

Summer at CERN

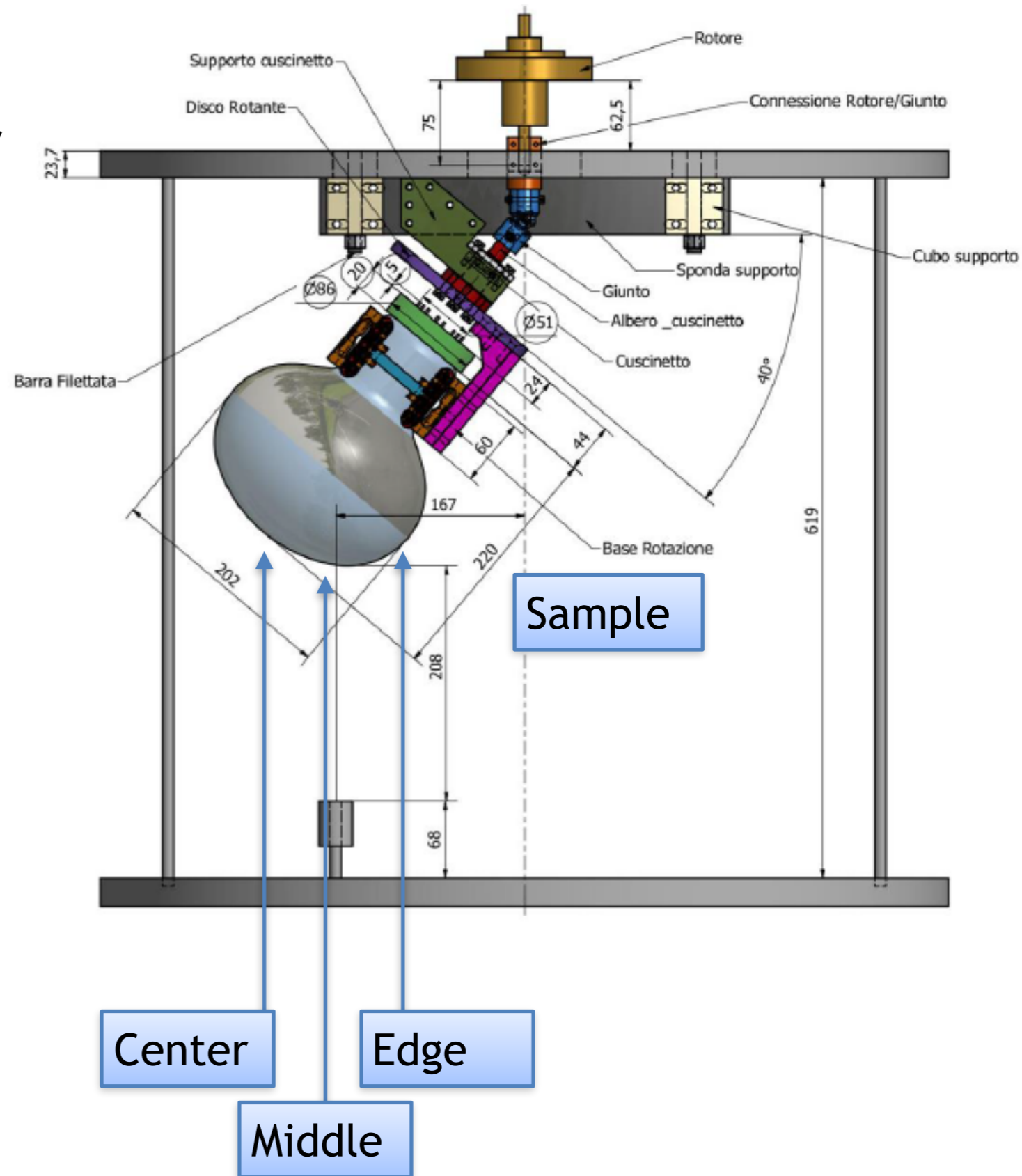
Past, present and future...

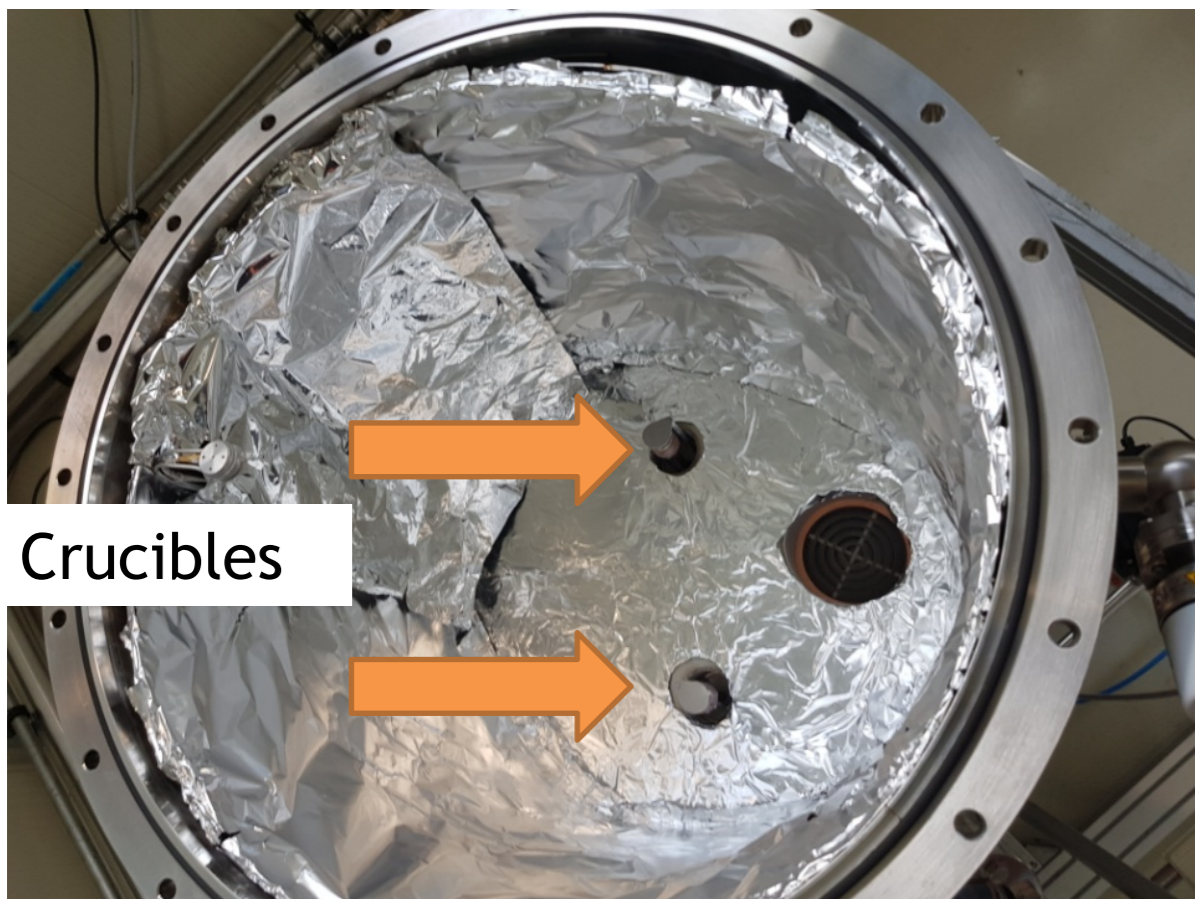
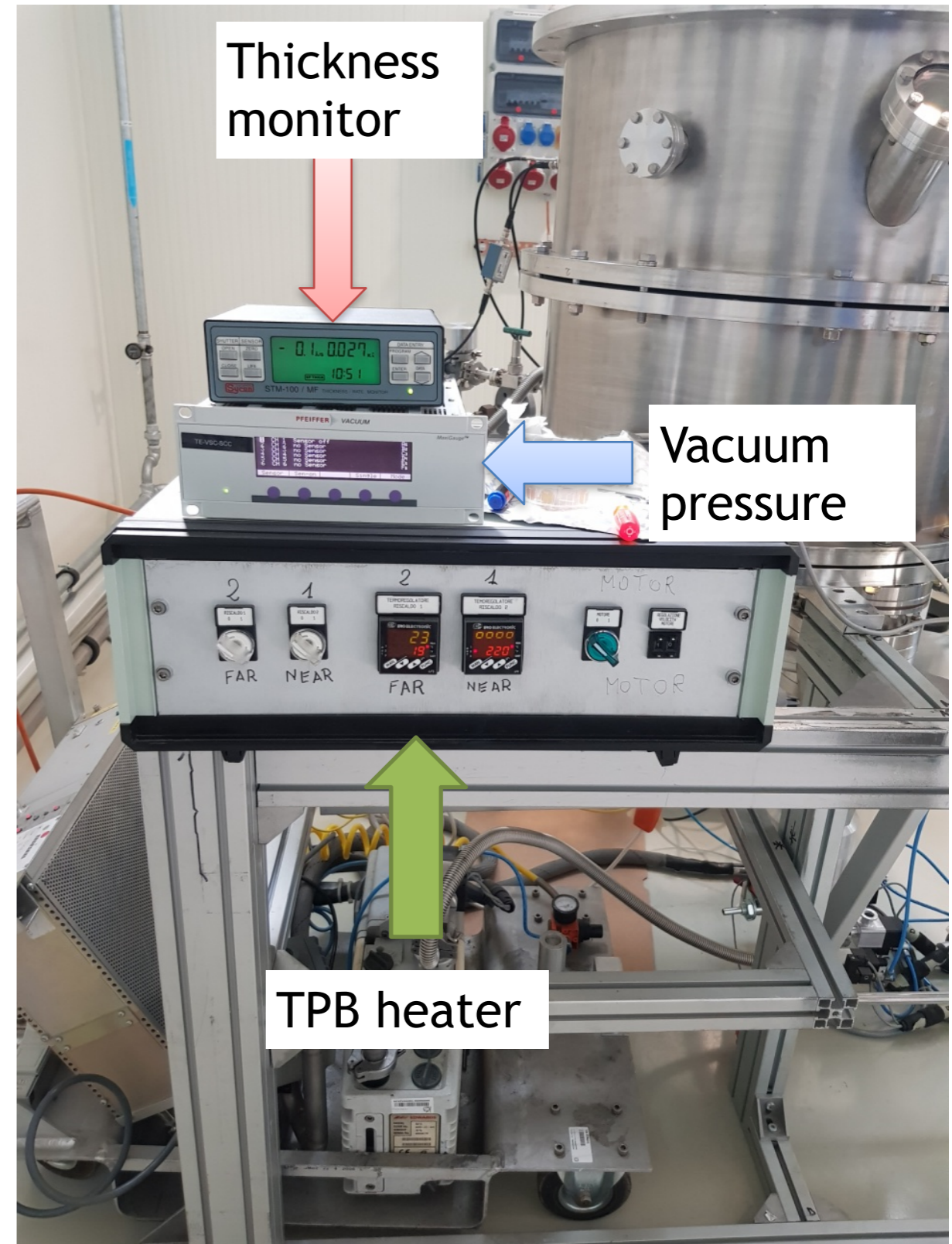
C. Jesús-Valls

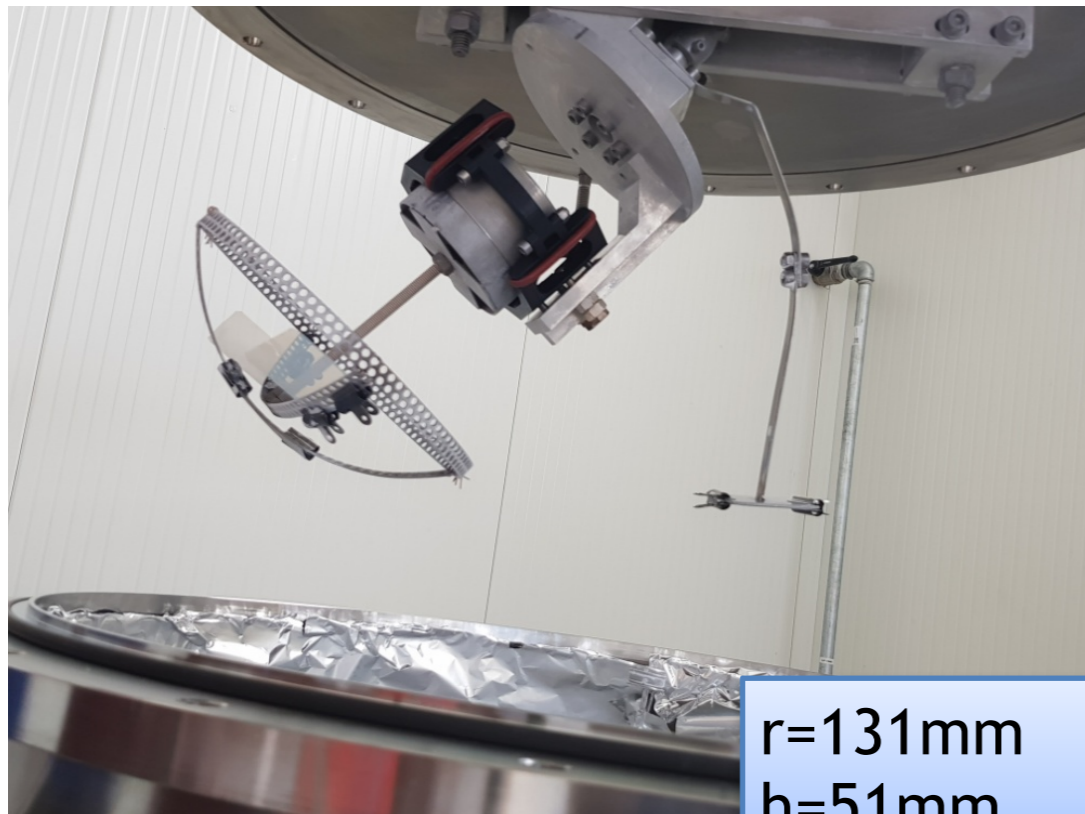
PMT COATING

Coating Facility

- using facility developed for ICARUS experiment
- ICARUS coated 360 PMTs with it
- position of the PMT optimized for uniform coating
- full description of the system: <https://arxiv.org/pdf/1807.07123.pdf>
- system available at CERN for everyone
- great support from CERN Thin Film facility



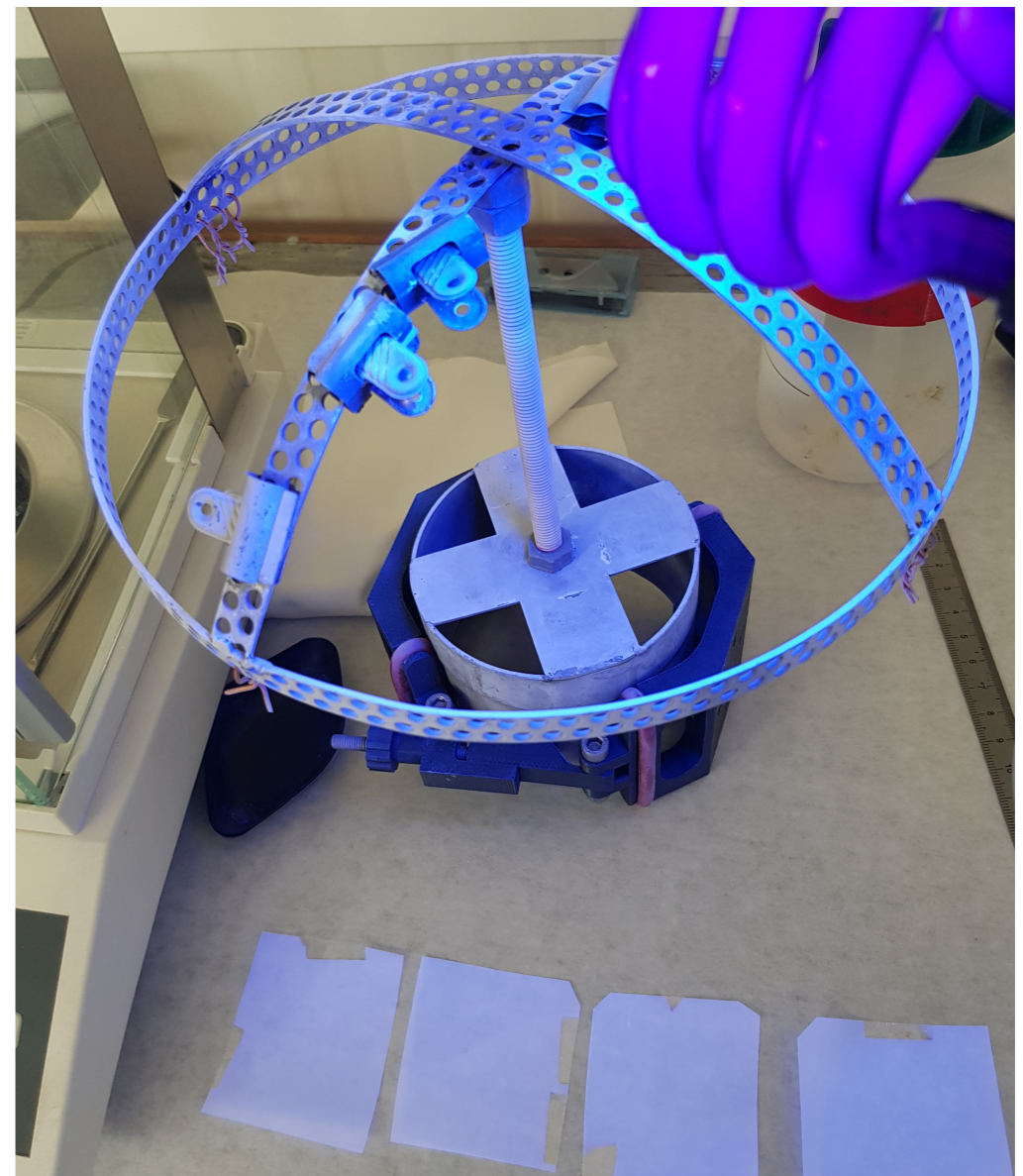




$r=131\text{mm}$
 $h=51\text{mm}$

- test samples also useful for thickness calibration
- all PMT have been coated with more than $0.2\text{mg}/\text{cm}^2$ in the middle.
- maximum difference in thickness is expected to be $<20\%$ from the extreme edge to the middle (middle-center $<5\%$).

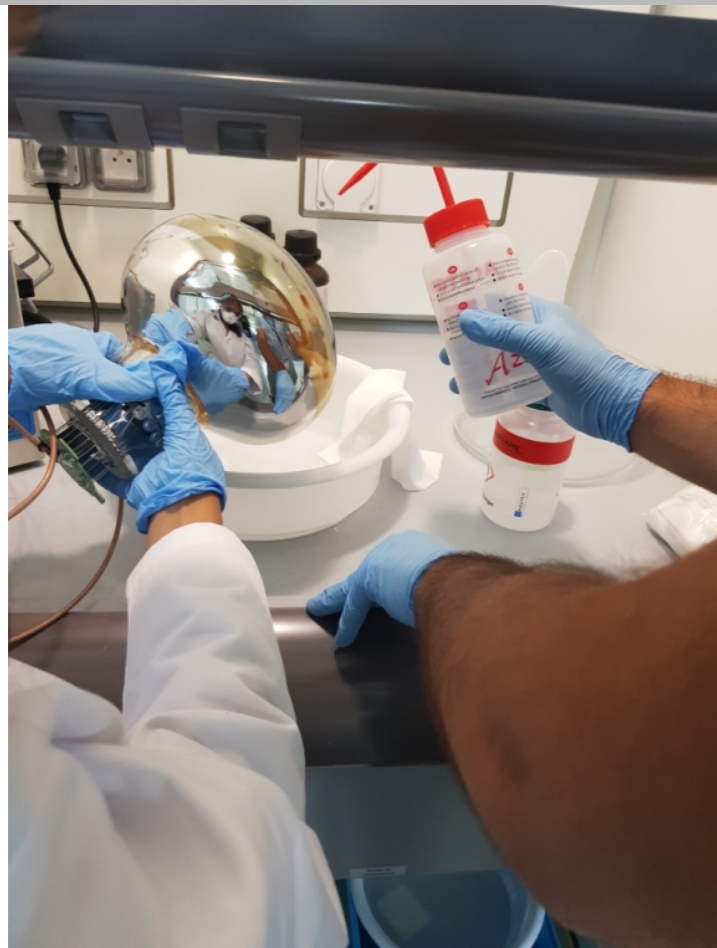
- test sample coating to test the system after transport and to learn
- spend 1.5 weeks on learning process
- easy but annoying job





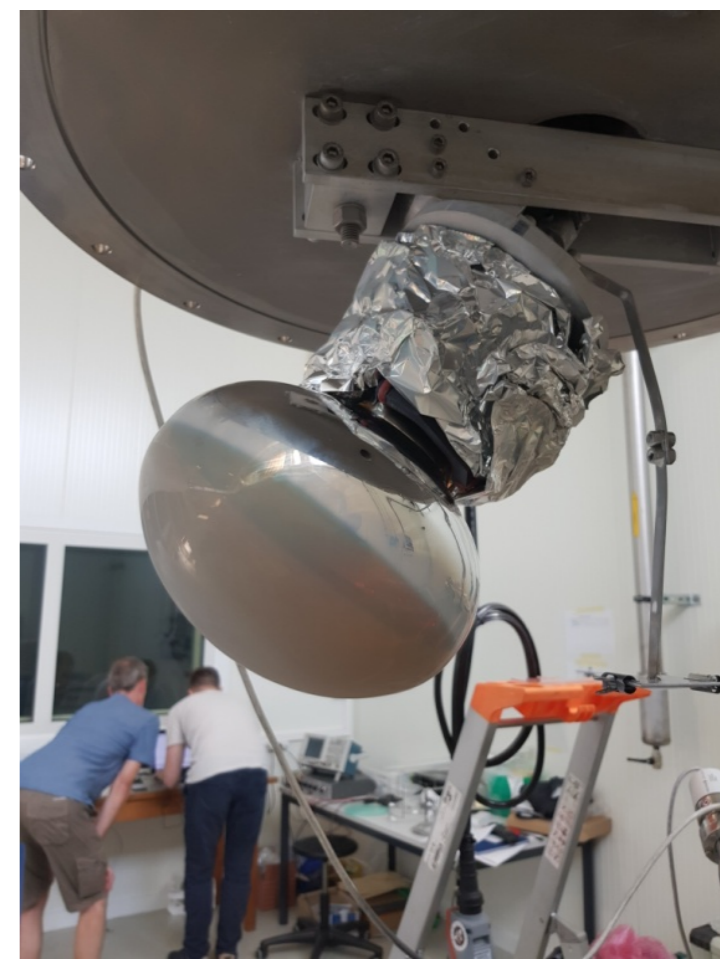
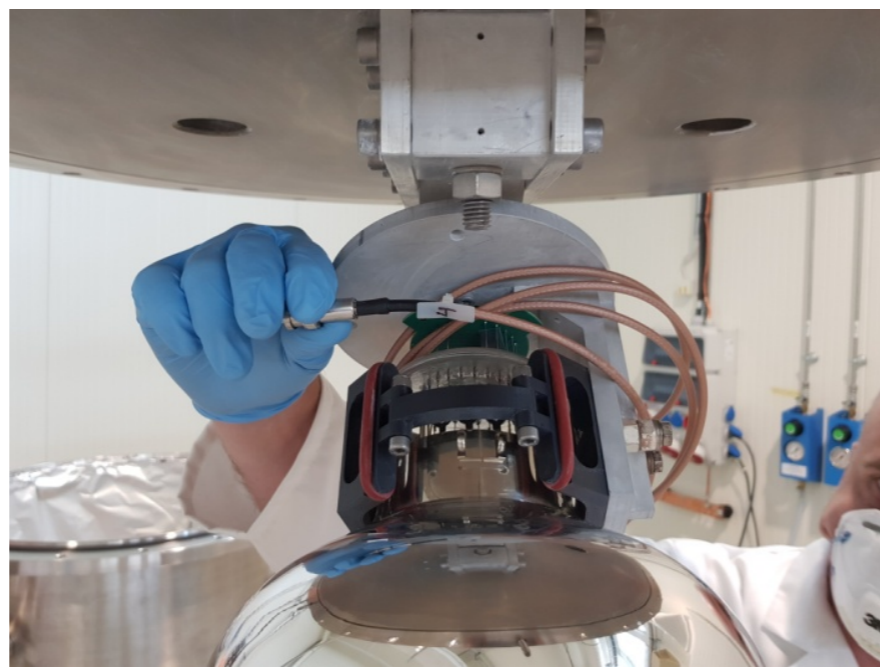
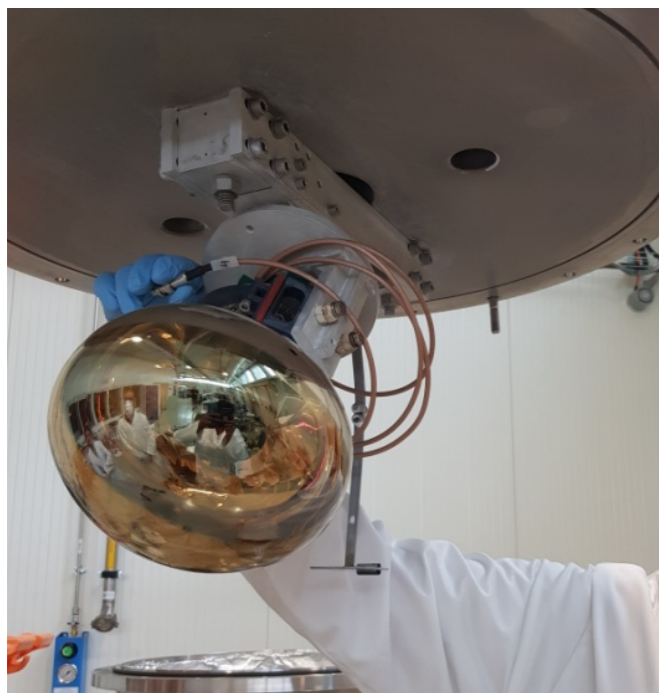
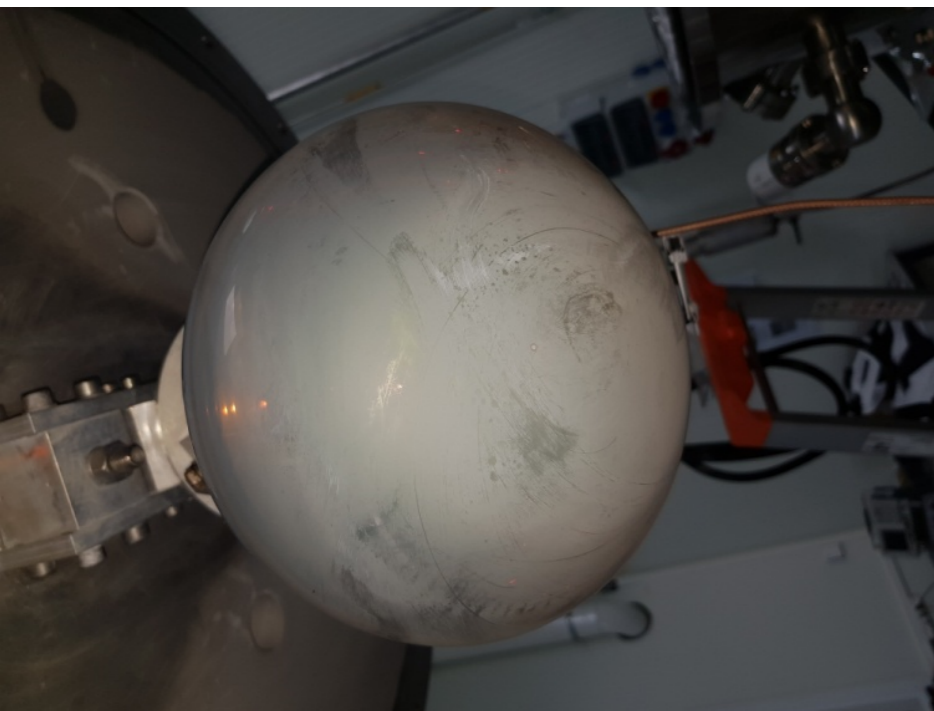
Procedure

- visual inspection of box => all fine
- PMT was left in box, HV cable was taken out, the box was covered with black sheets
- Dark count test at 1200 V => all fine
- Photo from PMT outside box to have reference for PMT orientation
- dismanteling of the PMT+support
- cleaning of the PMT surface with acetone and isopropanol + drying the surface => **crucial for good coating quality**



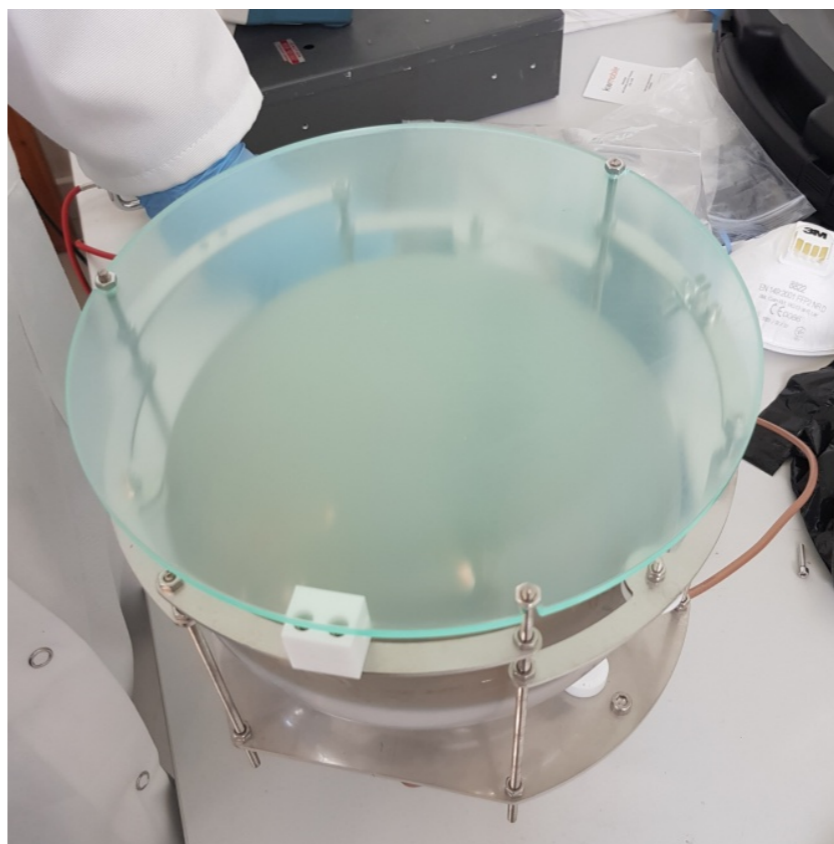
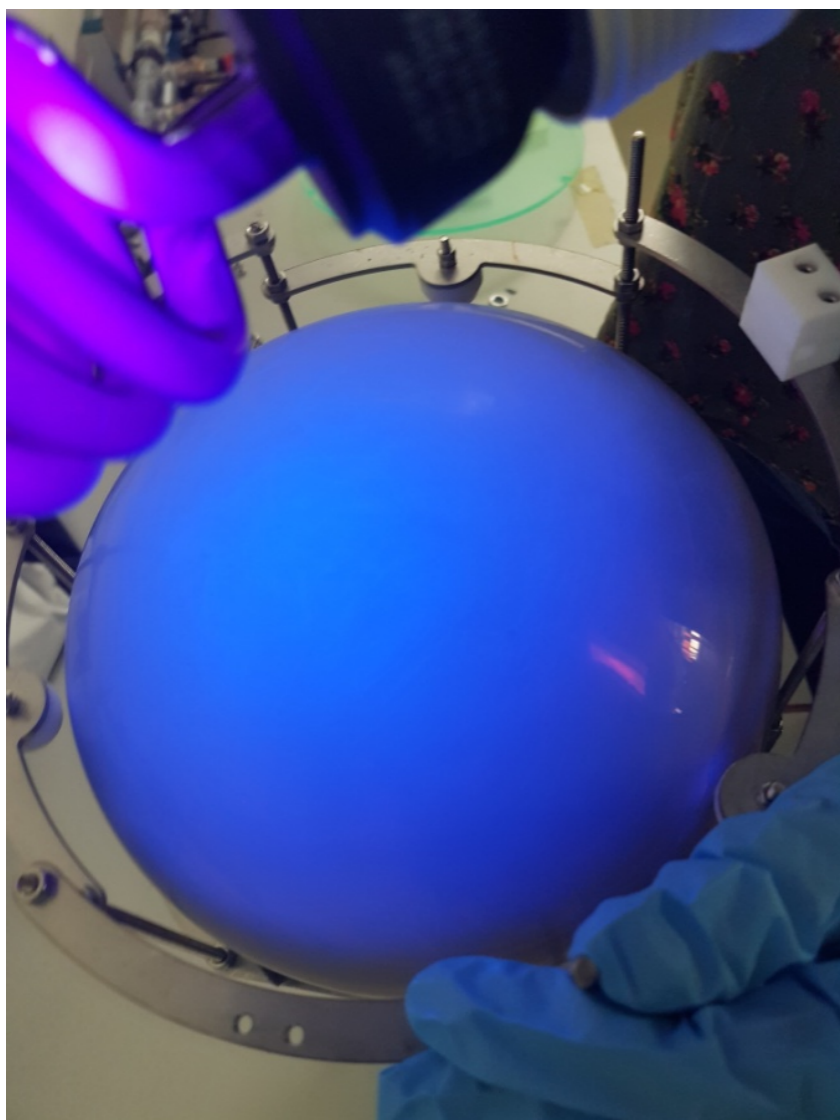
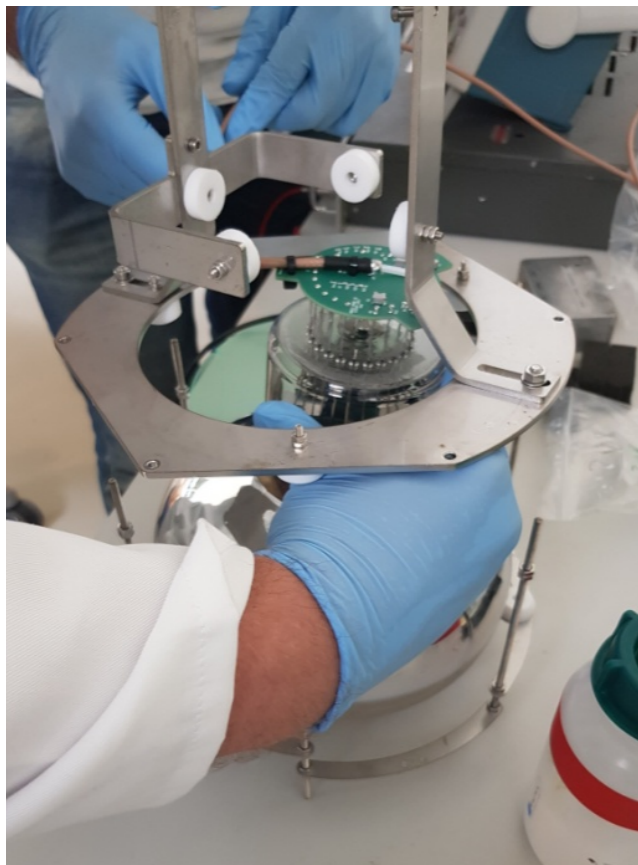
Procedure

- Fixing the PMT in the vessel
- Placing carefully the cable so that the PMT can rotate
- cover cable and base with alu foil
- pumping down to below $3 \cdot 10^{-5}$ mbar
- heating up TPB to 220 C monitoring every minute the parameters



Procedure

- Re-assembling of the support structure with PMT
- checking coating with UV lamp + photo
- adding acrylic plate to protect TPB
- place in box with silikat gel to keep dry => 35 of 40
- Dark count test after coating



Q_{eff} Measurement

- 40 PMTs now stored in EHN1
- 4 PMTs will be brought back to Meyrin for Q_{eff} measurement
- producing missing part currently => after 7th of September
- only >200 nm and quantitative not qualitative
- trying to contact ICARUS group to explore possibility to measure at 128 nm
- not clear if setup still exists

Absolute QE measurements

from Thin Film & Glass lab

<http://cdsweb.cern.ch/record/1164394/files/PH-EP-Tech-Note-2009-001.pdf>

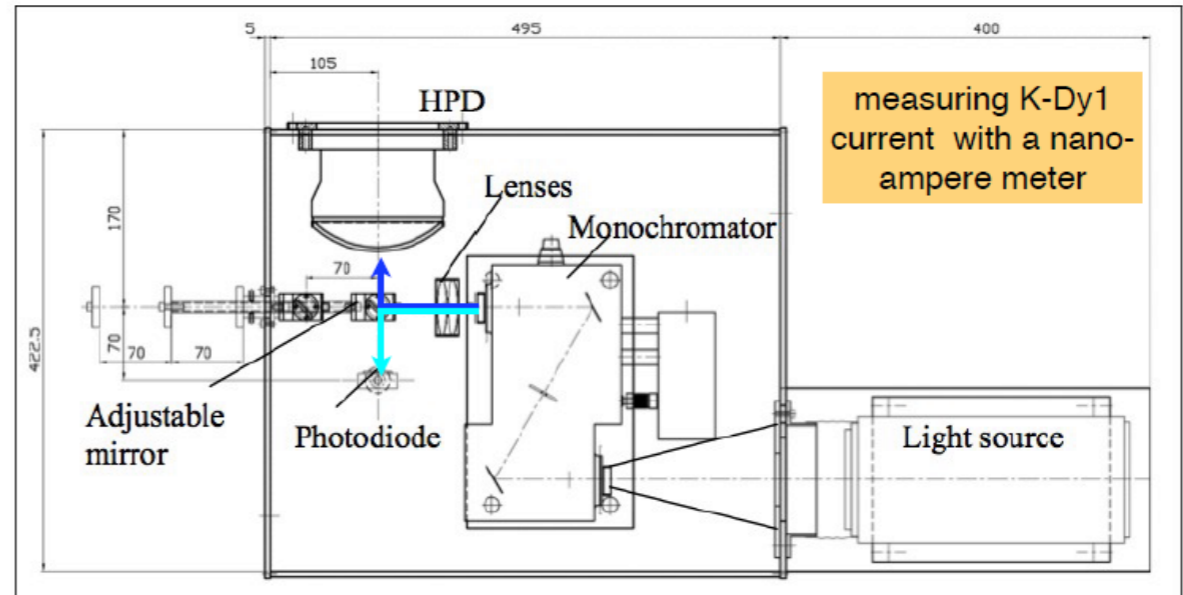
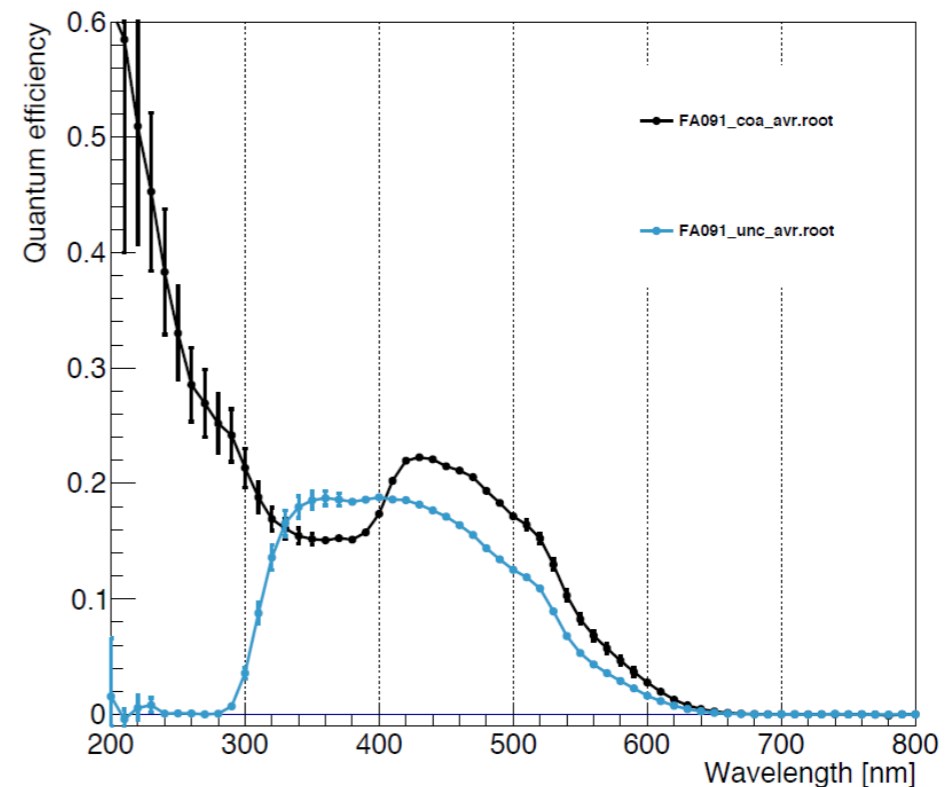


Figure 15: Overview of the mechanical setup of the measurement system. Dimensions in mm.

$$\epsilon_{PMT} = \frac{I_{PMT}}{I_{PD}} \epsilon_{PD}$$

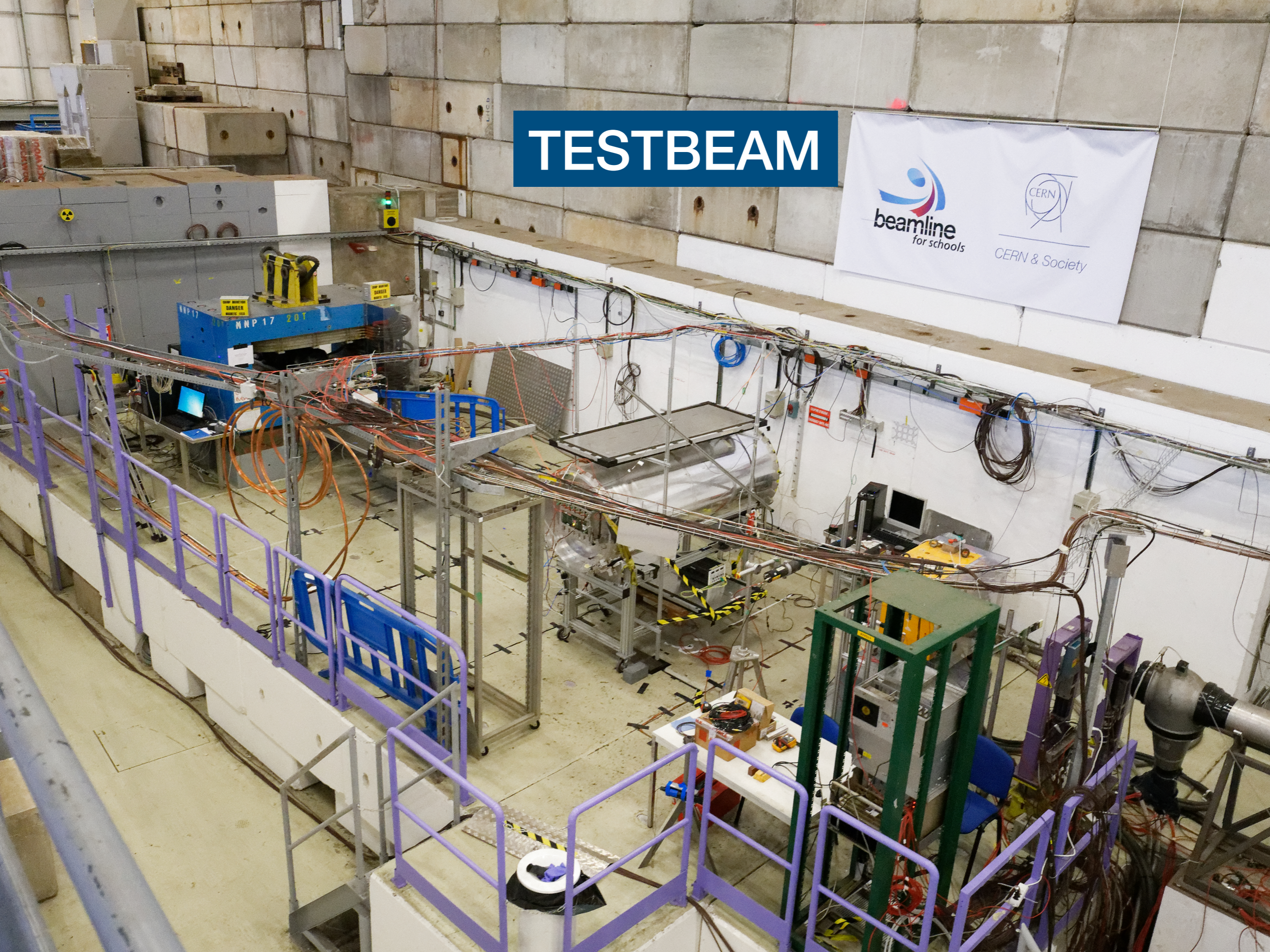
Sosuke's slide
@ArDM meeting

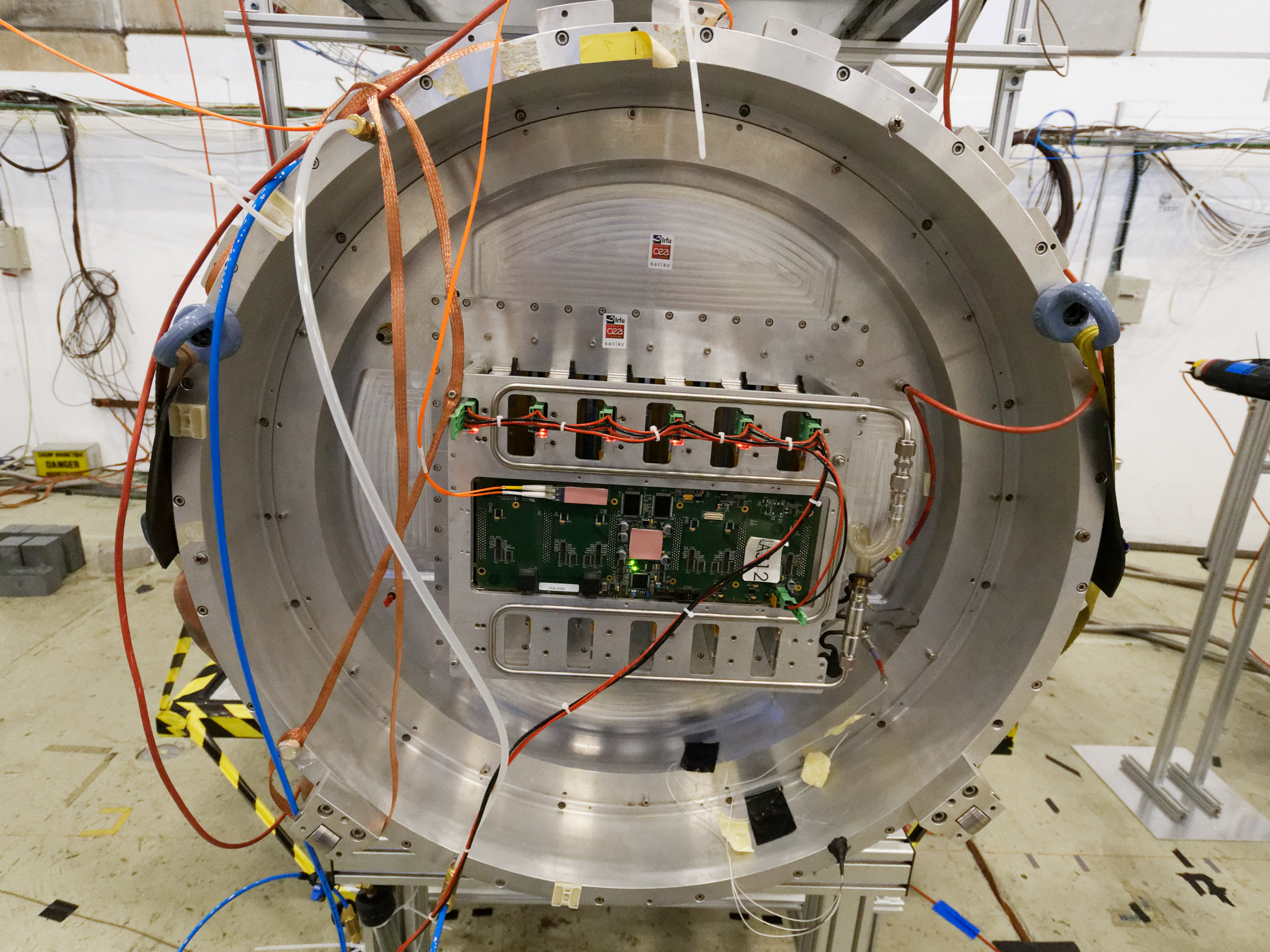


Summary + Outlook

- 40 PMTs were coated with ICARUS facility at CERN
- 2 PMTs/day coated (ICARUS w/o base 4 PMTs/day)
- Learned a lot:
 - Drying process after cleaning crucial
 - PMTs ideally stored after coating in sealed Mylar/Alu bag
- Visual inspection with UV lamp fine
- Open question: Surface preparation => ArDM seemed to have no preferences but ICARUS sand blasted all 360 PMTs
- Trying to organize Q_{eff} measurement at 128 nm
- PMTs at EHN1 waiting for installation

TESTBEAM





Test Beam data summary

Different runs contain:

- Different particles:
 1. “**Electrons**” = Scintillators * Cherenkov
 2. “**Pions**” (+ **muons**) = Scintillators * not_protons * not_electrons
 3. “**Protons**” (+ kaons) = S1_delayed * S2 (delay ~ proton TOF between S1 and S2)
 4. “**CR**” from icarus trigger module but only out of SPILL
 5. In addition there were **⁵⁵Fe clusters**, from a radioactive source.
- Different **momentum** for the beam.
- Different **HV** in MM.
- **Temperature** dependence.
- **Pressure** dependence.
- 6 different **FECs**, with 4 **ADC** each one.
- Different **threshold** conditions.
- Different **shaping** time.

**Preliminary data shows high quality of data
Many possible studies**

Track Selection Algorithm

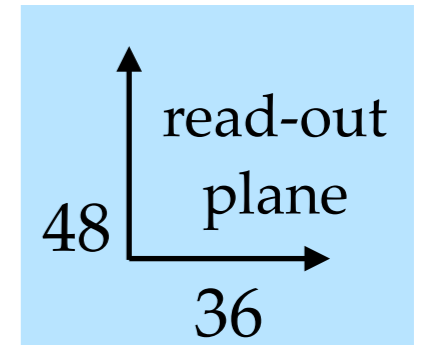
Selection Algorithm (Developed by S.Suvurov)

Common structure for the analysis:

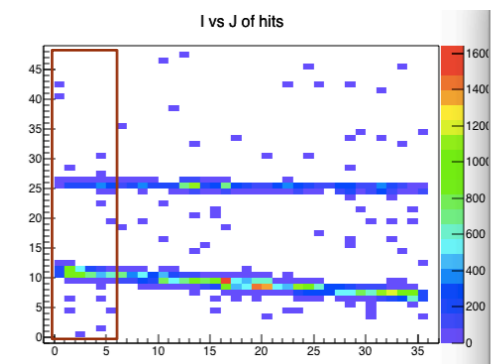
- Loop over channels
- Search of the maxADC in each channel.
- Store information.

- A. Store maxADC for each channel.
- B. Variation: Store ADC integrated in a window around the maxADC.
- C. Store time of max ADC.

Selection Input



1728 pads
7x10 mm²



Selection (COSMIC / BEAM)

- Cosmic and Beam selection works in the same way (just rotating definitions).
- In one margin (Top for Cosmics, right for Beam) a clustering algorithm classify the pads in sets of pads (clusters) with common position and time requirements.
- A linear fit is applied to the largest cluster.
- If the quality of the fit is good all pads around the fitted straight line are selected.

Selection Output

Track Selection Known Problems

Problem!

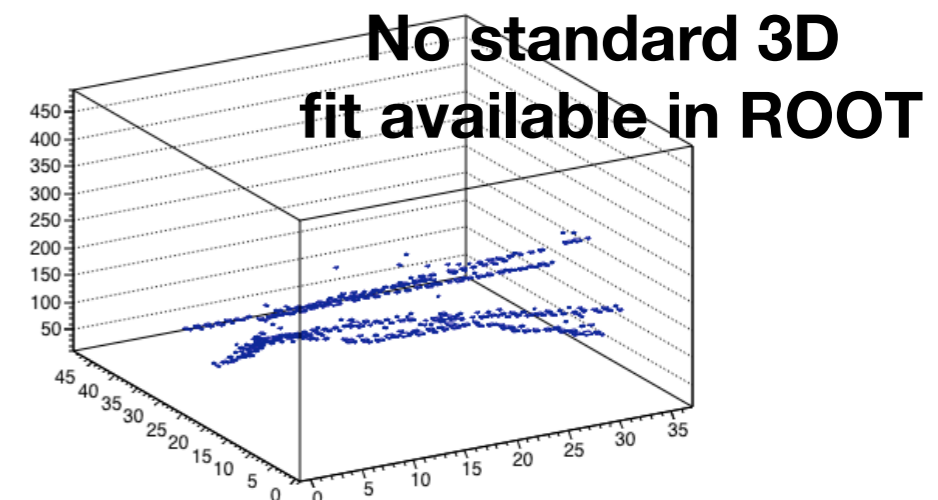
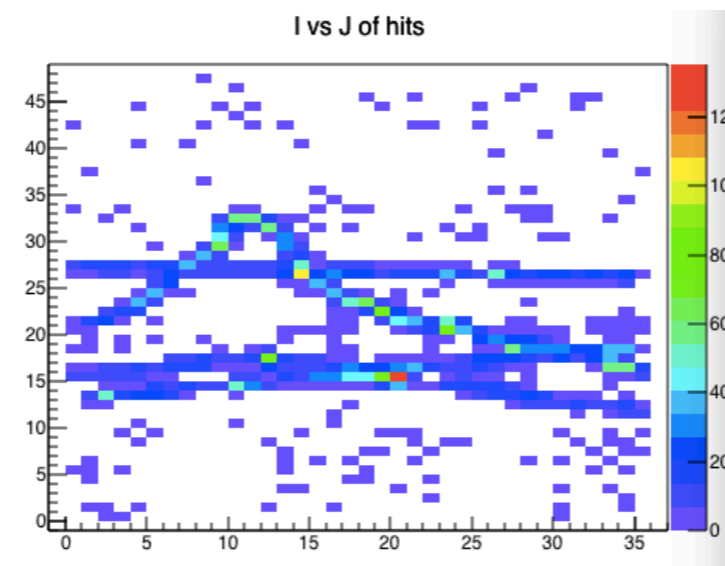
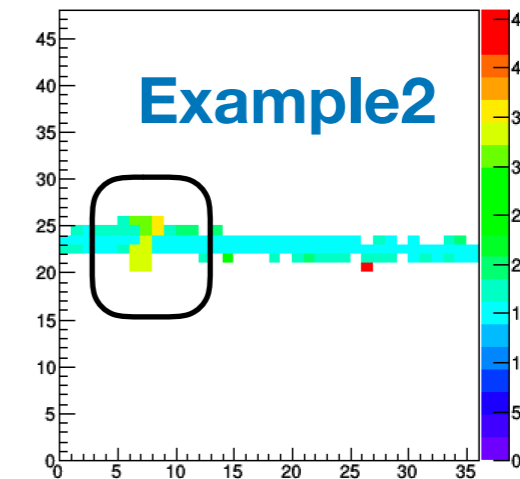
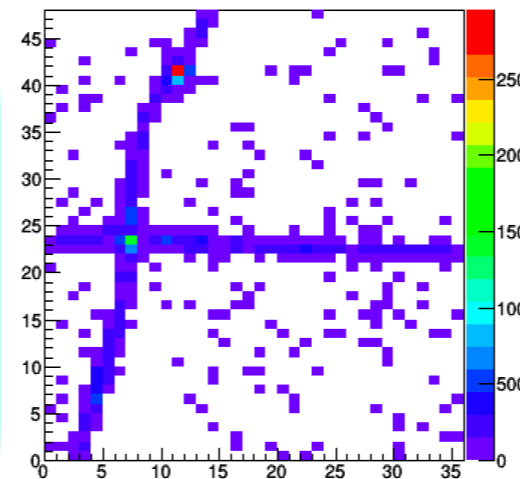
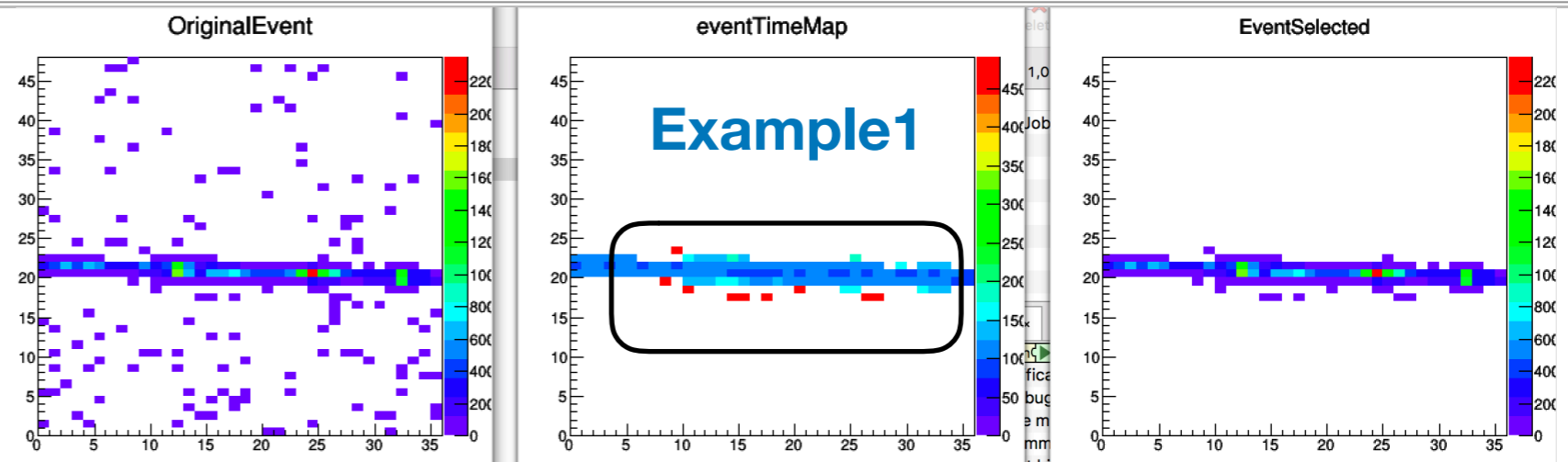
Since we have 3D tracks selected from a 2D developed algorithm we are selecting “wrong” noise pads.

No easy 2D solution

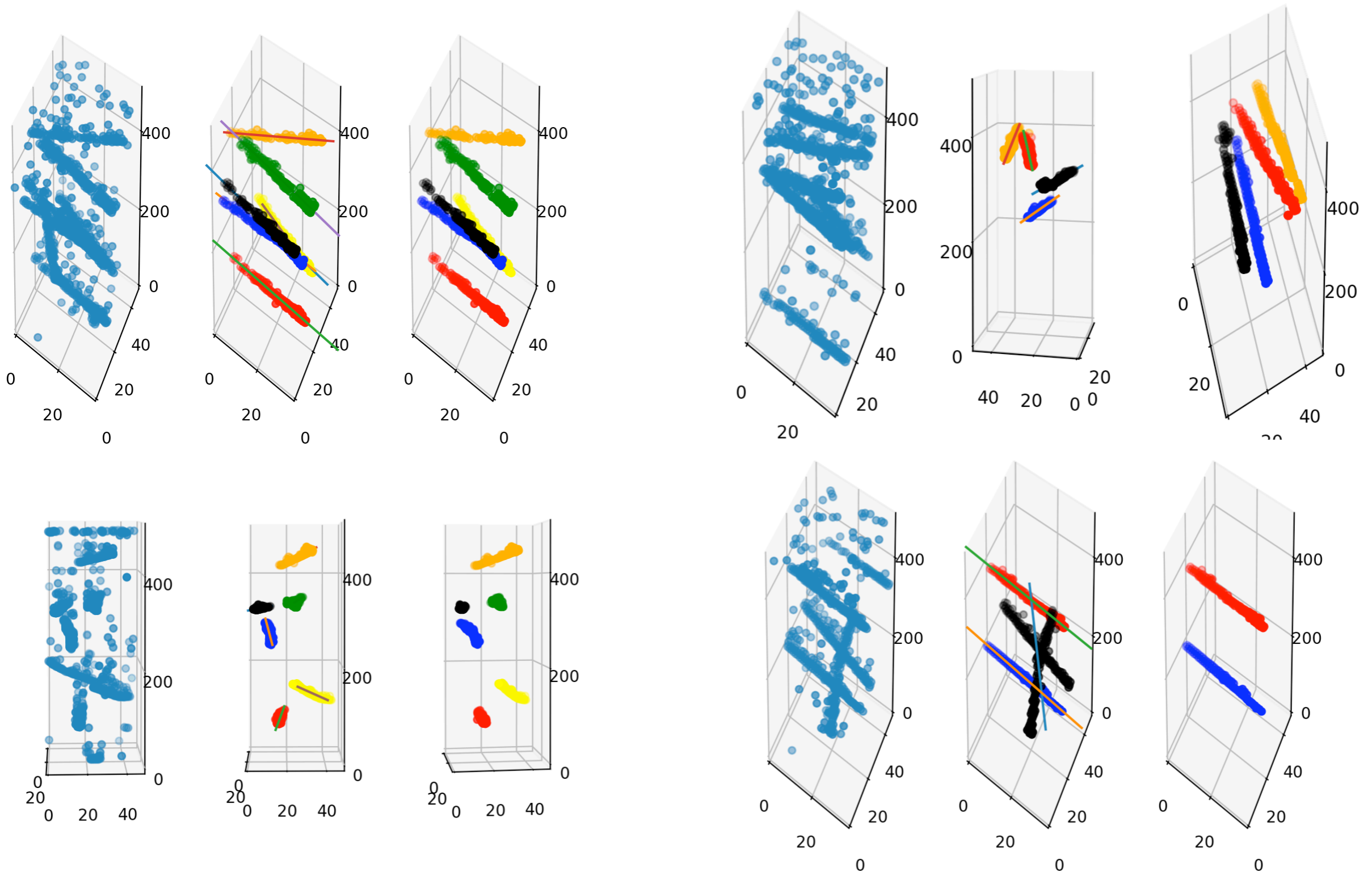
Cut in time after selection can be applied to solve this (done in PRF studies) However does not work for COSMICS that are not perpendicular to time axis!!

Solution? Move to 3D

Exploit full power of 3D information to implement better selection.

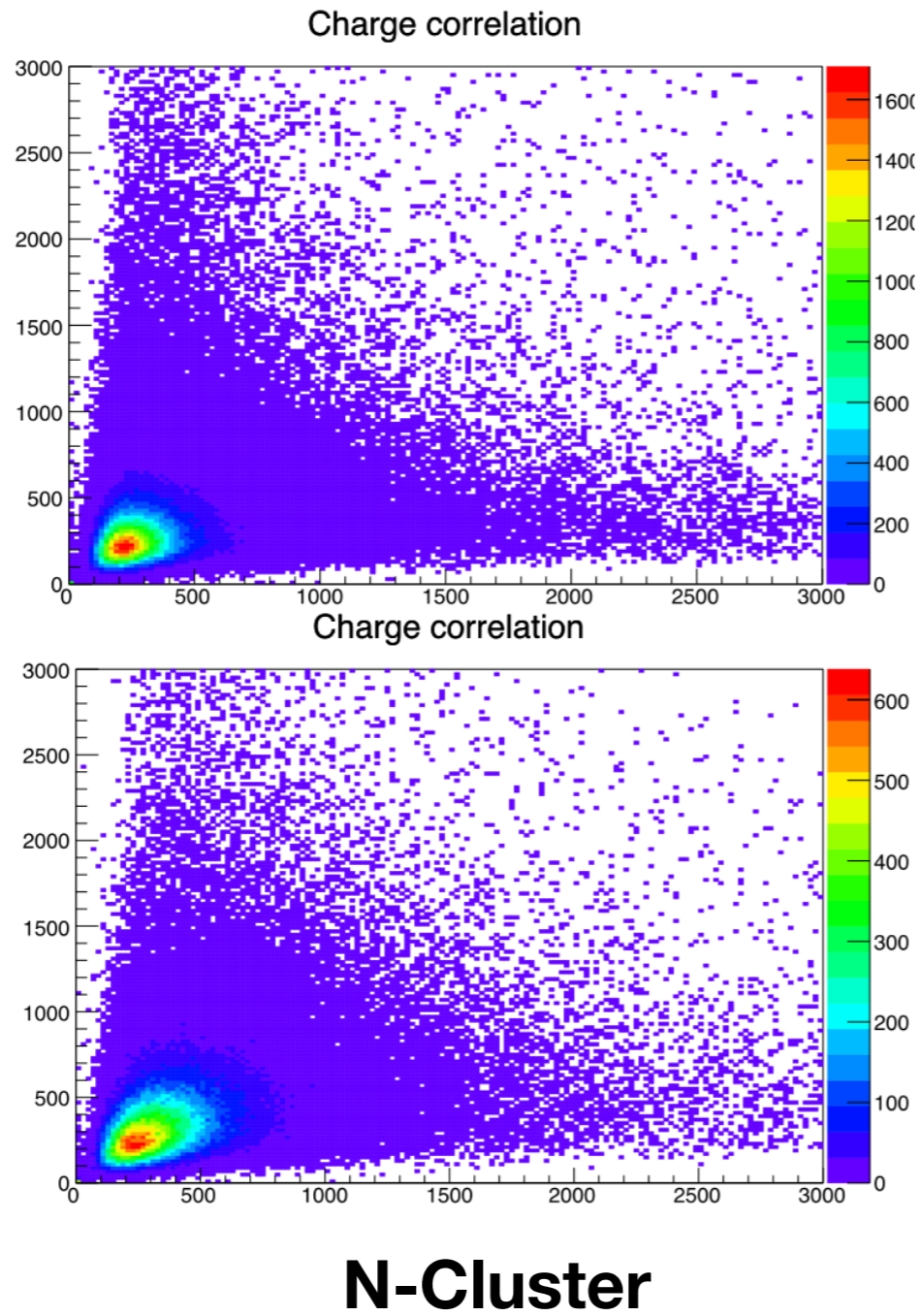


3D selection



ANALYSIS STATUS

N+1 Cluster

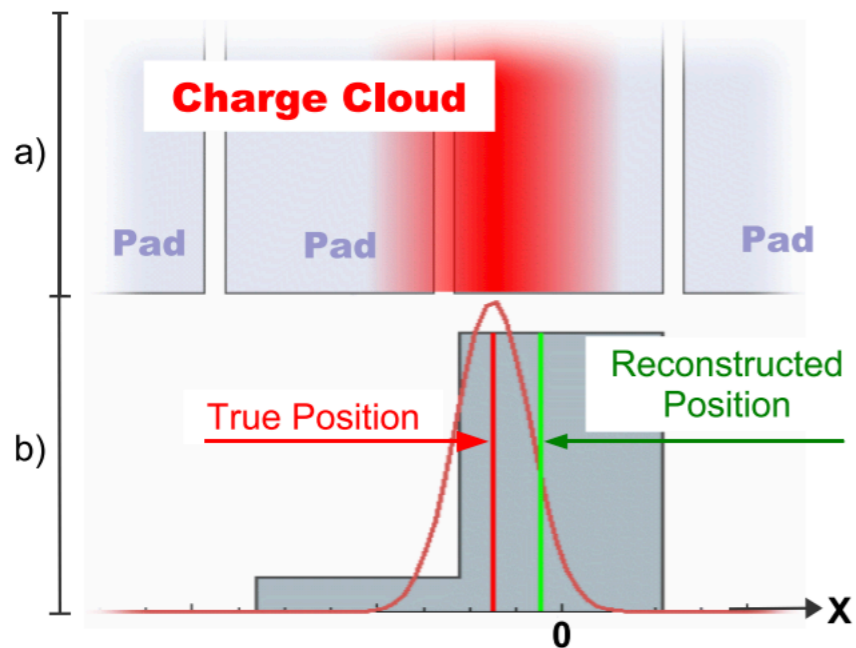


Sample:

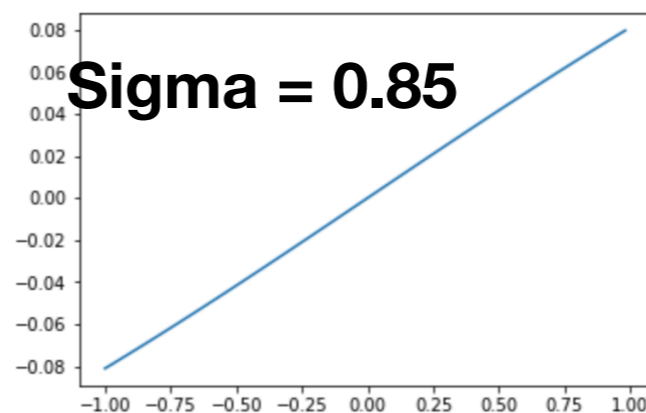
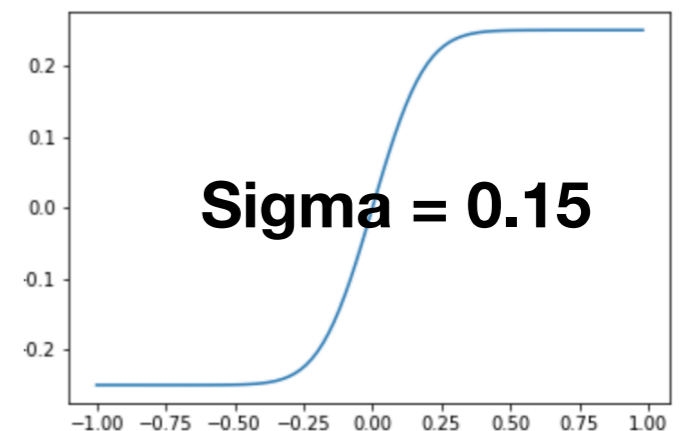
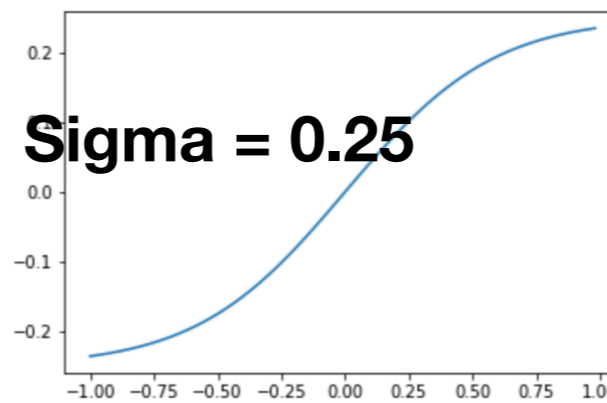
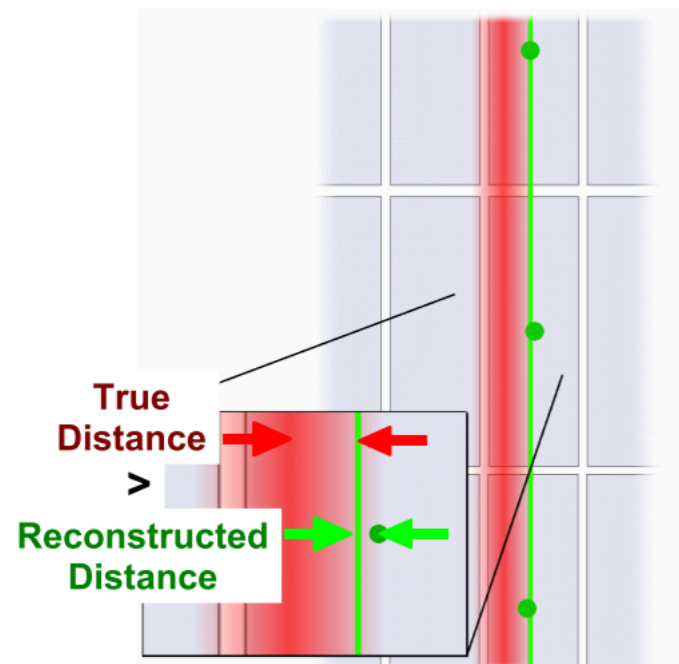
2 GeV/c muon

0.8 GeV/c pion trigger
(pion + electron)

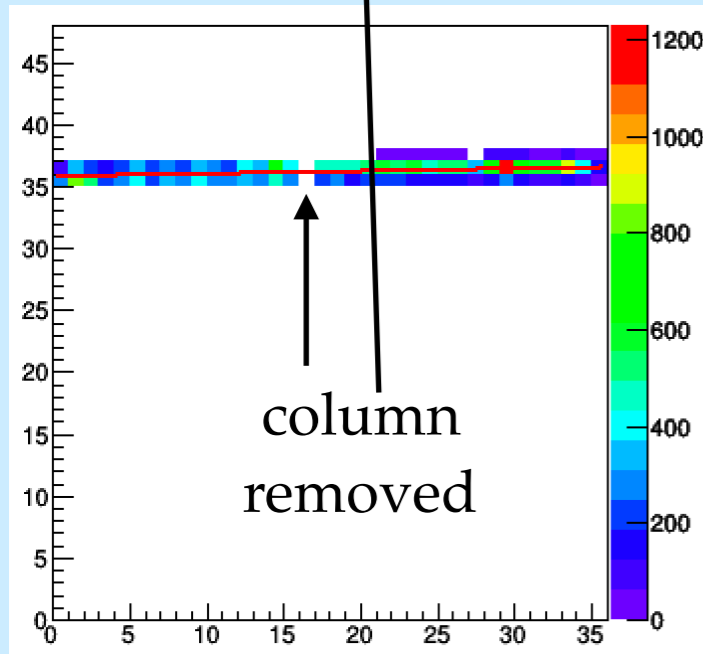
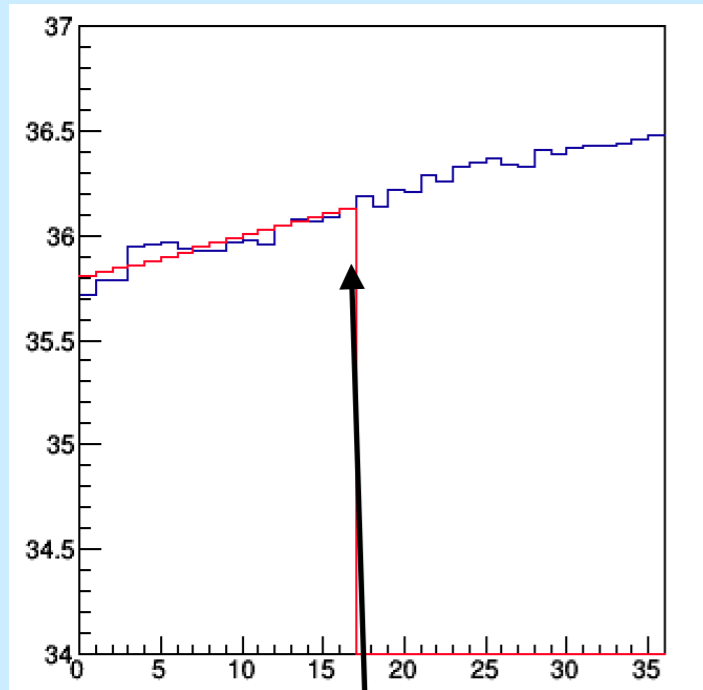
Charge correlation



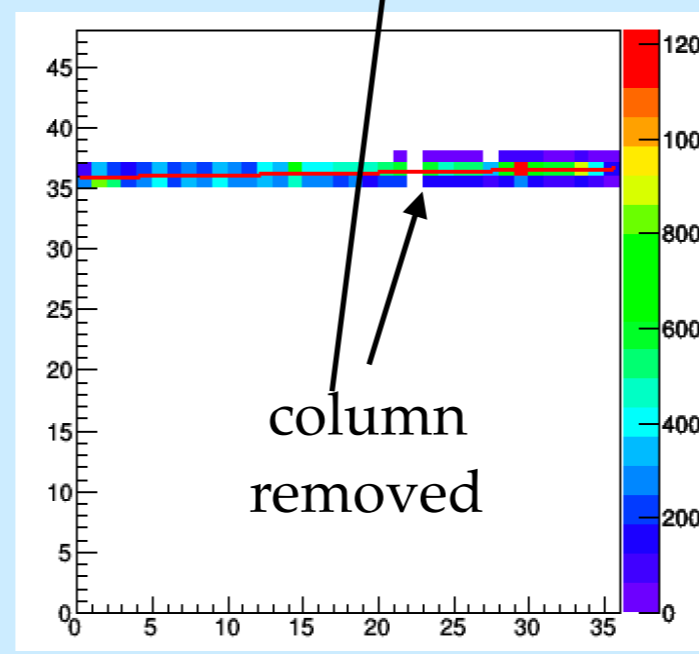
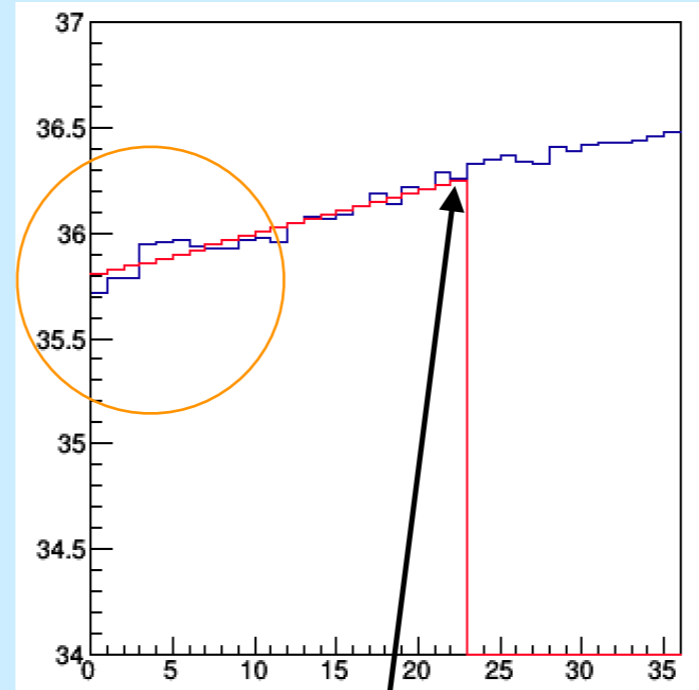
There is source of bias coming from the size of pads... the way to characterize and correct his effect is using the Pad Response Function (PRF)



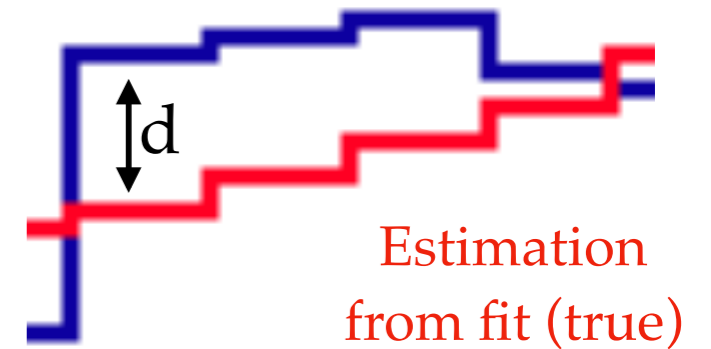
Variables definition



Column Iterations



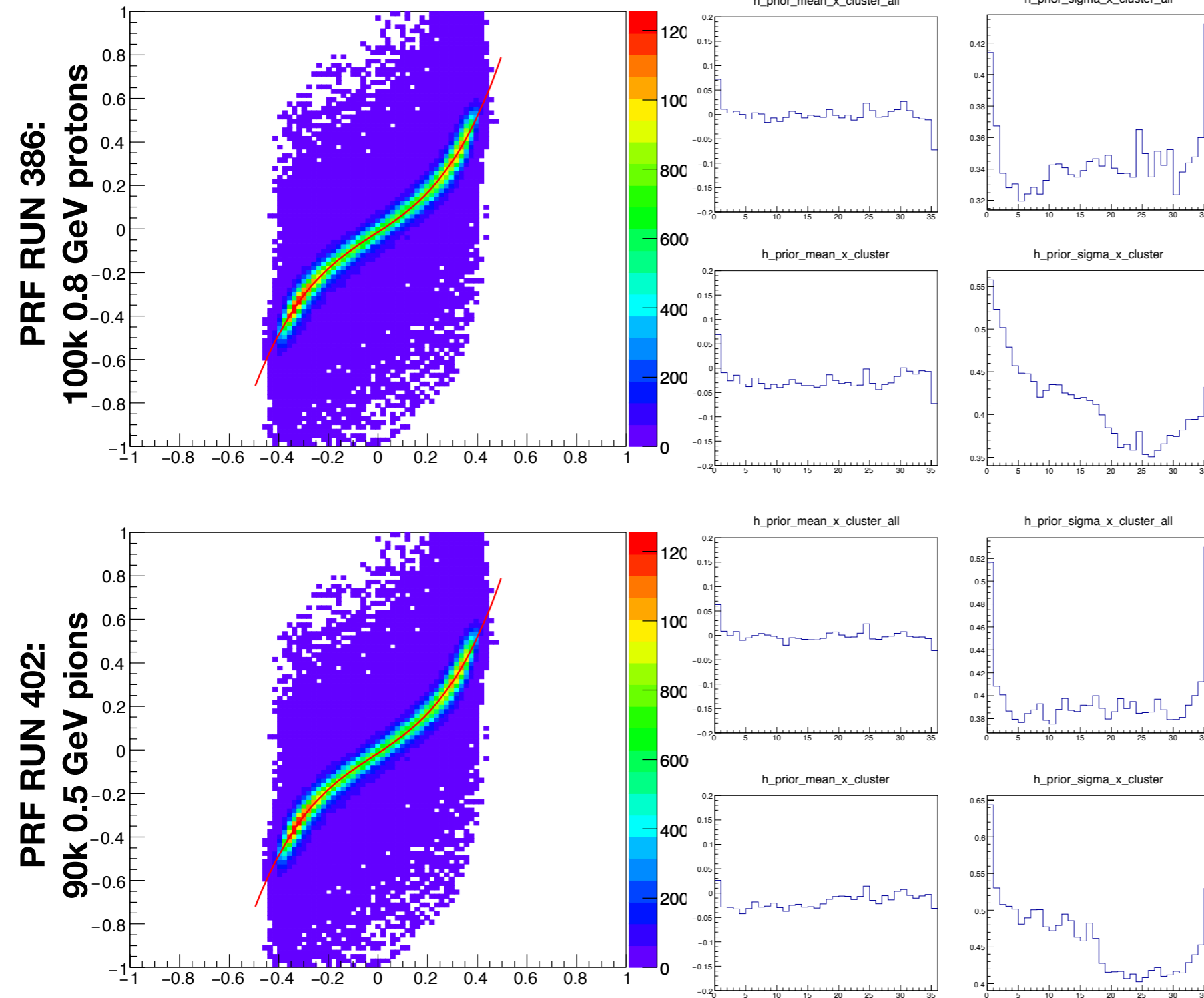
Gravity Center of the column (reconstructed)



Fit can be done in 2 ways (also 2 possible distances):

1. All-1 Columns (d_1)
2. All Columns (d_2)

Pad response function and non-corrected point resolution



NPAD: 236312 Entries

Point resolution fitting all clusters: 0.354401
 Point resolution fitting all-1 clusters: 0.442227
 Final (convined) point resolution: 0.395886

1PAD: 16541 Entries

Point resolution fitting all clusters: 0
 Point resolution fitting all-1 clusters: 0
 Final (convined) point resolution: 0

2PAD: 204600 Entries

Point resolution fitting all clusters: 0.34006
 Point resolution fitting all-1 clusters: 0.424929
 Final (convined) point resolution: 0.380133

3PAD: 15171 Entries

Point resolution fitting all clusters: 0.352337
 Point resolution fitting all-1 clusters: 0.418354
 Final (convined) point resolution: 0.383929

NPAD: 132536 Entries

Point resolution fitting all clusters: 0.425528
 Point resolution fitting all-1 clusters: 0.528037
 Final (convined) point resolution: 0.47402

1PAD: 10429 Entries

Point resolution fitting all clusters: 0
 Point resolution fitting all-1 clusters: 0
 Final (convined) point resolution: 0

2PAD: 103721 Entries

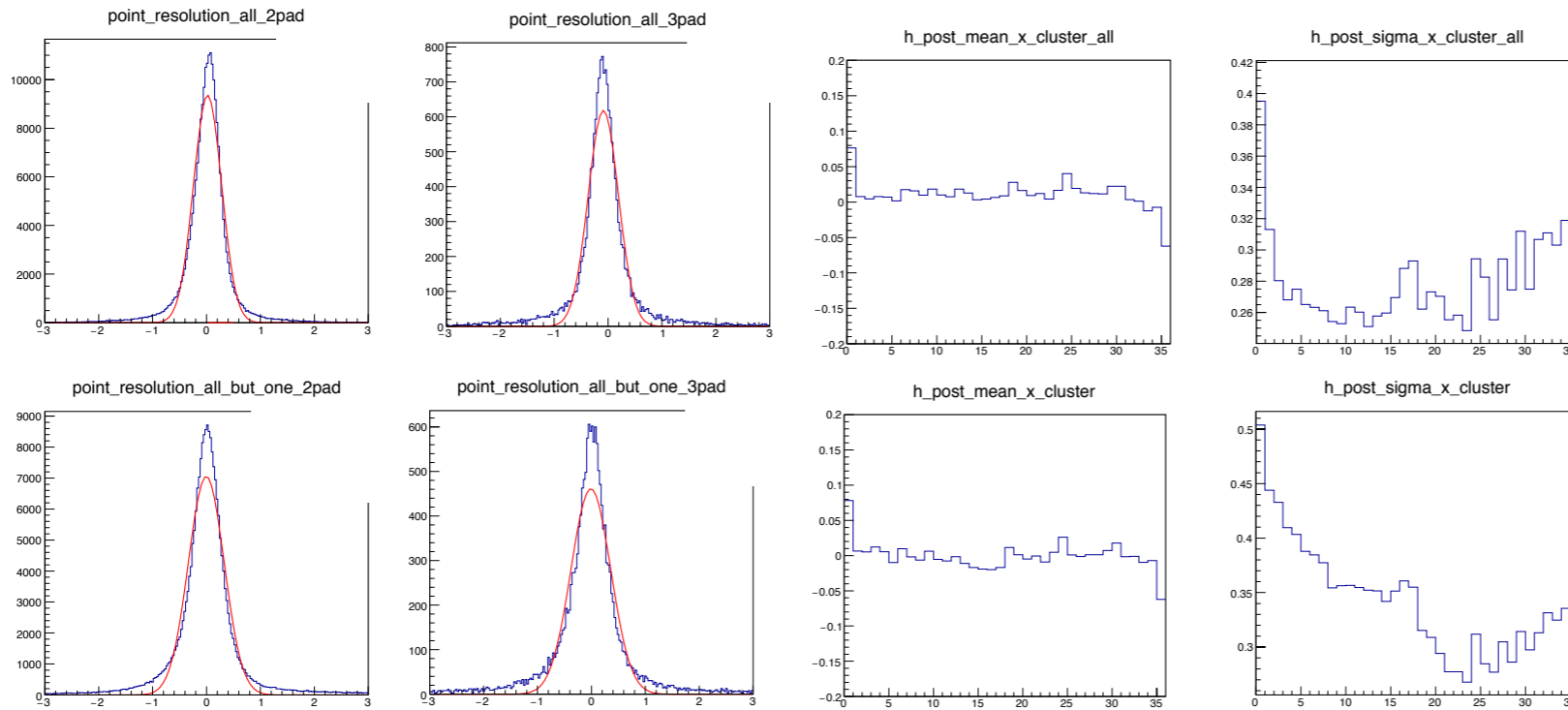
Point resolution fitting all clusters: 0.405356
 Point resolution fitting all-1 clusters: 0.512995
 Final (convined) point resolution: 0.456011

3PAD: 18386 Entries

Point resolution fitting all clusters: 0.403901
 Point resolution fitting all-1 clusters: 0.48147
 Final (convined) point resolution: 0.440984

corrected point resolution

**PRF RUN 386:
100k 0.8 GeV protons**



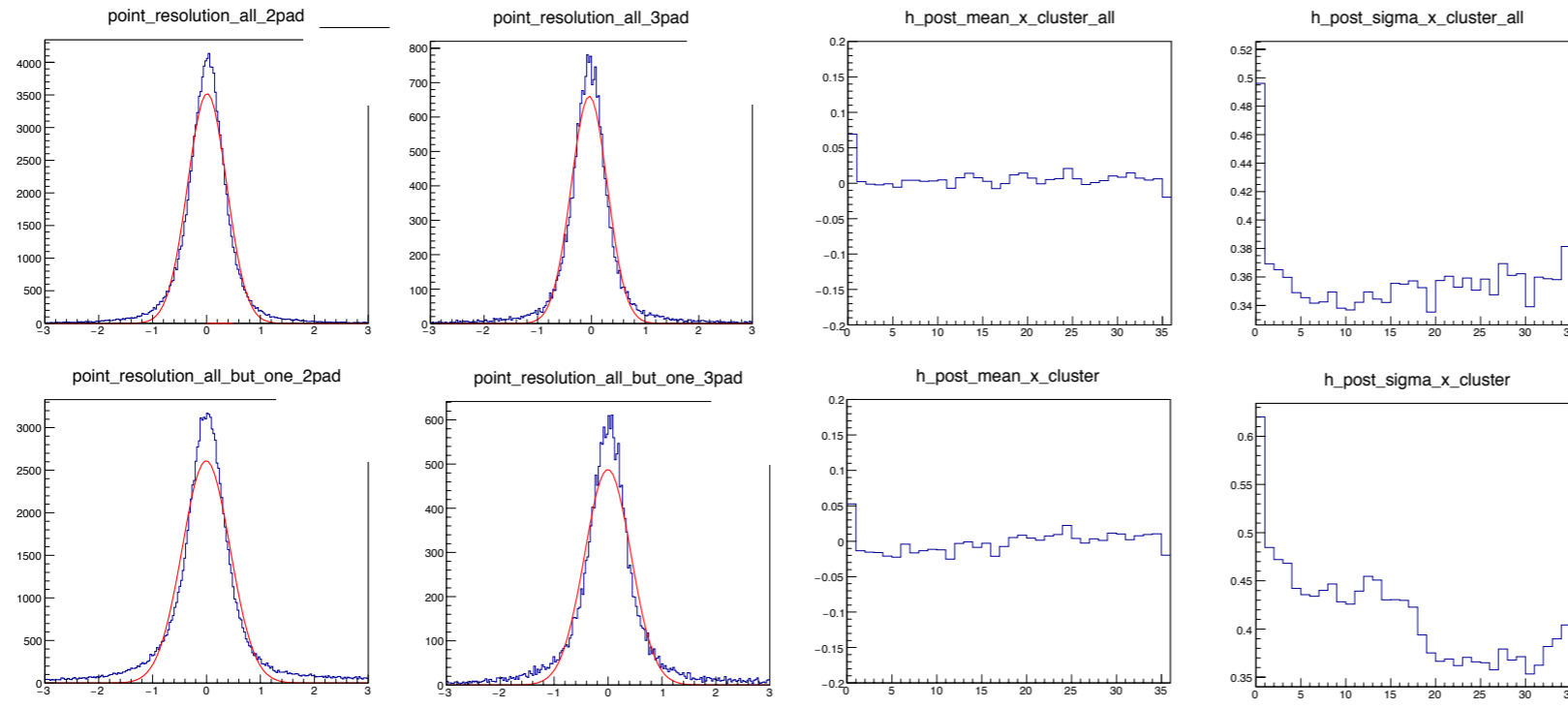
NPAD: 261663 Entries
 Point resolution fitting all clusters: 0.277438
 Point resolution fitting all-1 clusters: 0.34883
 Final (combined) point resolution: **0.311093**

2PAD: 226554 Entries
 Point resolution fitting all clusters: 0.261646
 Point resolution fitting all-1 clusters: 0.331938
 Final (combined) point resolution: 0.294704

3PAD: 17011 Entries
 Point resolution fitting all clusters: 0.28182
 Point resolution fitting all-1 clusters: 0.369522
 Final (combined) point resolution: 0.322705

The improvement after using PRF information is: **21.4186%**

**PRF RUN 402:
90k 0.5 GeV pions**



NPAD: 146711 Entries
 Point resolution fitting all clusters: 0.376979
 Point resolution fitting all-1 clusters: 0.461613
 Final (combined) point resolution: **0.417155**

2PAD: 114867 Entries
 Point resolution fitting all clusters: 0.361851
 Point resolution fitting all-1 clusters: 0.444799
 Final (combined) point resolution: 0.401187

3PAD: 20457 Entries
 Point resolution fitting all clusters: 0.338284
 Point resolution fitting all-1 clusters: 0.430047
 Final (combined) point resolution: 0.381416

The improvement after using PRF information is: **11.9962%**

PRF different approach by Sergey

- Started with center of charge
- Perform few iterations $x \rightarrow \text{PRF} \rightarrow x \rightarrow \text{PRF}$
- Look at the PRF function and fit with Gaussian & Lorentzian function

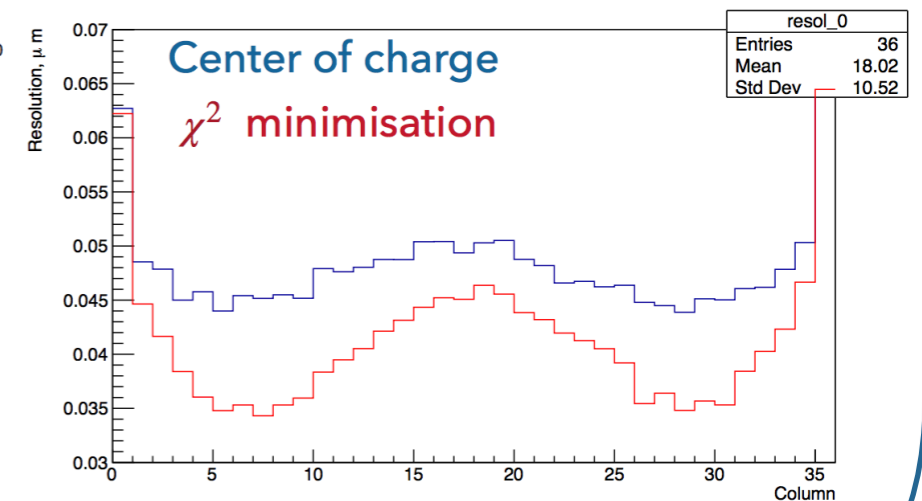
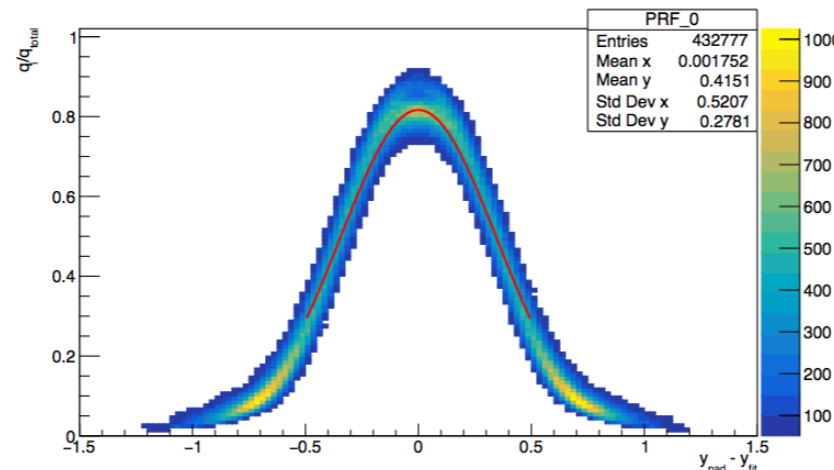
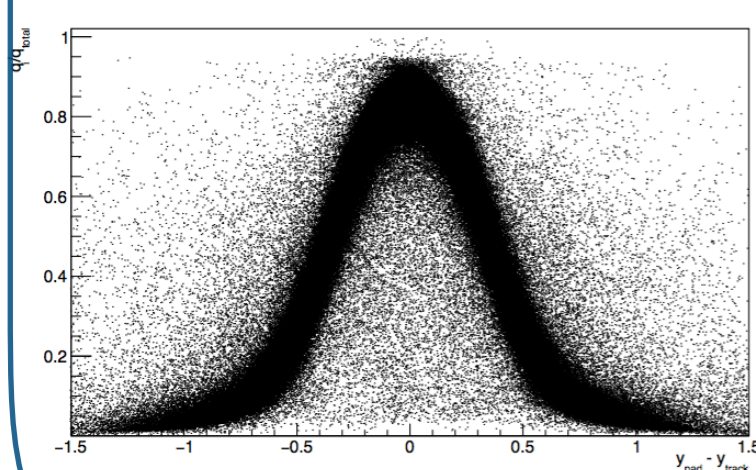
$$\text{PRF}(x, r, w) = \frac{\exp[-4 \ln 2 (1-r)x^2/w^2]}{1+4rx^2/w^2}$$

$$\chi^2 = \sum_{\text{column}} \left[\frac{a - \text{PRF}(y_{\text{track}} - y_{\text{pad}})}{\sigma_a} \right]^2$$

$a = q_i/q_{\text{total}}$
 $\sigma_a = \sqrt{q_i}/q_{\text{total}}$

From S.Suvorov

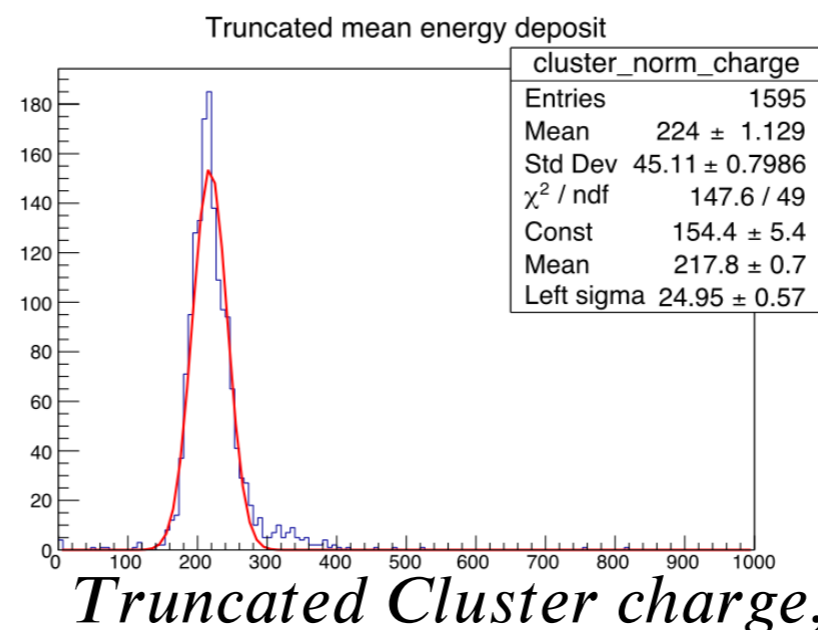
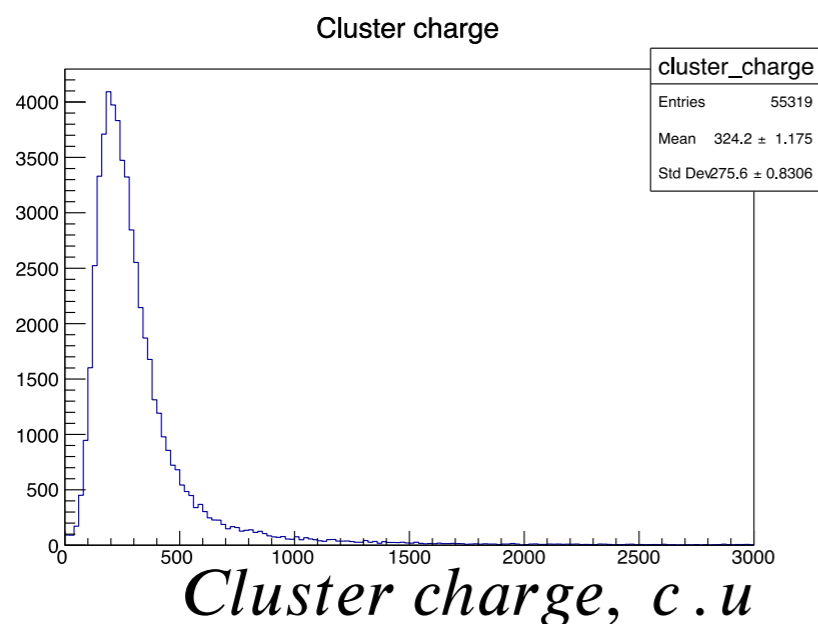
$$\text{PRF} = q_i/q_{\text{cluster}}$$



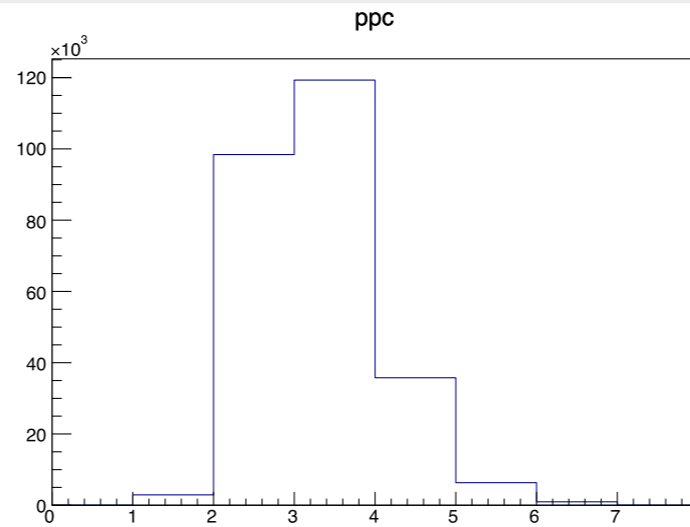
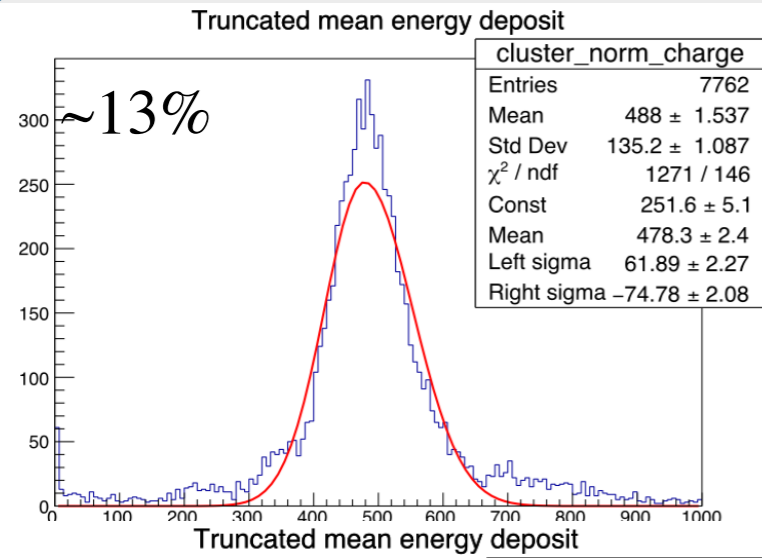
■ The method:

1. Apply the track selection
2. Take the maximum ADC vs. time for each pad
3. Sum up all pads in a column to make a 'cluster'
4. Truncate the most higher energetic clusters
take 70% of clusters sorted by increasing charge
5. Fit \rightarrow profit!

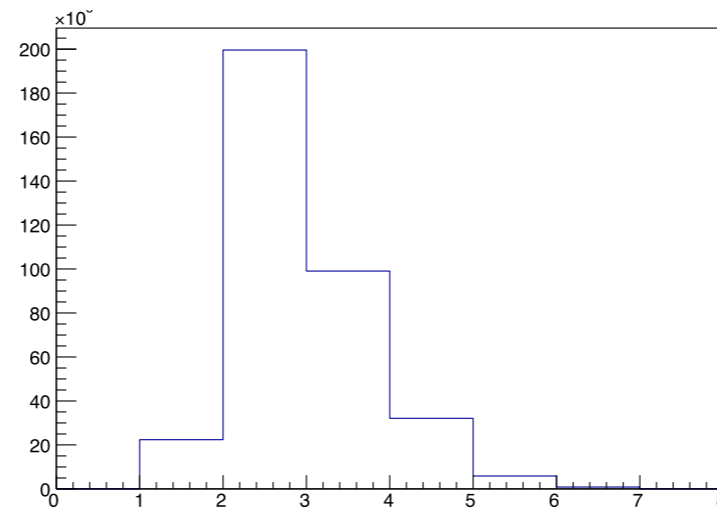
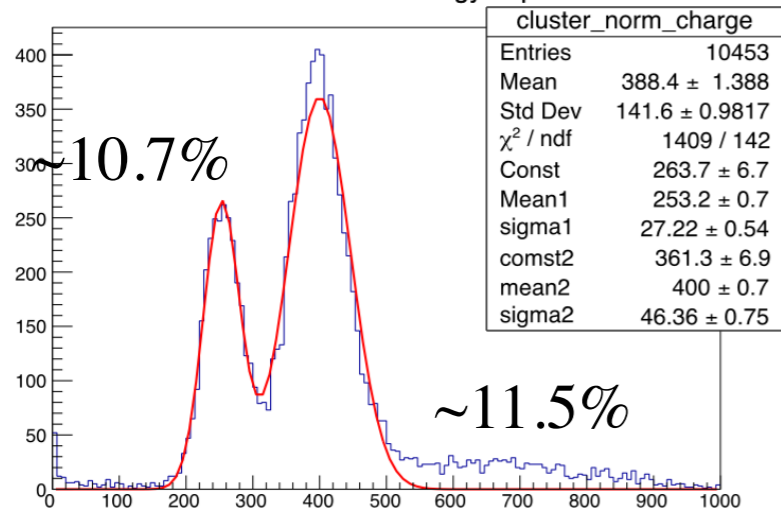
From S.Suvorov



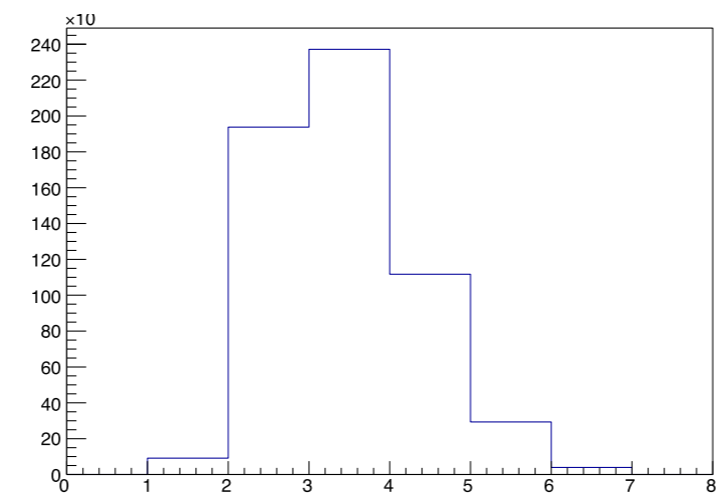
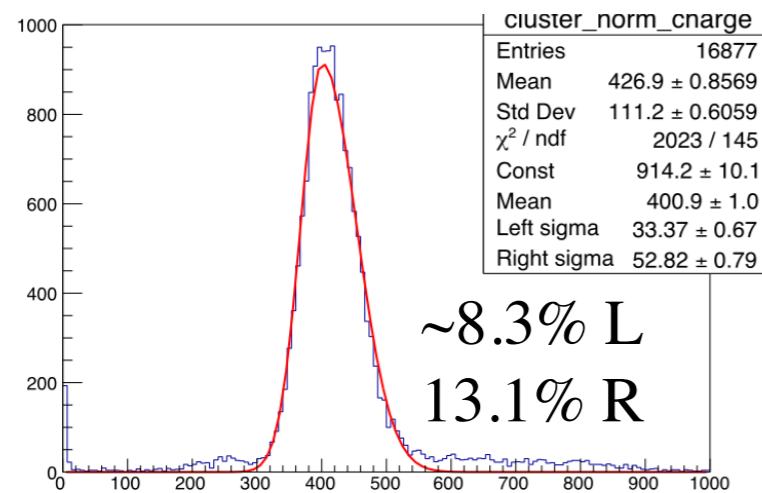
0.5 GeV/c muons



0.8 GeV/c proton

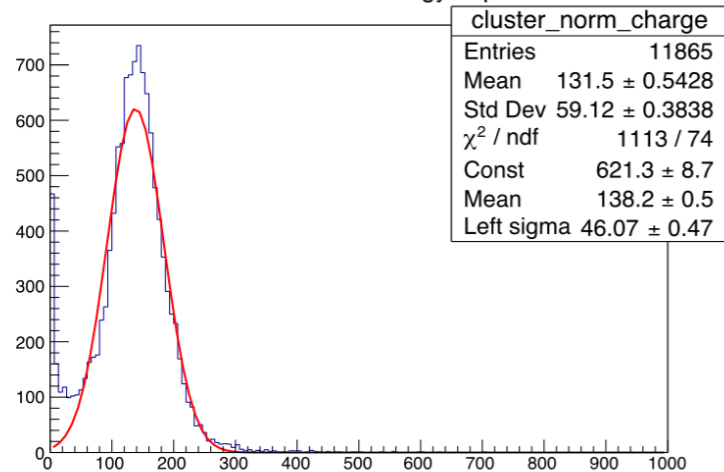


0.8 GeV/c pion trigger

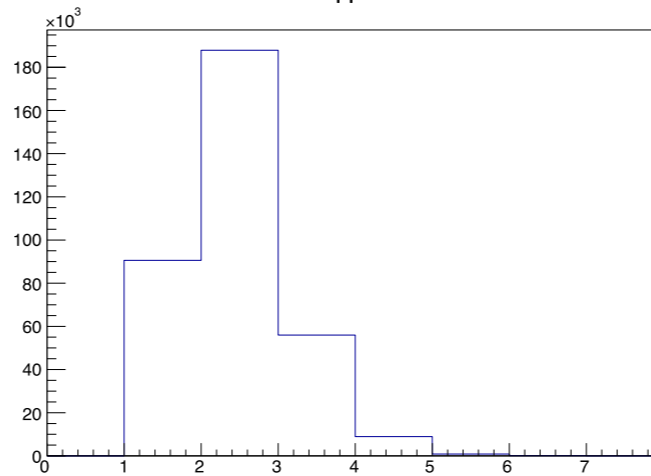


0.8 GeV/c electron trigger

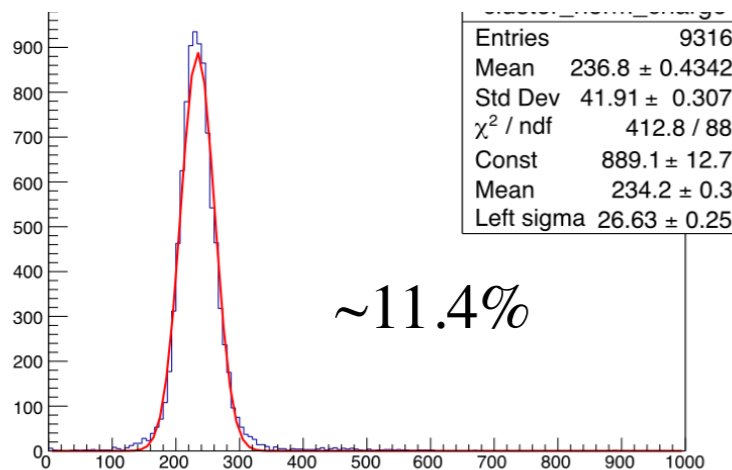
Truncated mean energy deposit



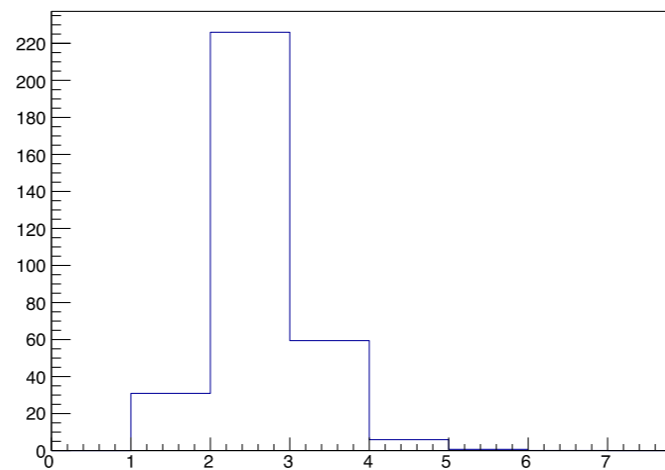
ppc



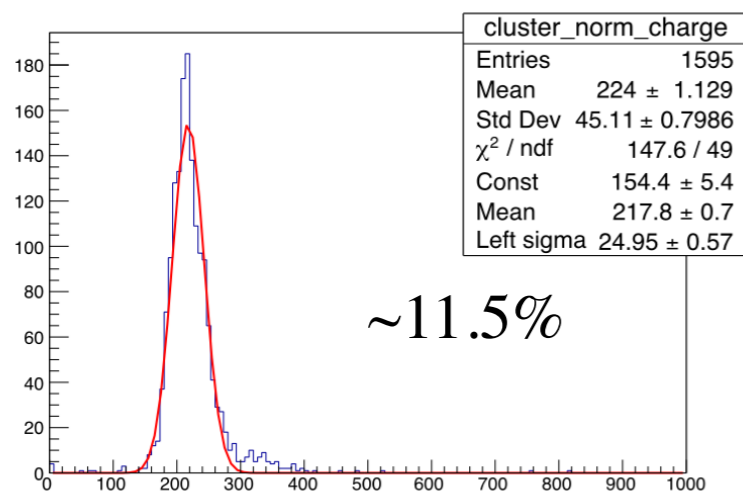
cosmic



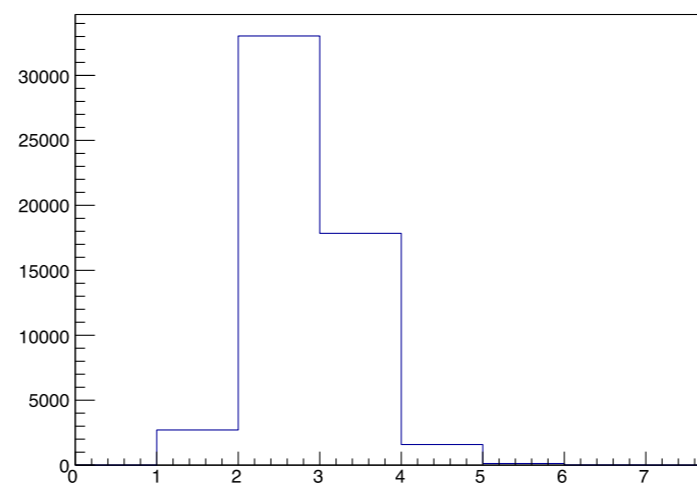
~11.4%



2 GeV/c muon



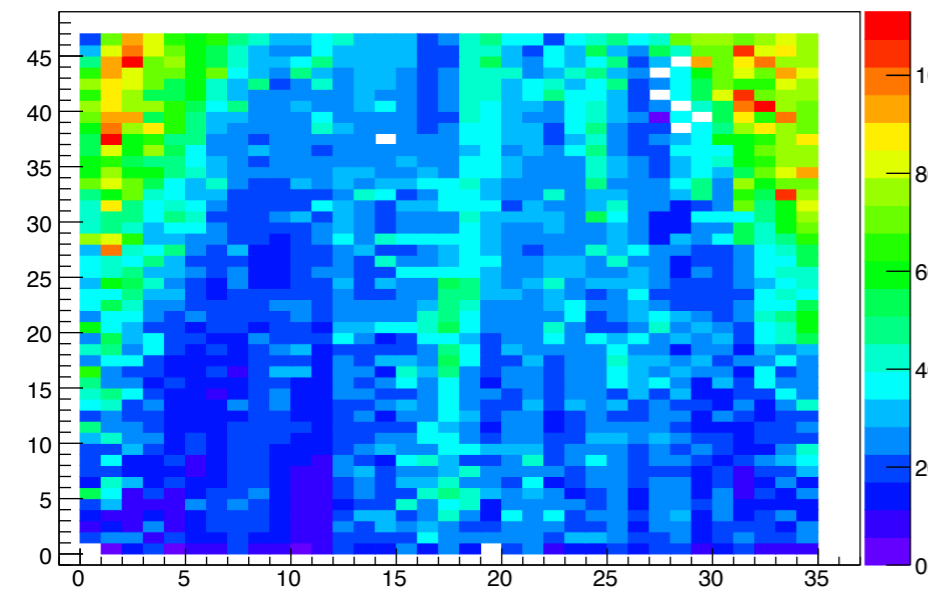
~11.5%



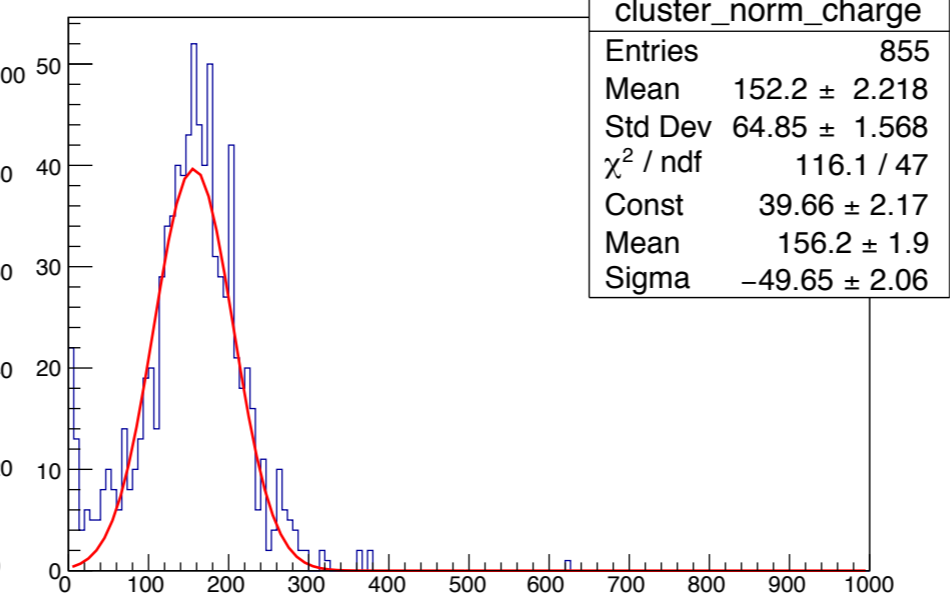
0.5 GeV/c muon

dE / dx resolution

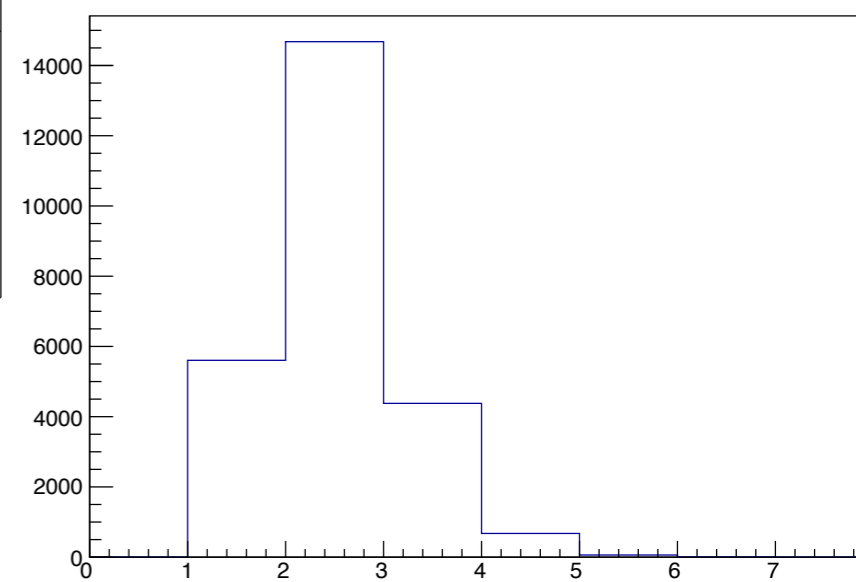
I vs J of hits



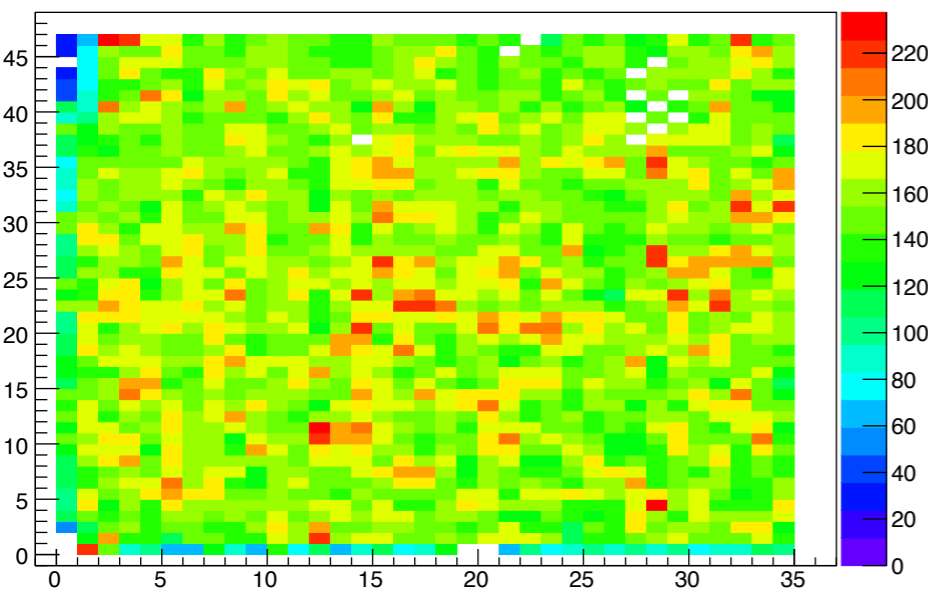
Truncated mean energy deposit



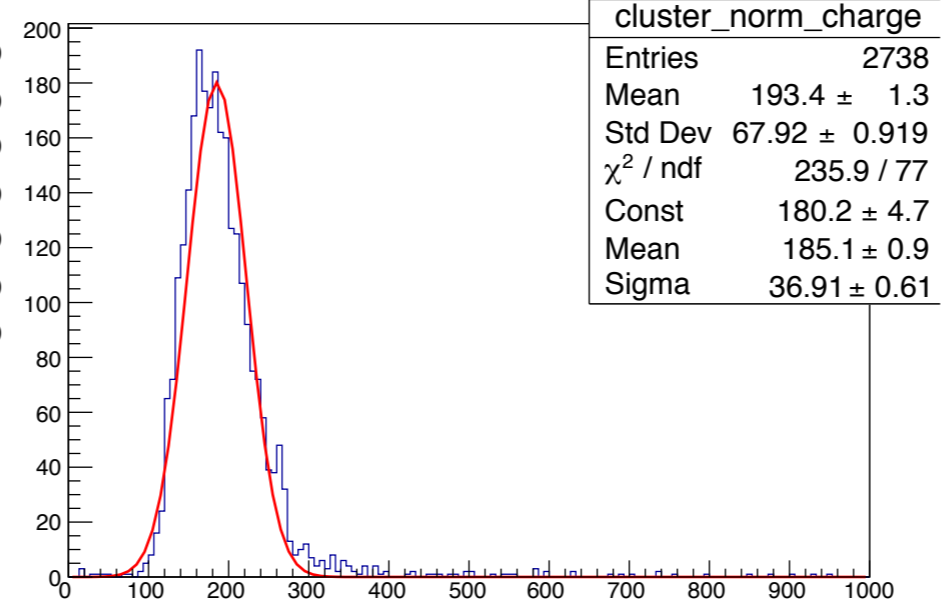
ppc



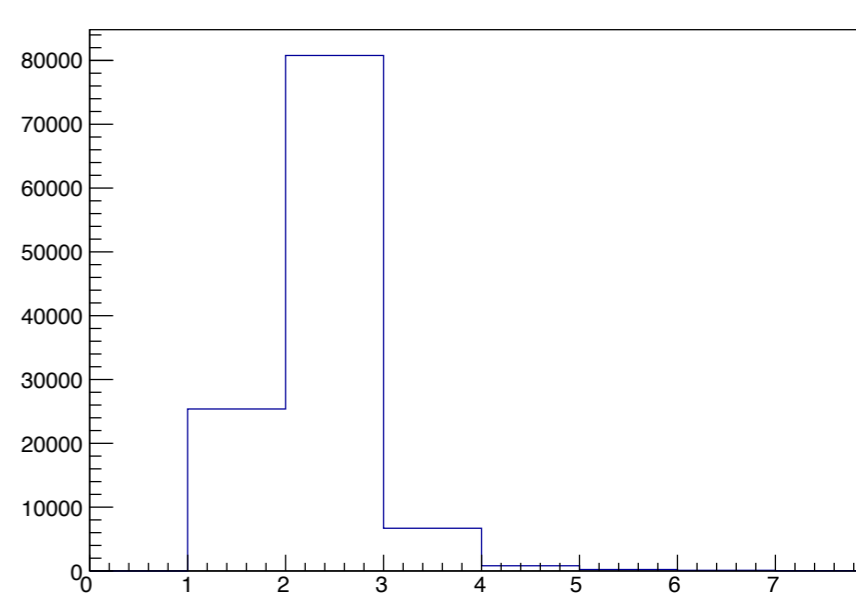
I vs J of hits



Truncated mean energy deposit



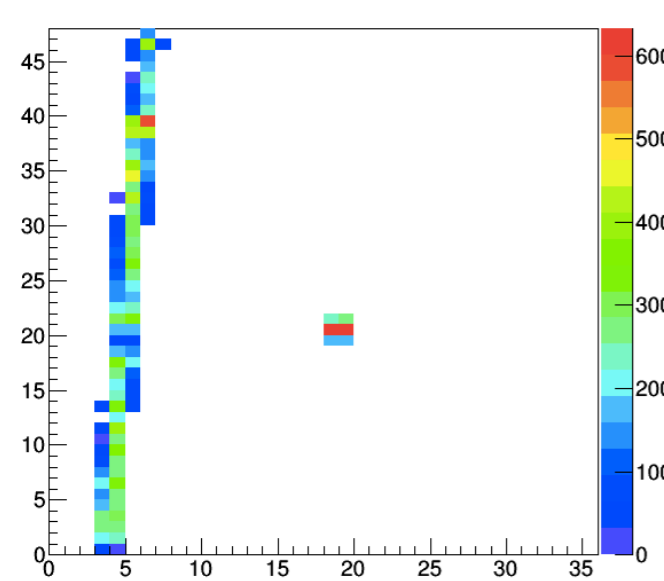
ppc



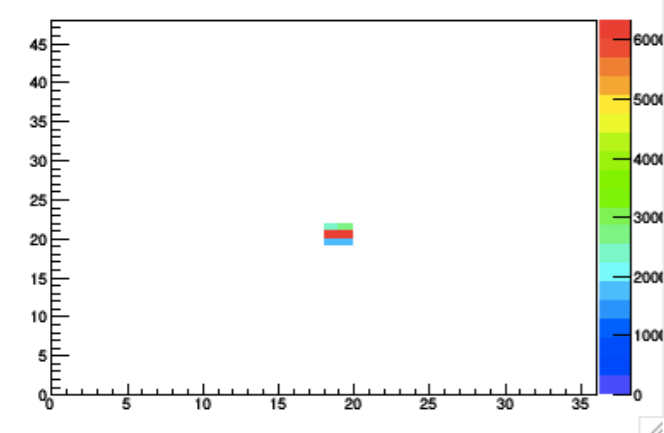
Iron 1 Cluster Selection

Duplicated Study

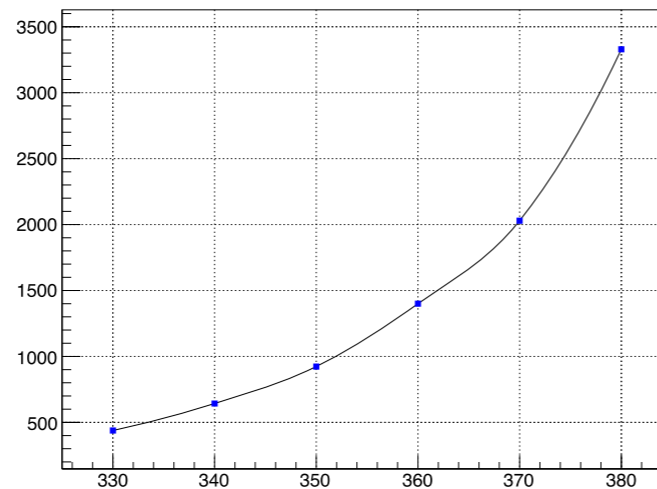
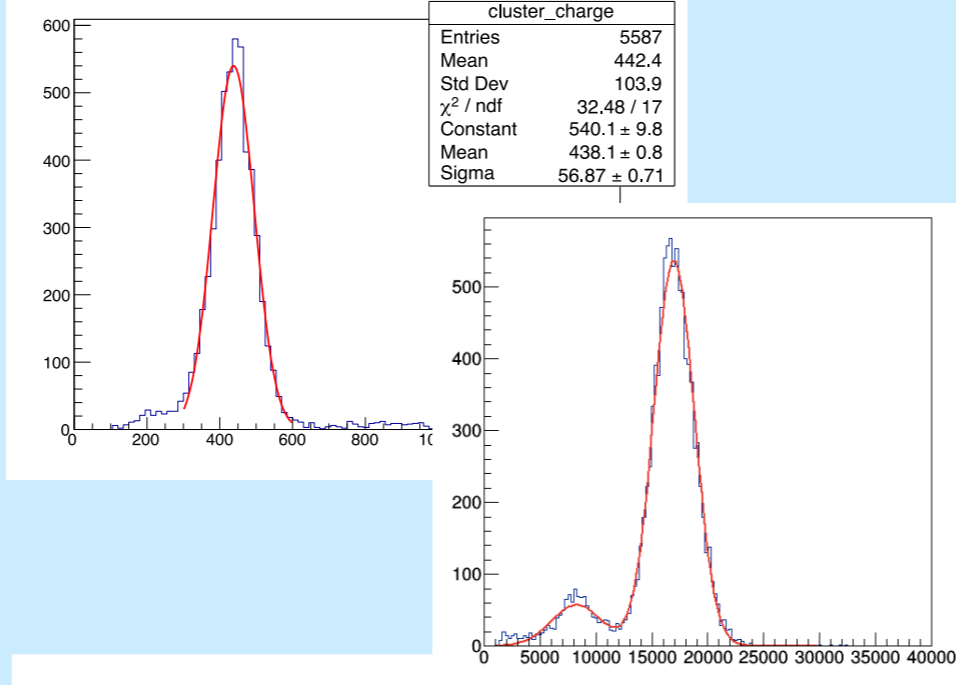
I vs J of hits



I vs J of hits



Cluster charge



Iron 2 Cluster Selection

A 2 cluster selection was also developed but it is only reliable for events with no tracks!

... very few statistics.

Looking into it it could be possible to compute experimentally Argon absorption length (from reabsorbed 2.9 keV photons)

It is tricky... abs length at that wavelength is know to be ~3cm, 2 iron clusters are expected to happen very close.