\text{WIMP}} \sim \frac{1}{T_{eq} M_{pl}}$$

# Spectral Distortions from Coscattering

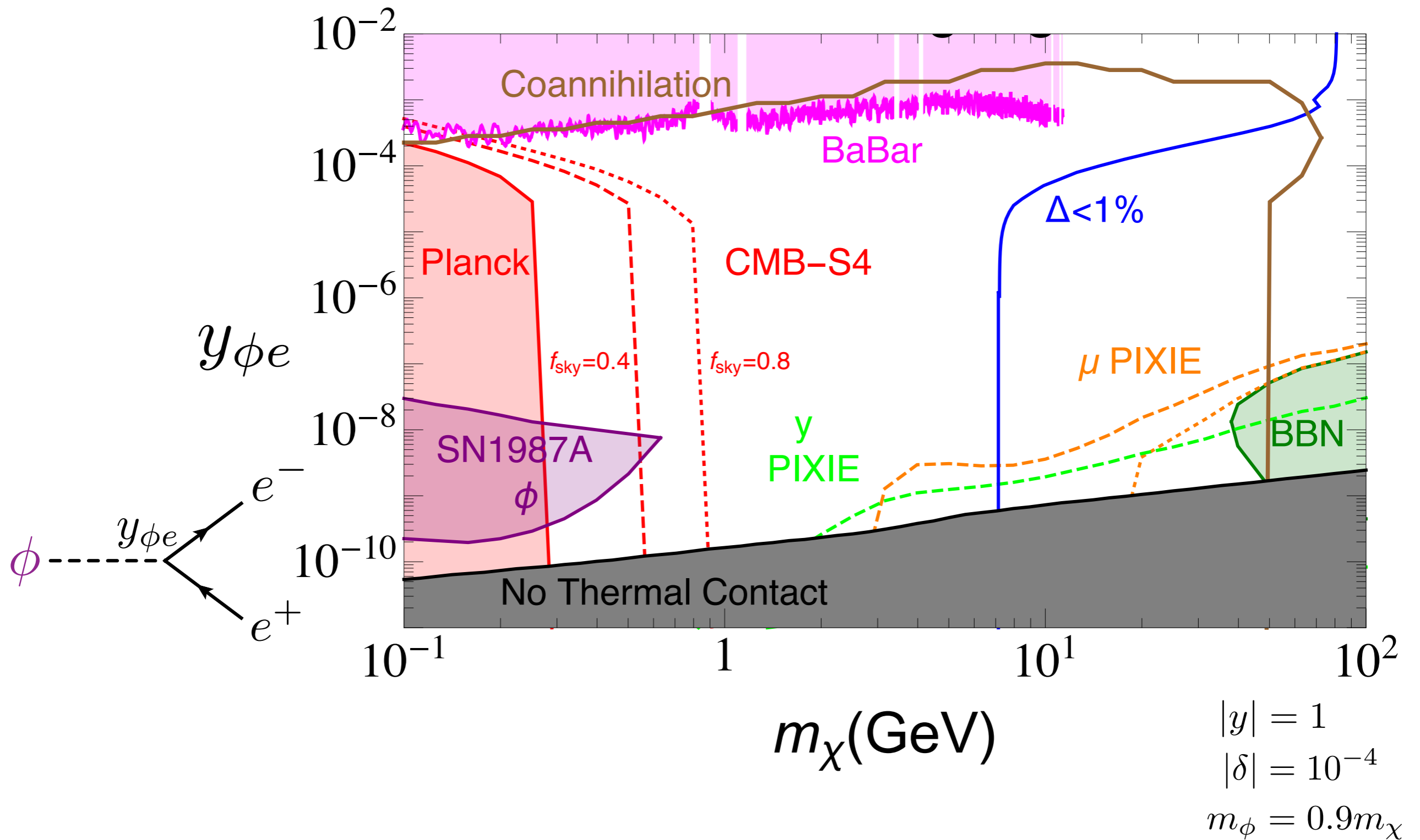


$\mu$  - distortion

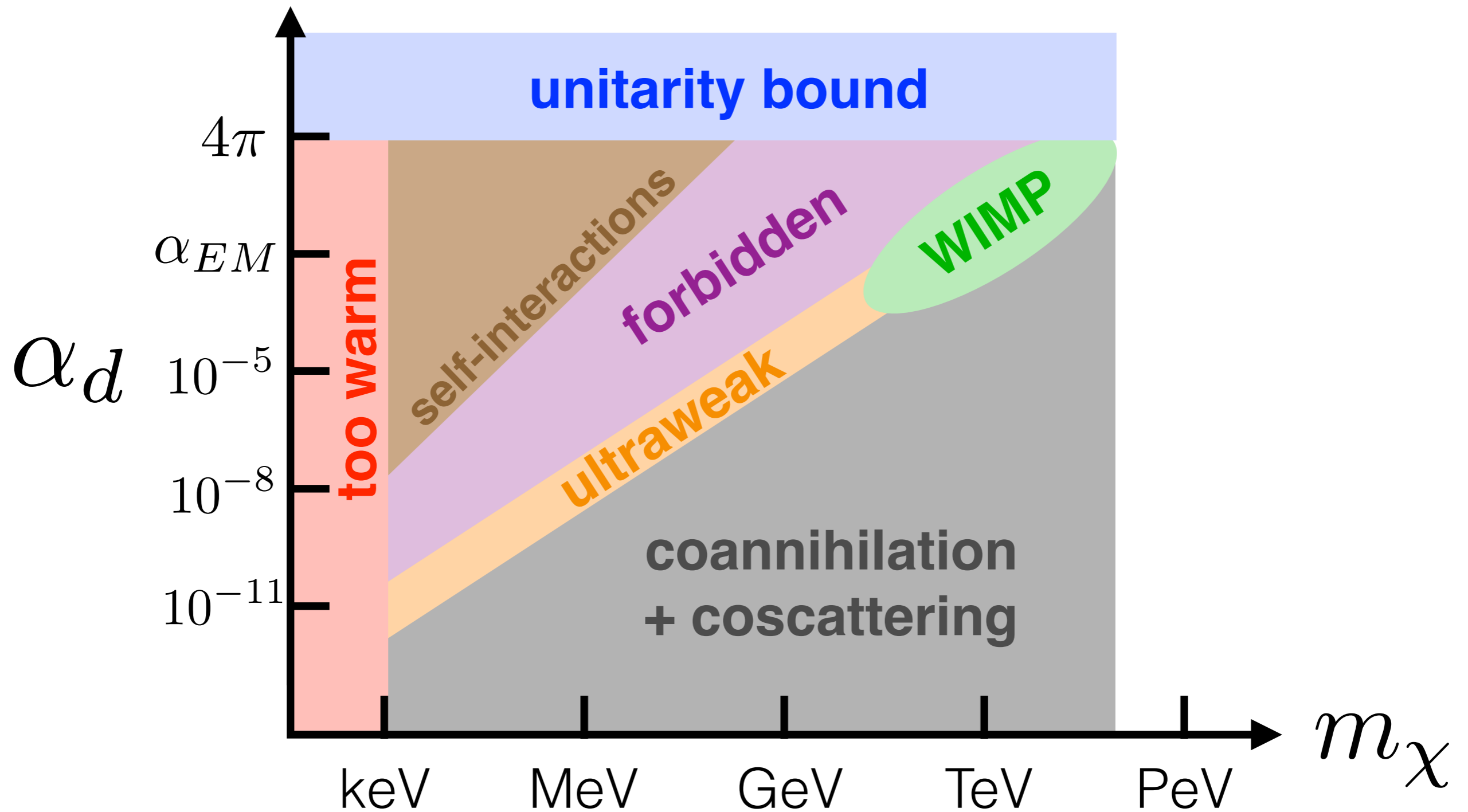
$$f_{\gamma}(E) = \frac{1}{e^{E/(kT)+\mu} - 1}$$



# Coscattering Parameter Space



# Dark Freezeout Map



Mar. 29, 1976

THE

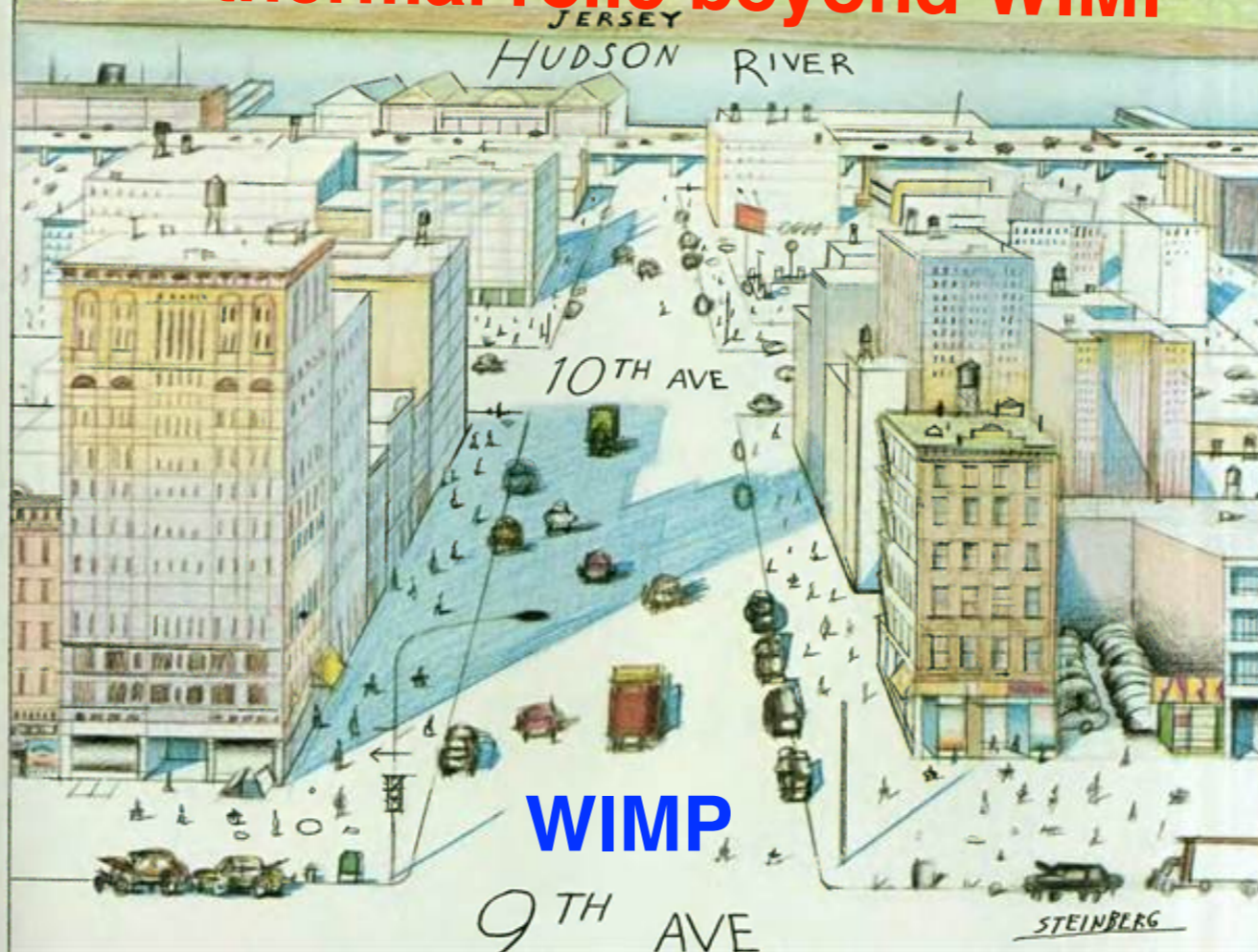
Price 75 cents

# NEW YORKER

beyond thermal relic

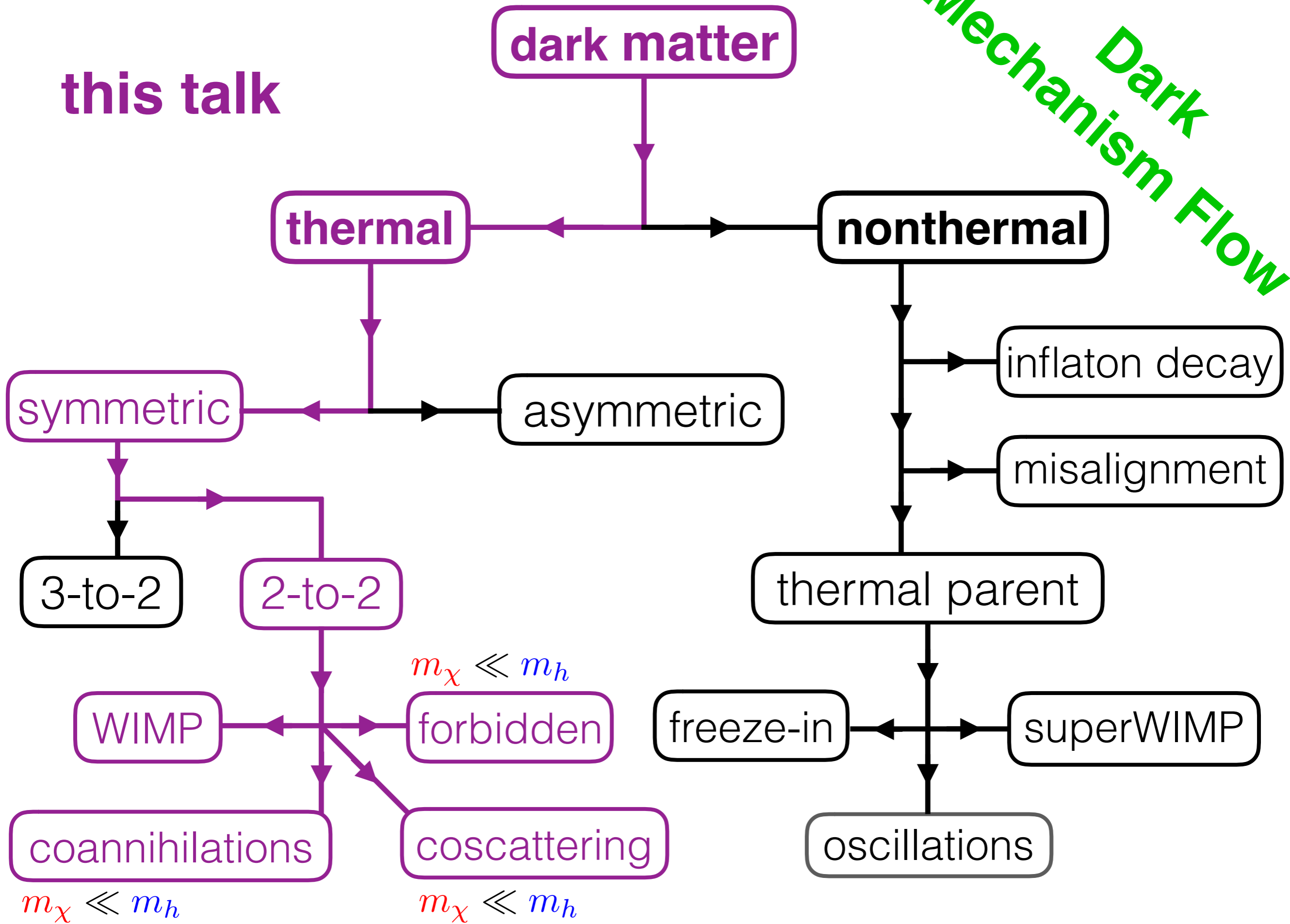


thermal relic beyond WIMP

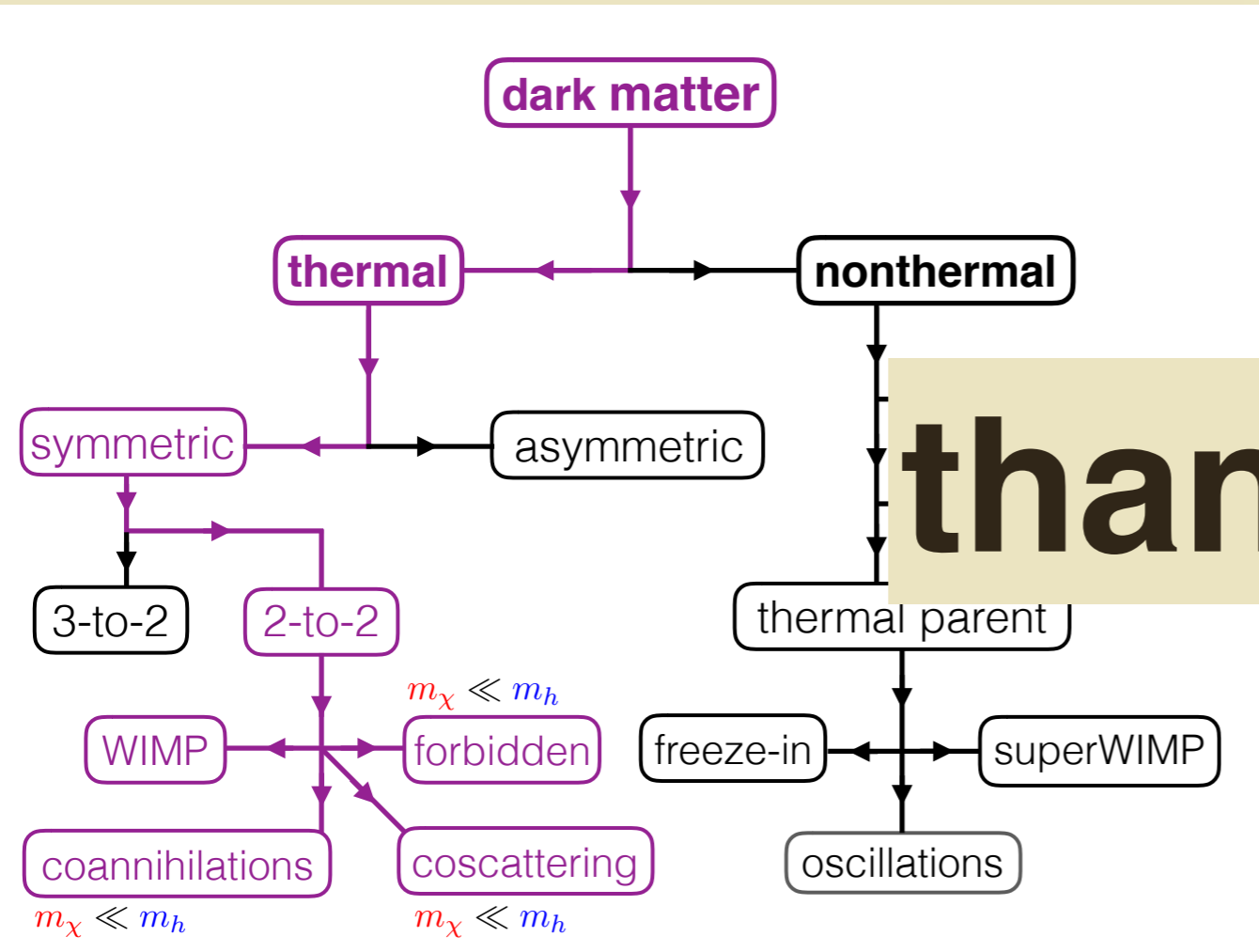


WIMP

this talk



Mechanism Dark Flow



thanks!

